

$\tau$ 

$$J = \frac{1}{2}$$

$\tau$  discovery paper was PERL 75.  $e^+ e^- \rightarrow \tau^+ \tau^-$  cross-section threshold behavior and magnitude are consistent with pointlike spin-1/2 Dirac particle. BRANDELIK 78 ruled out pointlike spin-0 or spin-1 particle. FELDMAN 78 ruled out  $J = 3/2$ . KIRKBY 79 also ruled out  $J=\text{integer}$ ,  $J = 3/2$ .

## $\tau$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1776.86±0.12 OUR AVERAGE</b>				
1776.91±0.12 <sup>+0.10</sup> <sub>-0.13</sub>	1171	1 ABLIKIM	14D BES3	$23.3 \text{ pb}^{-1}, E_{\text{cm}}^{\text{ee}}=3.54\text{--}3.60 \text{ GeV}$
1776.68±0.12±0.41	682k	2 AUBERT	09AK BABR	$423 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$
1776.81 <sup>+0.25</sup> <sub>-0.23</sub> ±0.15	81	ANASHIN	07 KEDR	$6.7 \text{ pb}^{-1}, E_{\text{cm}}^{\text{ee}}=3.54\text{--}3.78 \text{ GeV}$
1776.61±0.13±0.35		2 BELOUS	07 BELL	$414 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$
1775.1 ±1.6 ±1.0	13.3k	3 ABBIENDI	00A OPAL	1990–1995 LEP runs
1778.2 ±0.8 ±1.2		ANASTASSOV	97 CLEO	$E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$
1776.96 <sup>+0.18</sup> <sub>-0.21</sub> <sup>+0.25</sup> <sub>-0.17</sub>	65	4 BAI	96 BES	$E_{\text{cm}}^{\text{ee}}=3.54\text{--}3.57 \text{ GeV}$
1776.3 ±2.4 ±1.4	11k	5 ALBRECHT	92M ARG	$E_{\text{cm}}^{\text{ee}}=9.4\text{--}10.6 \text{ GeV}$
1783 <sup>+3</sup> <sub>-4</sub>	692	6 BACINO	78B DLCO	$E_{\text{cm}}^{\text{ee}}=3.1\text{--}7.4 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1777.8 ±0.7 ±1.7	35k	7 BALEST	93 CLEO	Repl. by ANASTASSOV 97
1776.9 <sup>+0.4</sup> <sub>-0.5</sub> ±0.2	14	8 BAI	92 BES	Repl. by BAI 96

<sup>1</sup> ABLIKIM 14D fit  $\sigma(e^+ e^- \rightarrow \tau^+ \tau^-)$  at different energies near threshold.

<sup>2</sup> AUBERT 09AK and BELOUS 07 fit  $\tau$  pseudomass spectrum in  $\tau \rightarrow \pi\pi^+\pi^-\nu_\tau$  decays.  
Result assumes  $m_{\nu_\tau} = 0$ .

<sup>3</sup> ABBIENDI 00A fit  $\tau$  pseudomass spectrum in  $\tau \rightarrow \pi^\pm \leq 2\pi^0\nu_\tau$  and  
 $\tau \rightarrow \pi^\pm\pi^+\pi^- \leq 1\pi^0\nu_\tau$  decays. Result assumes  $m_{\nu_\tau}=0$ .

<sup>4</sup> BAI 96 fit  $\sigma(e^+ e^- \rightarrow \tau^+ \tau^-)$  at different energies near threshold.

<sup>5</sup> ALBRECHT 92M fit  $\tau$  pseudomass spectrum in  $\tau^- \rightarrow 2\pi^-\pi^+\nu_\tau$  decays. Result assumes  $m_{\nu_\tau}=0$ .

<sup>6</sup> BACINO 78B value comes from  $e^\pm X^\mp$  threshold. Published mass 1782 MeV increased by 1 MeV using the high precision  $\psi(2S)$  mass measurement of ZHOLENTZ 80 to eliminate the absolute SPEAR energy calibration uncertainty.

<sup>7</sup> BALEST 93 fit spectra of minimum kinematically allowed  $\tau$  mass in events of the type  $e^+ e^- \rightarrow \tau^+ \tau^- \rightarrow (\pi^+ n\pi^0\nu_\tau)(\pi^- m\pi^0\nu_\tau)$   $n \leq 2$ ,  $m \leq 2$ ,  $1 \leq n+m \leq 3$ . If  $m_{\nu_\tau} \neq 0$ , result increases by  $(m_{\nu_\tau}^2/1100 \text{ MeV})$ .

<sup>8</sup> BAI 92 fit  $\sigma(e^+ e^- \rightarrow \tau^+ \tau^-)$  near threshold using  $e\mu$  events.

$$(m_{\tau^+} - m_{\tau^-})/m_{\text{average}}$$

A test of *CPT* invariance.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.8 \times 10^{-4}$	90	BELOUS	07 BELL	$414 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<5.5 \times 10^{-4}$  90 <sup>1</sup> AUBERT 09AK BABR  $423 \text{ fb}^{-1}$ ,  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$<3.0 \times 10^{-3}$  90 ABBIENDI 00A OPAL 1990–1995 LEP runs

<sup>1</sup> AUBERT 09AK quote both the listed upper limit and  $(m_{\tau^+} - m_{\tau^-})/m_{\text{average}} = (-3.4 \pm 1.3 \pm 0.3) \times 10^{-4}$ .

## $\tau$ MEAN LIFE

VALUE ( $10^{-15} \text{ s}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>290.3 \pm 0.5</math> OUR AVERAGE</b>				
290.17 $\pm$ 0.53 $\pm$ 0.33	1.1M	BELOUS	14	BELL $711 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
290.9 $\pm$ 1.4 $\pm$ 1.0		ABDALLAH	04T	DLPH 1991–1995 LEP runs
293.2 $\pm$ 2.0 $\pm$ 1.5		ACCIARRI	00B	L3 1991–1995 LEP runs
290.1 $\pm$ 1.5 $\pm$ 1.1		BARATE	97R	ALEP 1989–1994 LEP runs
289.2 $\pm$ 1.7 $\pm$ 1.2		ALEXANDER	96E	OPAL 1990–1994 LEP runs
289.0 $\pm$ 2.8 $\pm$ 4.0	57.4k	BALEST	96	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
291.2 $\pm$ 2.0 $\pm$ 1.2		BARATE	97I	ALEP Repl. by BARATE 97R
291.4 $\pm$ 3.0		ABREU	96B	DLPH Repl. by ABDALLAH 04T
290.1 $\pm$ 4.0	34k	ACCIARRI	96K	L3 Repl. by ACCIARRI 00B
297 $\pm$ 9 $\pm$ 5	1671	ABE	95Y	SLD 1992–1993 SLC runs
304 $\pm$ 14 $\pm$ 7	4100	BATTLE	92	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
301 $\pm$ 29	3780	KLEINWORT	89	JADE $E_{\text{cm}}^{\text{ee}} = 35\text{--}46 \text{ GeV}$
288 $\pm$ 16 $\pm$ 17	807	AMIDEI	88	MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
306 $\pm$ 20 $\pm$ 14	695	BRAUNSCH...	88C	TASS $E_{\text{cm}}^{\text{ee}} = 36 \text{ GeV}$
299 $\pm$ 15 $\pm$ 10	1311	ABACHI	87C	HRS $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
295 $\pm$ 14 $\pm$ 11	5696	ALBRECHT	87P	ARG $E_{\text{cm}}^{\text{ee}} = 9.3\text{--}10.6 \text{ GeV}$
309 $\pm$ 17 $\pm$ 7	3788	BAND	87B	MAC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
325 $\pm$ 14 $\pm$ 18	8470	BEBEK	87C	CLEO $E_{\text{cm}}^{\text{ee}} = 10.5 \text{ GeV}$
460 $\pm$ 190	102	FELDMAN	82	MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

## $(\tau_{\tau^+} - \tau_{\tau^-}) / \tau_{\text{average}}$

Test of *CPT* invariance.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<7.0 \times 10^{-3}$	90	<sup>1</sup> BELOUS	14	BELL $711 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> BELOUS 14 quote limit on the absolute value of the relative lifetime difference.

## $\tau$ MAGNETIC MOMENT ANOMALY

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

$$\mu_{\tau}/(e\hbar/2m_{\tau}) - 1 = (g_{\tau} - 2)/2$$

For a theoretical calculation  $[(g_{\tau} - 2)/2 = 117\,721(5) \times 10^{-8}]$ , see EIDELMAN 07.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&gt; -0.052</math> and <math>&lt; 0.013</math> (CL = 95%) OUR LIMIT</b>				
$> -0.052$ and $< 0.013$	95	<sup>1</sup> ABDALLAH	04K	DLPH $e^+ e^- \rightarrow e^+ e^- \tau^+ \tau^-$ at LEP2

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.107	95	<sup>2</sup> ACHARD	04G L3	$e^+ e^- \rightarrow e^+ e^- \tau^+ \tau^-$ at LEP2
> -0.007 and < 0.005	95	<sup>3</sup> GONZALEZ-S..00	RVUE	$e^+ e^- \rightarrow \tau^+ \tau^-$ and $W \rightarrow \tau \nu_\tau$
> -0.052 and < 0.058	95	<sup>4</sup> ACCIARRI	98E L3	1991–1995 LEP runs
> -0.068 and < 0.065	95	<sup>5</sup> ACKERSTAFF	98N OPAL	1990–1995 LEP runs
> -0.004 and < 0.006	95	<sup>6</sup> ESCRIBANO	97 RVUE	$Z \rightarrow \tau^+ \tau^-$ at LEP
<0.01	95	<sup>7</sup> ESCRIBANO	93 RVUE	$Z \rightarrow \tau^+ \tau^-$ at LEP
<0.12	90	GRIFOLS	91 RVUE	$Z \rightarrow \tau \tau \gamma$ at LEP
<0.023	95	<sup>8</sup> SILVERMAN	83 RVUE	$e^+ e^- \rightarrow \tau^+ \tau^-$ at PETRA

<sup>1</sup> ABDALLAH 04K limit is derived from  $e^+ e^- \rightarrow e^+ e^- \tau^+ \tau^-$  total cross-section measurements at  $\sqrt{s}$  between 183 and 208 GeV. In addition to the limits, the authors also quote a value of  $-0.018 \pm 0.017$ .

<sup>2</sup> ACHARD 04G limit is derived from  $e^+ e^- \rightarrow e^+ e^- \tau^+ \tau^-$  total cross-section measurements at  $\sqrt{s}$  between 189 and 206 GeV, and is on the absolute value of the magnetic moment anomaly.

<sup>3</sup> GONZALEZ-SPRINBERG 00 use data on tau lepton production at LEP1, SLC, and LEP2, and data from colliders and LEP2 to determine limits. Assume imaginary component is zero.

<sup>4</sup> ACCIARRI 98E use  $Z \rightarrow \tau^+ \tau^- \gamma$  events. In addition to the limits, the authors also quote a value of  $0.004 \pm 0.027 \pm 0.023$ .

<sup>5</sup> ACKERSTAFF 98N use  $Z \rightarrow \tau^+ \tau^- \gamma$  events. The limit applies to an average of the form factor for off-shell  $\tau$ 's having  $p^2$  ranging from  $m_\tau^2$  to  $(M_Z - m_\tau)^2$ .

<sup>6</sup> ESCRIBANO 97 use preliminary experimental results.

<sup>7</sup> ESCRIBANO 93 limit derived from  $\Gamma(Z \rightarrow \tau^+ \tau^-)$ , and is on the absolute value of the magnetic moment anomaly.

<sup>8</sup> SILVERMAN 83 limit is derived from  $e^+ e^- \rightarrow \tau^+ \tau^-$  total cross-section measurements for  $q^2$  up to (37 GeV)<sup>2</sup>.

## $\tau$ ELECTRIC DIPOLE MOMENT ( $d_\tau$ )

A nonzero value is forbidden by both  $T$  invariance and  $P$  invariance.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

### $\text{Re}(d_\tau)$

VALUE ( $10^{-16}$ ecm)	CL%	DOCUMENT ID	TECN	COMMENT
= <b>0.22 to 0.45</b>	95	<sup>1</sup> INAMI	03 BELL	$E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 2.3	90	<sup>2</sup> GROZIN	09A RVUE	From e EDM limit
< 3.7	95	<sup>3</sup> ABDALLAH	04K DLPH	$e^+ e^- \rightarrow e^+ e^- \tau^+ \tau^-$ at LEP2
< 11.4	95	<sup>4</sup> ACHARD	04G L3	$e^+ e^- \rightarrow e^+ e^- \tau^+ \tau^-$ at LEP2
< 4.6	95	<sup>5</sup> ALBRECHT	00 ARG	$E_{\text{cm}}^{ee} = 10.4$ GeV
> -3.1 and < 3.1	95	ACCIARRI	98E L3	1991–1995 LEP runs
> -3.8 and < 3.6	95	<sup>6</sup> ACKERSTAFF	98N OPAL	1990–1995 LEP runs
< 0.11	95	<sup>7,8</sup> ESCRIBANO	97 RVUE	$Z \rightarrow \tau^+ \tau^-$ at LEP
< 0.5	95	<sup>9</sup> ESCRIBANO	93 RVUE	$Z \rightarrow \tau^+ \tau^-$ at LEP
< 7	90	GRIFOLS	91 RVUE	$Z \rightarrow \tau \tau \gamma$ at LEP
< 1.6	90	DELAGUILA	90 RVUE	$e^+ e^- \rightarrow \tau^+ \tau^-$ $E_{\text{cm}}^{ee} = 35$ GeV

<sup>1</sup> INAMI 03 use  $e^+ e^- \rightarrow \tau^+ \tau^-$  events.

<sup>2</sup> GROZIN 09A calculate the contribution to the electron electric dipole moment from the  $\tau$  electric dipole moment appearing in loops, which is  $\Delta d_e = 6.9 \times 10^{-12} d_\tau$ . Dividing the REGAN 02 upper limit  $|d_e| \leq 1.6 \times 10^{-27}$  e cm at CL=90% by  $6.9 \times 10^{-12}$  gives this limit.

<sup>3</sup> ABDALLAH 04K limit is derived from  $e^+ e^- \rightarrow e^+ e^- \tau^+ \tau^-$  total cross-section measurements at  $\sqrt{s}$  between 183 and 208 GeV and is on the absolute value of  $d_\tau$ .

<sup>4</sup> ACHARD 04G limit is derived from  $e^+ e^- \rightarrow e^+ e^- \tau^+ \tau^-$  total cross-section measurements at  $\sqrt{s}$  between 189 and 206 GeV, and is on the absolute value of  $d_\tau$ .

<sup>5</sup> ALBRECHT 00 use  $e^+ e^- \rightarrow \tau^+ \tau^-$  events. Limit is on the absolute value of  $\text{Re}(d_\tau)$ .

<sup>6</sup> ACKERSTAFF 98N use  $Z \rightarrow \tau^+ \tau^- \gamma$  events. The limit applies to an average of the form factor for off-shell  $\tau$ 's having  $p^2$  ranging from  $m_\tau^2$  to  $(M_Z - m_\tau)^2$ .

<sup>7</sup> ESCRIBANO 97 derive the relationship  $|d_\tau| = \cot \theta_W |d_\tau^W|$  using effective Lagrangian methods, and use a conference result  $|d_\tau^W| < 5.8 \times 10^{-18}$  e cm at 95% CL (L. Silvestris, ICHEP96) to obtain this result.

<sup>8</sup> ESCRIBANO 97 use preliminary experimental results.

<sup>9</sup> ESCRIBANO 93 limit derived from  $\Gamma(Z \rightarrow \tau^+ \tau^-)$ , and is on the absolute value of the electric dipole moment.

## Im( $d_\tau$ )

VALUE ( $10^{-16}$ e cm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>-0.25 to 0.008</b>	95	<sup>1</sup> INAMI 03	BELL	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
< 1.8	95	<sup>2</sup> ALBRECHT 00	ARG	$E_{\text{cm}}^{\text{ee}} = 10.4$ GeV

<sup>1</sup> INAMI 03 use  $e^+ e^- \rightarrow \tau^+ \tau^-$  events.

<sup>2</sup> ALBRECHT 00 use  $e^+ e^- \rightarrow \tau^+ \tau^-$  events. Limit is on the absolute value of  $\text{Im}(d_\tau)$ .

## $\tau$ WEAK DIPOLE MOMENT ( $d_\tau^W$ )

A nonzero value is forbidden by  $CP$  invariance.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

## Re( $d_\tau^W$ )

VALUE ( $10^{-17}$ e cm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.50</b>	95	<sup>1</sup> HEISTER 03F	ALEP	1990–1995 LEP runs
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
<3.0	90	<sup>1</sup> ACCIARRI 98C	L3	1991–1995 LEP runs
<0.56	95	ACKERSTAFF 97L	OPAL	1991–1995 LEP runs
<0.78	95	<sup>2</sup> AKERS 95F	OPAL	Repl. by ACKERSTAFF 97L
<1.5	95	<sup>2</sup> BUSKULIC 95C	ALEP	Repl. by HEISTER 03F
<7.0	95	<sup>2</sup> ACTON 92F	OPAL	$Z \rightarrow \tau^+ \tau^-$ at LEP
<3.7	95	<sup>2</sup> BUSKULIC 92J	ALEP	Repl. by BUSKULIC 95C

<sup>1</sup> Limit is on the absolute value of the real part of the weak dipole moment.

<sup>2</sup> Limit is on the absolute value of the real part of the weak dipole moment, and applies for  $q^2 = m_Z^2$ .

**Im( $d_\tau^w$ )**

<u>VALUE</u> ( $10^{-17}$ ecm)	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<1.1	95	1 HEISTER	03F ALEP	1990–1995 LEP runs
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
<1.5	95	ACKERSTAFF 97L	OPAL	1991–1995 LEP runs
<4.5	95	2 AKERS	95F OPAL	Repl. by ACKERSTAFF 97L

<sup>1</sup> HEISTER 03F limit is on the absolute value of the imaginary part of the weak dipole moment.  
<sup>2</sup> Limit is on the absolute value of the imaginary part of the weak dipole moment, and applies for  $q^2 = m_Z^2$ .

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 **$\tau^-$  WEAK ANOMALOUS MAGNETIC DIPOLE MOMENT ( $\alpha_\tau^w$ )**

Electroweak radiative corrections are expected to contribute at the  $10^{-6}$  level. See BERNABEU 95.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

**Re( $\alpha_\tau^w$ )**

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<1.1 × 10 <sup>-3</sup>	95	1 HEISTER	03F ALEP	1990–1995 LEP runs
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
> -0.0024 and < 0.0025	95	2 GONZALEZ-S..00	RVUE	$e^+ e^- \rightarrow \tau^+ \tau^-$ and $W \rightarrow \tau \nu_\tau$
<4.5 × 10 <sup>-3</sup>	90	1 ACCIARRI	98C L3	1991–1995 LEP runs

<sup>1</sup> Limit is on the absolute value of the real part of the weak anomalous magnetic dipole moment.

<sup>2</sup> GONZALEZ-SPRINBERG 00 use data on tau lepton production at LEP1, SLC, and LEP2, and data from colliders and LEP2 to determine limits. Assume imaginary component is zero.

**Im( $\alpha_\tau^w$ )**

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<2.7 × 10 <sup>-3</sup>	95	1 HEISTER	03F ALEP	1990–1995 LEP runs
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
<9.9 × 10 <sup>-3</sup>	90	1 ACCIARRI	98C L3	1991–1995 LEP runs

<sup>1</sup> Limit is on the absolute value of the imaginary part of the weak anomalous magnetic dipole moment.

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 **$\tau^-$  DECAY MODES**

$\tau^+$  modes are charge conjugates of the modes below. “ $h^\pm$ ” stands for  $\pi^\pm$  or  $K^\pm$ . “ $\ell$ ” stands for  $e$  or  $\mu$ . “Neutrals” stands for  $\gamma$ 's and/or  $\pi^0$ 's.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
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**Modes with one charged particle**

$\Gamma_1$	particle $^- \geq 0$ neutrals $\geq 0 K^0 \nu_\tau$	(85.24 $\pm$ 0.06) %
	(“1-prong”)	
$\Gamma_2$	particle $^- \geq 0$ neutrals $\geq 0 K_L^0 \nu_\tau$	(84.58 $\pm$ 0.06) %
$\Gamma_3$	$\mu^- \bar{\nu}_\mu \nu_\tau$	[a] (17.39 $\pm$ 0.04) %
$\Gamma_4$	$\mu^- \bar{\nu}_\mu \nu_\tau \gamma$	[b] ( $3.67 \pm 0.08$ ) $\times 10^{-3}$
$\Gamma_5$	$e^- \bar{\nu}_e \nu_\tau$	[a] (17.82 $\pm$ 0.04) %
$\Gamma_6$	$e^- \bar{\nu}_e \nu_\tau \gamma$	[b] ( $1.83 \pm 0.05$ ) %
$\Gamma_7$	$h^- \geq 0 K_L^0 \nu_\tau$	(12.03 $\pm$ 0.05) %
$\Gamma_8$	$h^- \nu_\tau$	(11.51 $\pm$ 0.05) %
$\Gamma_9$	$\pi^- \nu_\tau$	[a] (10.82 $\pm$ 0.05) %
$\Gamma_{10}$	$K^- \nu_\tau$	[a] ( $6.96 \pm 0.10$ ) $\times 10^{-3}$
$\Gamma_{11}$	$h^- \geq 1$ neutrals $\nu_\tau$	(37.01 $\pm$ 0.09) %
$\Gamma_{12}$	$h^- \geq 1 \pi^0 \nu_\tau$ (ex. $K^0$ )	(36.51 $\pm$ 0.09) %
$\Gamma_{13}$	$h^- \pi^0 \nu_\tau$	(25.93 $\pm$ 0.09) %
$\Gamma_{14}$	$\pi^- \pi^0 \nu_\tau$	[a] (25.49 $\pm$ 0.09) %
$\Gamma_{15}$	$\pi^- \pi^0$ non- $\rho(770)$ $\nu_\tau$	( $3.0 \pm 3.2$ ) $\times 10^{-3}$
$\Gamma_{16}$	$K^- \pi^0 \nu_\tau$	[a] ( $4.33 \pm 0.15$ ) $\times 10^{-3}$
$\Gamma_{17}$	$h^- \geq 2 \pi^0 \nu_\tau$	(10.81 $\pm$ 0.09) %
$\Gamma_{18}$	$h^- 2 \pi^0 \nu_\tau$	(9.48 $\pm$ 0.10) %
$\Gamma_{19}$	$h^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ )	(9.32 $\pm$ 0.10) %
$\Gamma_{20}$	$\pi^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ )	[a] (9.26 $\pm$ 0.10) %
$\Gamma_{21}$	$\pi^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ ), scalar	< 9 $\times 10^{-3}$ CL=95%
$\Gamma_{22}$	$\pi^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ ), vector	< 7 $\times 10^{-3}$ CL=95%
$\Gamma_{23}$	$K^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ )	[a] ( $6.5 \pm 2.2$ ) $\times 10^{-4}$
$\Gamma_{24}$	$h^- \geq 3 \pi^0 \nu_\tau$	(1.34 $\pm$ 0.07) %
$\Gamma_{25}$	$h^- \geq 3 \pi^0 \nu_\tau$ (ex. $K^0$ )	(1.25 $\pm$ 0.07) %
$\Gamma_{26}$	$h^- 3 \pi^0 \nu_\tau$	(1.18 $\pm$ 0.07) %
$\Gamma_{27}$	$\pi^- 3 \pi^0 \nu_\tau$ (ex. $K^0$ )	[a] (1.04 $\pm$ 0.07) %
$\Gamma_{28}$	$K^- 3 \pi^0 \nu_\tau$ (ex. $K^0$ , $\eta$ )	[a] (4.8 $\pm$ 2.1) $\times 10^{-4}$
$\Gamma_{29}$	$h^- 4 \pi^0 \nu_\tau$ (ex. $K^0$ )	(1.6 $\pm$ 0.4) $\times 10^{-3}$
$\Gamma_{30}$	$h^- 4 \pi^0 \nu_\tau$ (ex. $K^0$ , $\eta$ )	[a] (1.1 $\pm$ 0.4) $\times 10^{-3}$
$\Gamma_{31}$	$a_1(1260) \nu_\tau \rightarrow \pi^- \gamma \nu_\tau$	(3.8 $\pm$ 1.5) $\times 10^{-4}$
$\Gamma_{32}$	$K^- \geq 0 \pi^0 \geq 0 K^0 \geq 0 \gamma \nu_\tau$	( $1.552 \pm 0.029$ ) %
$\Gamma_{33}$	$K^- \geq 1 (\pi^0 \text{ or } K^0 \text{ or } \gamma) \nu_\tau$	(8.59 $\pm$ 0.28) $\times 10^{-3}$

**Modes with  $K^0$ 's**

$\Gamma_{34}$	$K_S^0$ (particles) $^- \nu_\tau$	(9.43 $\pm$ 0.28) $\times 10^{-3}$
$\Gamma_{35}$	$h^- \bar{K}^0 \nu_\tau$	(9.87 $\pm$ 0.14) $\times 10^{-3}$
$\Gamma_{36}$	$\pi^- \bar{K}^0 \nu_\tau$	[a] (8.38 $\pm$ 0.14) $\times 10^{-3}$
$\Gamma_{37}$	$\pi^- \bar{K}^0$ (non- $K^*(892)^-$ ) $\nu_\tau$	(5.4 $\pm$ 2.1) $\times 10^{-4}$
$\Gamma_{38}$	$K^- K^0 \nu_\tau$	[a] (1.486 $\pm$ 0.034) $\times 10^{-3}$
$\Gamma_{39}$	$K^- K^0 \geq 0 \pi^0 \nu_\tau$	(2.99 $\pm$ 0.07) $\times 10^{-3}$

$\Gamma_{40}$	$h^- \bar{K}^0 \pi^0 \nu_\tau$	$( 5.32 \pm 0.13 ) \times 10^{-3}$
$\Gamma_{41}$	$\pi^- \bar{K}^0 \pi^0 \nu_\tau$	$[a] ( 3.82 \pm 0.13 ) \times 10^{-3}$
$\Gamma_{42}$	$\bar{K}^0 \rho^- \nu_\tau$	$( 2.2 \pm 0.5 ) \times 10^{-3}$
$\Gamma_{43}$	$K^- K^0 \pi^0 \nu_\tau$	$[a] ( 1.50 \pm 0.07 ) \times 10^{-3}$
$\Gamma_{44}$	$\pi^- \bar{K}^0 \geq 1 \pi^0 \nu_\tau$	$( 4.08 \pm 0.25 ) \times 10^{-3}$
$\Gamma_{45}$	$\pi^- \bar{K}^0 \pi^0 \pi^0 \nu_\tau (\text{ex. } K^0)$	$[a] ( 2.6 \pm 2.3 ) \times 10^{-4}$
$\Gamma_{46}$	$K^- K^0 \pi^0 \pi^0 \nu_\tau$	$< 1.6 \times 10^{-4} \text{ CL}=95\%$
$\Gamma_{47}$	$\pi^- K^0 \bar{K}^0 \nu_\tau$	$( 1.55 \pm 0.24 ) \times 10^{-3}$
$\Gamma_{48}$	$\pi^- K_S^0 K_S^0 \nu_\tau$	$[a] ( 2.35 \pm 0.06 ) \times 10^{-4}$
$\Gamma_{49}$	$\pi^- K_S^0 K_L^0 \nu_\tau$	$[a] ( 1.08 \pm 0.24 ) \times 10^{-3}$
$\Gamma_{50}$	$\pi^- K_L^0 K_L^0 \nu_\tau$	$( 2.35 \pm 0.06 ) \times 10^{-4}$
$\Gamma_{51}$	$\pi^- K^0 \bar{K}^0 \pi^0 \nu_\tau$	$( 3.6 \pm 1.2 ) \times 10^{-4}$
$\Gamma_{52}$	$\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$	$[a] ( 1.82 \pm 0.21 ) \times 10^{-5}$
$\Gamma_{53}$	$K^{*-} K^0 \pi^0 \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$	$( 1.08 \pm 0.21 ) \times 10^{-5}$
$\Gamma_{54}$	$f_1(1285) \pi^- \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$	$( 6.8 \pm 1.5 ) \times 10^{-6}$
$\Gamma_{55}$	$f_1(1420) \pi^- \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$	$( 2.4 \pm 0.8 ) \times 10^{-6}$
$\Gamma_{56}$	$\pi^- K_S^0 K_L^0 \pi^0 \nu_\tau$	$[a] ( 3.2 \pm 1.2 ) \times 10^{-4}$
$\Gamma_{57}$	$\pi^- K_L^0 K_L^0 \pi^0 \nu_\tau$	$( 1.82 \pm 0.21 ) \times 10^{-5}$
$\Gamma_{58}$	$K^- K_S^0 K_S^0 \nu_\tau$	$< 6.3 \times 10^{-7} \text{ CL}=90\%$
$\Gamma_{59}$	$K^- K_S^0 K_S^0 \pi^0 \nu_\tau$	$< 4.0 \times 10^{-7} \text{ CL}=90\%$
$\Gamma_{60}$	$K^0 h^+ h^- h^- \geq 0 \text{ neutrals } \nu_\tau$	$< 1.7 \times 10^{-3} \text{ CL}=95\%$
$\Gamma_{61}$	$K^0 h^+ h^- h^- \nu_\tau$	$[a] ( 2.5 \pm 2.0 ) \times 10^{-4}$

**Modes with three charged particles**

$\Gamma_{62}$	$h^- h^- h^+ \geq 0 \text{ neutrals } \geq 0 K_L^0 \nu_\tau$	$( 15.20 \pm 0.06 ) \%$
$\Gamma_{63}$	$h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau$ (ex. $K_S^0 \rightarrow \pi^+ \pi^-$ ) ("3-prong")	$( 14.55 \pm 0.06 ) \%$
$\Gamma_{64}$	$h^- h^- h^+ \nu_\tau$	$( 9.80 \pm 0.05 ) \%$
$\Gamma_{65}$	$h^- h^- h^+ \nu_\tau (\text{ex. } K^0)$	$( 9.46 \pm 0.05 ) \%$
$\Gamma_{66}$	$h^- h^- h^+ \nu_\tau (\text{ex. } K^0, \omega)$	$( 9.43 \pm 0.05 ) \%$
$\Gamma_{67}$	$\pi^- \pi^+ \pi^- \nu_\tau$	$( 9.31 \pm 0.05 ) \%$
$\Gamma_{68}$	$\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)$	$( 9.02 \pm 0.05 ) \%$
$\Gamma_{69}$	$\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0),$ non-axial vector	$< 2.4 \% \text{ CL}=95\%$
$\Gamma_{70}$	$\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0, \omega)$	$[a] ( 8.99 \pm 0.05 ) \%$
$\Gamma_{71}$	$h^- h^- h^+ \geq 1 \text{ neutrals } \nu_\tau$	$( 5.29 \pm 0.05 ) \%$
$\Gamma_{72}$	$h^- h^- h^+ \geq 1 \pi^0 \nu_\tau (\text{ex. } K^0)$	$( 5.09 \pm 0.05 ) \%$
$\Gamma_{73}$	$h^- h^- h^+ \pi^0 \nu_\tau$	$( 4.76 \pm 0.05 ) \%$
$\Gamma_{74}$	$h^- h^- h^+ \pi^0 \nu_\tau (\text{ex. } K^0)$	$( 4.57 \pm 0.05 ) \%$
$\Gamma_{75}$	$h^- h^- h^+ \pi^0 \nu_\tau (\text{ex. } K^0, \omega)$	$( 2.79 \pm 0.07 ) \%$

$\Gamma_{76}$	$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$	$( 4.62 \pm 0.05 ) \%$
$\Gamma_{77}$	$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)$	$( 4.49 \pm 0.05 ) \%$
$\Gamma_{78}$	$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0, \omega)$	[a] $( 2.74 \pm 0.07 ) \%$
$\Gamma_{79}$	$h^- \rho \pi^0 \nu_\tau$	
$\Gamma_{80}$	$h^- \rho^+ h^- \nu_\tau$	
$\Gamma_{81}$	$h^- \rho^- h^+ \nu_\tau$	
$\Gamma_{82}$	$h^- h^- h^+ \geq 2\pi^0 \nu_\tau (\text{ex. } K^0)$	$( 5.17 \pm 0.31 ) \times 10^{-3}$
$\Gamma_{83}$	$h^- h^- h^+ 2\pi^0 \nu_\tau$	$( 5.05 \pm 0.31 ) \times 10^{-3}$
$\Gamma_{84}$	$h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0)$	$( 4.95 \pm 0.31 ) \times 10^{-3}$
$\Gamma_{85}$	$h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0, \omega, \eta)$	[a] $( 10 \pm 4 ) \times 10^{-4}$
$\Gamma_{86}$	$h^- h^- h^+ 3\pi^0 \nu_\tau$	$( 2.13 \pm 0.30 ) \times 10^{-4}$
$\Gamma_{87}$	$2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0)$	$( 1.95 \pm 0.30 ) \times 10^{-4}$
$\Gamma_{88}$	$2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta, f_1(1285))$	$( 1.7 \pm 0.4 ) \times 10^{-4}$
$\Gamma_{89}$	$2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285))$	[a] $( 1.4 \pm 2.7 ) \times 10^{-5}$
$\Gamma_{90}$	$K^- h^+ h^- \geq 0 \text{ neutrals } \nu_\tau$	$( 6.29 \pm 0.14 ) \times 10^{-3}$
$\Gamma_{91}$	$K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0)$	$( 4.37 \pm 0.07 ) \times 10^{-3}$
$\Gamma_{92}$	$K^- h^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)$	$( 8.6 \pm 1.2 ) \times 10^{-4}$
$\Gamma_{93}$	$K^- \pi^+ \pi^- \geq 0 \text{ neutrals } \nu_\tau$	$( 4.77 \pm 0.14 ) \times 10^{-3}$
$\Gamma_{94}$	$K^- \pi^+ \pi^- \geq 0 \pi^0 \nu_\tau (\text{ex. } K^0)$	$( 3.73 \pm 0.13 ) \times 10^{-3}$
$\Gamma_{95}$	$K^- \pi^+ \pi^- \nu_\tau$	$( 3.45 \pm 0.07 ) \times 10^{-3}$
$\Gamma_{96}$	$K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)$	$( 2.93 \pm 0.07 ) \times 10^{-3}$
$\Gamma_{97}$	$K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0, \omega)$	[a] $( 2.93 \pm 0.07 ) \times 10^{-3}$
$\Gamma_{98}$	$K^- \rho^0 \nu_\tau \rightarrow K^- \pi^+ \pi^- \nu_\tau$	$( 1.4 \pm 0.5 ) \times 10^{-3}$
$\Gamma_{99}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$	$( 1.31 \pm 0.12 ) \times 10^{-3}$
$\Gamma_{100}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)$	$( 7.9 \pm 1.2 ) \times 10^{-4}$
$\Gamma_{101}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0, \eta)$	$( 7.6 \pm 1.2 ) \times 10^{-4}$
$\Gamma_{102}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0, \omega)$	$( 3.7 \pm 0.9 ) \times 10^{-4}$
$\Gamma_{103}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0, \omega, \eta)$	[a] $( 3.9 \pm 1.4 ) \times 10^{-4}$
$\Gamma_{104}$	$K^- \pi^+ K^- \geq 0 \text{ neut. } \nu_\tau$	$< 9 \times 10^{-4} \text{ CL=95\%}$
$\Gamma_{105}$	$K^- K^+ \pi^- \geq 0 \text{ neut. } \nu_\tau$	$( 1.496 \pm 0.033 ) \times 10^{-3}$
$\Gamma_{106}$	$K^- K^+ \pi^- \nu_\tau$	[a] $( 1.435 \pm 0.027 ) \times 10^{-3}$
$\Gamma_{107}$	$K^- K^+ \pi^- \pi^0 \nu_\tau$	[a] $( 6.1 \pm 1.8 ) \times 10^{-5}$
$\Gamma_{108}$	$K^- K^+ K^- \nu_\tau$	$( 2.2 \pm 0.8 ) \times 10^{-5} \text{ S=5.4}$
$\Gamma_{109}$	$K^- K^+ K^- \nu_\tau (\text{ex. } \phi)$	$< 2.5 \times 10^{-6} \text{ CL=90\%}$
$\Gamma_{110}$	$K^- K^+ K^- \pi^0 \nu_\tau$	$< 4.8 \times 10^{-6} \text{ CL=90\%}$
$\Gamma_{111}$	$\pi^- K^+ \pi^- \geq 0 \text{ neut. } \nu_\tau$	$< 2.5 \times 10^{-3} \text{ CL=95\%}$
$\Gamma_{112}$	$e^- e^- e^+ \bar{\nu}_e \nu_\tau$	$( 2.8 \pm 1.5 ) \times 10^{-5}$
$\Gamma_{113}$	$\mu^- e^- e^+ \bar{\nu}_\mu \nu_\tau$	$< 3.2 \times 10^{-5} \text{ CL=90\%}$
$\Gamma_{114}$	$\pi^- e^- e^+ \nu_\tau$	seen
$\Gamma_{115}$	$\pi^- \mu^- \mu^+ \nu_\tau$	$< 1.14 \times 10^{-5} \text{ CL=90\%}$

**Modes with five charged particles**

$\Gamma_{116}$	$3h^- 2h^+ \geq 0$ neutrals $\nu_\tau$ (ex. $K_S^0 \rightarrow \pi^- \pi^+$ ) ("5-prong")	$(9.9 \pm 0.4) \times 10^{-4}$
$\Gamma_{117}$	$3h^- 2h^+ \nu_\tau$ (ex. $K^0$ )	$(8.29 \pm 0.31) \times 10^{-4}$
$\Gamma_{118}$	$3\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0, \omega$ )	$(8.27 \pm 0.31) \times 10^{-4}$
$\Gamma_{119}$	$3\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0, \omega, f_1(1285)$ )	[a] $(7.75 \pm 0.30) \times 10^{-4}$
$\Gamma_{120}$	$K^- 2\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0$ )	[a] $(6 \pm 12) \times 10^{-7}$
$\Gamma_{121}$	$K^+ 3\pi^- \pi^+ \nu_\tau$	$< 5.0 \times 10^{-6}$ CL=90%
$\Gamma_{122}$	$K^+ K^- 2\pi^- \pi^+ \nu_\tau$	$< 4.5 \times 10^{-7}$ CL=90%
$\Gamma_{123}$	$3h^- 2h^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	$(1.65 \pm 0.11) \times 10^{-4}$
$\Gamma_{124}$	$3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	$(1.63 \pm 0.11) \times 10^{-4}$
$\Gamma_{125}$	$3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \eta, f_1(1285)$ )	$(1.11 \pm 0.10) \times 10^{-4}$
$\Gamma_{126}$	$3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \eta, \omega, f_1(1285)$ )	[a] $(3.8 \pm 0.9) \times 10^{-5}$
$\Gamma_{127}$	$K^- 2\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	[a] $(1.1 \pm 0.6) \times 10^{-6}$
$\Gamma_{128}$	$K^+ 3\pi^- \pi^+ \pi^0 \nu_\tau$	$< 8 \times 10^{-7}$ CL=90%
$\Gamma_{129}$	$3h^- 2h^+ 2\pi^0 \nu_\tau$	$< 3.4 \times 10^{-6}$ CL=90%

**Miscellaneous other allowed modes**

$\Gamma_{130}$	$(5\pi)^- \nu_\tau$	$(7.8 \pm 0.5) \times 10^{-3}$
$\Gamma_{131}$	$4h^- 3h^+ \geq 0$ neutrals $\nu_\tau$ ("7-prong")	$< 3.0 \times 10^{-7}$ CL=90%
$\Gamma_{132}$	$4h^- 3h^+ \nu_\tau$	$< 4.3 \times 10^{-7}$ CL=90%
$\Gamma_{133}$	$4h^- 3h^+ \pi^0 \nu_\tau$	$< 2.5 \times 10^{-7}$ CL=90%
$\Gamma_{134}$	$X^- (S=-1) \nu_\tau$	$(2.92 \pm 0.04) \%$
$\Gamma_{135}$	$K^*(892)^- \geq 0$ neutrals $\geq 0K_L^0 \nu_\tau$	$(1.42 \pm 0.18) \%$ S=1.4
$\Gamma_{136}$	$K^*(892)^- \nu_\tau$	$(1.20 \pm 0.07) \%$ S=1.8
$\Gamma_{137}$	$K^*(892)^- \nu_\tau \rightarrow \pi^- \bar{K}^0 \nu_\tau$	$(7.82 \pm 0.26) \times 10^{-3}$
$\Gamma_{138}$	$K^*(892)^0 K^- \geq 0$ neutrals $\nu_\tau$	$(3.2 \pm 1.4) \times 10^{-3}$
$\Gamma_{139}$	$K^*(892)^0 K^- \nu_\tau$	$(2.1 \pm 0.4) \times 10^{-3}$
$\Gamma_{140}$	$\bar{K}^*(892)^0 \pi^- \geq 0$ neutrals $\nu_\tau$	$(3.8 \pm 1.7) \times 10^{-3}$
$\Gamma_{141}$	$\bar{K}^*(892)^0 \pi^- \nu_\tau$	$(2.2 \pm 0.5) \times 10^{-3}$
$\Gamma_{142}$	$(\bar{K}^*(892)\pi)^- \nu_\tau \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$	$(1.0 \pm 0.4) \times 10^{-3}$
$\Gamma_{143}$	$K_1(1270)^- \nu_\tau$	$(4.7 \pm 1.1) \times 10^{-3}$
$\Gamma_{144}$	$K_1(1400)^- \nu_\tau$	$(1.7 \pm 2.6) \times 10^{-3}$ S=1.7
$\Gamma_{145}$	$K^*(1410)^- \nu_\tau$	$(1.5 \pm 1.4) \times 10^{-3}$
$\Gamma_{146}$	$K_0^*(1430)^- \nu_\tau$	$< 5 \times 10^{-4}$ CL=95%
$\Gamma_{147}$	$K_2^*(1430)^- \nu_\tau$	$< 3 \times 10^{-3}$ CL=95%
$\Gamma_{148}$	$a_0(980)^- \geq 0$ neutrals $\nu_\tau$	

$\Gamma_{149}$	$\eta\pi^-\nu_\tau$	$< 9.9 \times 10^{-5}$	CL=95%
$\Gamma_{150}$	$\eta\pi^-\pi^0\nu_\tau$	[a] $(1.39 \pm 0.07) \times 10^{-3}$	
$\Gamma_{151}$	$\eta\pi^-\pi^0\pi^0\nu_\tau$	[a] $(2.0 \pm 0.4) \times 10^{-4}$	
$\Gamma_{152}$	$\eta K^-\nu_\tau$	[a] $(1.55 \pm 0.08) \times 10^{-4}$	
$\Gamma_{153}$	$\eta K^*(892)^-\nu_\tau$	$(1.38 \pm 0.15) \times 10^{-4}$	
$\Gamma_{154}$	$\eta K^-\pi^0\nu_\tau$	[a] $(4.8 \pm 1.2) \times 10^{-5}$	
$\Gamma_{155}$	$\eta K^-\pi^0(\text{non-}K^*(892))\nu_\tau$	$< 3.5 \times 10^{-5}$	CL=90%
$\Gamma_{156}$	$\eta\bar{K}^0\pi^-\nu_\tau$	[a] $(9.4 \pm 1.5) \times 10^{-5}$	
$\Gamma_{157}$	$\eta\bar{K}^0\pi^-\pi^0\nu_\tau$	$< 5.0 \times 10^{-5}$	CL=90%
$\Gamma_{158}$	$\eta K^-K^0\nu_\tau$	$< 9.0 \times 10^{-6}$	CL=90%
$\Gamma_{159}$	$\eta\pi^+\pi^-\pi^- \geq 0 \text{ neutrals } \nu_\tau$	$< 3 \times 10^{-3}$	CL=90%
$\Gamma_{160}$	$\eta\pi^-\pi^+\pi^-\nu_\tau (\text{ex. } K^0)$	[a] $(2.20 \pm 0.13) \times 10^{-4}$	
$\Gamma_{161}$	$\eta\pi^-\pi^+\pi^-\nu_\tau (\text{ex. } K^0, f_1(1285))$	$(9.9 \pm 1.6) \times 10^{-5}$	
$\Gamma_{162}$	$\eta a_1(1260)^-\nu_\tau \rightarrow \eta\pi^-\rho^0\nu_\tau$	$< 3.9 \times 10^{-4}$	CL=90%
$\Gamma_{163}$	$\eta\eta\pi^-\nu_\tau$	$< 7.4 \times 10^{-6}$	CL=90%
$\Gamma_{164}$	$\eta\eta\pi^-\pi^0\nu_\tau$	$< 2.0 \times 10^{-4}$	CL=95%
$\Gamma_{165}$	$\eta\eta K^-\nu_\tau$	$< 3.0 \times 10^{-6}$	CL=90%
$\Gamma_{166}$	$\eta'(958)\pi^-\nu_\tau$	$< 4.0 \times 10^{-6}$	CL=90%
$\Gamma_{167}$	$\eta'(958)\pi^-\pi^0\nu_\tau$	$< 1.2 \times 10^{-5}$	CL=90%
$\Gamma_{168}$	$\eta'(958)K^-\nu_\tau$	$< 2.4 \times 10^{-6}$	CL=90%
$\Gamma_{169}$	$\phi\pi^-\nu_\tau$	$(3.4 \pm 0.6) \times 10^{-5}$	
$\Gamma_{170}$	$\phi K^-\nu_\tau$	[a] $(4.4 \pm 1.6) \times 10^{-5}$	
$\Gamma_{171}$	$f_1(1285)\pi^-\nu_\tau$	$(3.9 \pm 0.5) \times 10^{-4}$	S=1.9
$\Gamma_{172}$	$f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau$	$(1.18 \pm 0.07) \times 10^{-4}$	S=1.3
$\Gamma_{173}$	$f_1(1285)\pi^-\nu_\tau \rightarrow 3\pi^-2\pi^+\nu_\tau$	[a] $(5.2 \pm 0.4) \times 10^{-5}$	
$\Gamma_{174}$	$\pi(1300)^-\nu_\tau \rightarrow (\rho\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau$	$< 1.0 \times 10^{-4}$	CL=90%
$\Gamma_{175}$	$\pi(1300)^-\nu_\tau \rightarrow ((\pi\pi)_{S-\text{wave}}\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau$	$< 1.9 \times 10^{-4}$	CL=90%
$\Gamma_{176}$	$h^-\omega \geq 0 \text{ neutrals } \nu_\tau$	$(2.40 \pm 0.08) \%$	
$\Gamma_{177}$	$h^-\omega\nu_\tau$	$(1.99 \pm 0.06) \%$	
$\Gamma_{178}$	$\pi^-\omega\nu_\tau$	[a] $(1.95 \pm 0.06) \%$	
$\Gamma_{179}$	$K^-\omega\nu_\tau$	[a] $(4.1 \pm 0.9) \times 10^{-4}$	
$\Gamma_{180}$	$h^-\omega\pi^0\nu_\tau$	[a] $(4.1 \pm 0.4) \times 10^{-3}$	
$\Gamma_{181}$	$h^-\omega 2\pi^0\nu_\tau$	$(1.4 \pm 0.5) \times 10^{-4}$	
$\Gamma_{182}$	$\pi^-\omega 2\pi^0\nu_\tau$	[a] $(7.2 \pm 1.6) \times 10^{-5}$	
$\Gamma_{183}$	$h^-\omega\nu_\tau$	$< 5.4 \times 10^{-7}$	CL=90%
$\Gamma_{184}$	$2h^-h^+\omega\nu_\tau$	$(1.20 \pm 0.22) \times 10^{-4}$	
$\Gamma_{185}$	$2\pi^-\pi^+\omega\nu_\tau (\text{ex. } K^0)$	[a] $(8.4 \pm 0.6) \times 10^{-5}$	

## Lepton Family number (*LF*), Lepton number (*L*), or Baryon number (*B*) violating modes

*L* means lepton number violation (e.g.  $\tau^- \rightarrow e^+ \pi^- \pi^-$ ). Following common usage, *LF* means lepton family violation *and not* lepton number violation (e.g.  $\tau^- \rightarrow e^- \pi^+ \pi^-$ ). *B* means baryon number violation.

$\Gamma_{186}$	$e^- \gamma$	<i>LF</i>	$< 3.3$	$\times 10^{-8}$	CL=90%
$\Gamma_{187}$	$e^- \gamma\gamma$		$< 2.5$	$\times 10^{-4}$	CL=90%
$\Gamma_{188}$	$\mu^- \gamma$	<i>LF</i>	$< 4.2$	$\times 10^{-8}$	CL=90%
$\Gamma_{189}$	$\mu^- \gamma\gamma$		$< 5.8$	$\times 10^{-4}$	CL=90%
$\Gamma_{190}$	$e^- \pi^0$	<i>LF</i>	$< 8.0$	$\times 10^{-8}$	CL=90%
$\Gamma_{191}$	$\mu^- \pi^0$	<i>LF</i>	$< 1.1$	$\times 10^{-7}$	CL=90%
$\Gamma_{192}$	$e^- K_S^0$	<i>LF</i>	$< 2.6$	$\times 10^{-8}$	CL=90%
$\Gamma_{193}$	$\mu^- K_S^0$	<i>LF</i>	$< 2.3$	$\times 10^{-8}$	CL=90%
$\Gamma_{194}$	$e^- \eta$	<i>LF</i>	$< 9.2$	$\times 10^{-8}$	CL=90%
$\Gamma_{195}$	$\mu^- \eta$	<i>LF</i>	$< 6.5$	$\times 10^{-8}$	CL=90%
$\Gamma_{196}$	$e^- \rho^0$	<i>LF</i>	$< 1.8$	$\times 10^{-8}$	CL=90%
$\Gamma_{197}$	$\mu^- \rho^0$	<i>LF</i>	$< 1.2$	$\times 10^{-8}$	CL=90%
$\Gamma_{198}$	$e^- \omega$	<i>LF</i>	$< 4.8$	$\times 10^{-8}$	CL=90%
$\Gamma_{199}$	$\mu^- \omega$	<i>LF</i>	$< 4.7$	$\times 10^{-8}$	CL=90%
$\Gamma_{200}$	$e^- K^*(892)^0$	<i>LF</i>	$< 3.2$	$\times 10^{-8}$	CL=90%
$\Gamma_{201}$	$\mu^- K^*(892)^0$	<i>LF</i>	$< 5.9$	$\times 10^{-8}$	CL=90%
$\Gamma_{202}$	$e^- \bar{K}^*(892)^0$	<i>LF</i>	$< 3.4$	$\times 10^{-8}$	CL=90%
$\Gamma_{203}$	$\mu^- \bar{K}^*(892)^0$	<i>LF</i>	$< 7.0$	$\times 10^{-8}$	CL=90%
$\Gamma_{204}$	$e^- \eta'(958)$	<i>LF</i>	$< 1.6$	$\times 10^{-7}$	CL=90%
$\Gamma_{205}$	$\mu^- \eta'(958)$	<i>LF</i>	$< 1.3$	$\times 10^{-7}$	CL=90%
$\Gamma_{206}$	$e^- f_0(980) \rightarrow e^- \pi^+ \pi^-$	<i>LF</i>	$< 3.2$	$\times 10^{-8}$	CL=90%
$\Gamma_{207}$	$\mu^- f_0(980) \rightarrow \mu^- \pi^+ \pi^-$	<i>LF</i>	$< 3.4$	$\times 10^{-8}$	CL=90%
$\Gamma_{208}$	$e^- \phi$	<i>LF</i>	$< 3.1$	$\times 10^{-8}$	CL=90%
$\Gamma_{209}$	$\mu^- \phi$	<i>LF</i>	$< 8.4$	$\times 10^{-8}$	CL=90%
$\Gamma_{210}$	$e^- e^+ e^-$	<i>LF</i>	$< 2.7$	$\times 10^{-8}$	CL=90%
$\Gamma_{211}$	$e^- \mu^+ \mu^-$	<i>LF</i>	$< 2.7$	$\times 10^{-8}$	CL=90%
$\Gamma_{212}$	$e^+ \mu^- \mu^-$	<i>LF</i>	$< 1.7$	$\times 10^{-8}$	CL=90%
$\Gamma_{213}$	$\mu^- e^+ e^-$	<i>LF</i>	$< 1.8$	$\times 10^{-8}$	CL=90%
$\Gamma_{214}$	$\mu^+ e^- e^-$	<i>LF</i>	$< 1.5$	$\times 10^{-8}$	CL=90%
$\Gamma_{215}$	$\mu^- \mu^+ \mu^-$	<i>LF</i>	$< 2.1$	$\times 10^{-8}$	CL=90%
$\Gamma_{216}$	$e^- \pi^+ \pi^-$	<i>LF</i>	$< 2.3$	$\times 10^{-8}$	CL=90%
$\Gamma_{217}$	$e^+ \pi^- \pi^-$	<i>L</i>	$< 2.0$	$\times 10^{-8}$	CL=90%
$\Gamma_{218}$	$\mu^- \pi^+ \pi^-$	<i>LF</i>	$< 2.1$	$\times 10^{-8}$	CL=90%
$\Gamma_{219}$	$\mu^+ \pi^- \pi^-$	<i>L</i>	$< 3.9$	$\times 10^{-8}$	CL=90%
$\Gamma_{220}$	$e^- \pi^+ K^-$	<i>LF</i>	$< 3.7$	$\times 10^{-8}$	CL=90%
$\Gamma_{221}$	$e^- \pi^- K^+$	<i>LF</i>	$< 3.1$	$\times 10^{-8}$	CL=90%
$\Gamma_{222}$	$e^+ \pi^- K^-$	<i>L</i>	$< 3.2$	$\times 10^{-8}$	CL=90%
$\Gamma_{223}$	$e^- K_S^0 K_S^0$	<i>LF</i>	$< 7.1$	$\times 10^{-8}$	CL=90%
$\Gamma_{224}$	$e^- K^+ K^-$	<i>LF</i>	$< 3.4$	$\times 10^{-8}$	CL=90%

$\Gamma_{225}$	$e^+ K^- K^-$	$L$	$< 3.3$	$\times 10^{-8}$	CL=90%
$\Gamma_{226}$	$\mu^- \pi^+ K^-$	$LF$	$< 8.6$	$\times 10^{-8}$	CL=90%
$\Gamma_{227}$	$\mu^- \pi^- K^+$	$LF$	$< 4.5$	$\times 10^{-8}$	CL=90%
$\Gamma_{228}$	$\mu^+ \pi^- K^-$	$L$	$< 4.8$	$\times 10^{-8}$	CL=90%
$\Gamma_{229}$	$\mu^- K_S^0 K_S^0$	$LF$	$< 8.0$	$\times 10^{-8}$	CL=90%
$\Gamma_{230}$	$\mu^- K^+ K^-$	$LF$	$< 4.4$	$\times 10^{-8}$	CL=90%
$\Gamma_{231}$	$\mu^+ K^- K^-$	$L$	$< 4.7$	$\times 10^{-8}$	CL=90%
$\Gamma_{232}$	$e^- \pi^0 \pi^0$	$LF$	$< 6.5$	$\times 10^{-6}$	CL=90%
$\Gamma_{233}$	$\mu^- \pi^0 \pi^0$	$LF$	$< 1.4$	$\times 10^{-5}$	CL=90%
$\Gamma_{234}$	$e^- \eta \eta$	$LF$	$< 3.5$	$\times 10^{-5}$	CL=90%
$\Gamma_{235}$	$\mu^- \eta \eta$	$LF$	$< 6.0$	$\times 10^{-5}$	CL=90%
$\Gamma_{236}$	$e^- \pi^0 \eta$	$LF$	$< 2.4$	$\times 10^{-5}$	CL=90%
$\Gamma_{237}$	$\mu^- \pi^0 \eta$	$LF$	$< 2.2$	$\times 10^{-5}$	CL=90%
$\Gamma_{238}$	$p e^- e^-$	$L, B$	$< 3.0$	$\times 10^{-8}$	CL=90%
$\Gamma_{239}$	$\bar{p} e^+ e^-$	$L, B$	$< 3.0$	$\times 10^{-8}$	CL=90%
$\Gamma_{240}$	$\bar{p} e^+ \mu^-$	$L, B$	$< 2.0$	$\times 10^{-8}$	CL=90%
$\Gamma_{241}$	$\bar{p} e^- \mu^+$	$L, B$	$< 1.8$	$\times 10^{-8}$	CL=90%
$\Gamma_{242}$	$p \mu^- \mu^-$	$L, B$	$< 4.0$	$\times 10^{-8}$	CL=90%
$\Gamma_{243}$	$\bar{p} \mu^+ \mu^-$	$L, B$	$< 1.8$	$\times 10^{-8}$	CL=90%
$\Gamma_{244}$	$\bar{p} \gamma$	$L, B$	$< 3.5$	$\times 10^{-6}$	CL=90%
$\Gamma_{245}$	$\bar{p} \pi^0$	$L, B$	$< 1.5$	$\times 10^{-5}$	CL=90%
$\Gamma_{246}$	$\bar{p} 2\pi^0$	$L, B$	$< 3.3$	$\times 10^{-5}$	CL=90%
$\Gamma_{247}$	$\bar{p} \eta$	$L, B$	$< 8.9$	$\times 10^{-6}$	CL=90%
$\Gamma_{248}$	$\bar{p} \pi^0 \eta$	$L, B$	$< 2.7$	$\times 10^{-5}$	CL=90%
$\Gamma_{249}$	$\Lambda \pi^-$	$L, B$	$< 7.2$	$\times 10^{-8}$	CL=90%
$\Gamma_{250}$	$\bar{\Lambda} \pi^-$	$L, B$	$< 1.4$	$\times 10^{-7}$	CL=90%
$\Gamma_{251}$	$e^- \text{light boson}$	$LF$	$< 2.7$	$\times 10^{-3}$	CL=95%
$\Gamma_{252}$	$\mu^- \text{light boson}$	$LF$	$< 5$	$\times 10^{-3}$	CL=95%

[a] Basis mode for the  $\tau$ .

[b] See the Particle Listings below for the energy limits used in this measurement.

## CONSTRAINED FIT INFORMATION

An overall fit to 87 branching ratios uses 170 measurements and one constraint to determine 46 parameters. The overall fit has a  $\chi^2 = 135.0$  for 125 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \cdot \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ .

$x_5$	18									
$x_9$	2	-1								
$x_{10}$	3	4	5							
$x_{14}$	-18	-19	-17	-5						
$x_{16}$	-1	-1	0	-2	-9					
$x_{20}$	-11	-11	-14	-4	-46	-1				
$x_{23}$	-1	0	-2	-3	-1	-14	-10			
$x_{27}$	-6	-5	-10	-1	0	1	-39	1		
$x_{28}$	-1	-1	-1	-2	0	-13	-3	-23	-11	
$x_{30}$	-4	-3	-11	-1	-9	0	7	-2	-44	2
$x_{36}$	-2	-2	-3	-1	-1	-1	-2	0	-1	0
$x_{38}$	-1	-1	1	0	0	0	0	-2	0	-2
$x_{41}$	-2	-2	-2	-1	-1	0	-2	0	-1	0
$x_{43}$	-1	-1	-1	-1	0	-3	0	-5	0	-5
$x_{45}$	-5	-5	-5	-2	-3	-1	-5	-2	-1	-2
$x_{48}$	-1	-1	2	0	-1	2	-1	-1	0	-1
$x_{49}$	-5	-5	-5	-2	-3	-1	-5	-2	-1	-2
$x_{52}$	0	0	0	0	0	0	0	-1	0	-1
$x_{56}$	-2	-2	-2	-1	-1	-1	-2	-1	-1	-1
$x_{61}$	-5	-5	-5	-2	-3	-1	-4	-2	-1	-2
$x_{70}$	-7	-9	4	-2	-6	3	-12	-2	-7	-1
$x_{78}$	-4	-4	-5	0	-9	0	1	1	-1	1
$x_{85}$	0	0	-2	0	-2	0	0	0	2	0
$x_{89}$	0	0	0	0	0	0	0	0	0	0
$x_{97}$	-2	-2	-1	-1	-1	-1	-4	-1	-2	-1
$x_{103}$	1	1	0	-1	1	-1	-1	-1	0	-1
$x_{106}$	-1	-2	2	-1	-1	1	-2	-1	-1	-1
$x_{107}$	0	0	0	0	0	0	0	0	0	0
$x_{119}$	-1	-1	1	0	-1	1	-2	-1	-1	0
$x_{120}$	0	0	0	0	0	0	0	0	0	0
$x_{126}$	0	0	0	0	0	0	0	0	0	0
$x_{127}$	0	0	0	0	0	0	0	0	0	0
$x_{150}$	-1	-1	-1	0	-1	0	-2	-1	0	-1
$x_{151}$	-1	-1	0	0	-1	0	-1	0	0	0
$x_{152}$	0	0	0	0	0	0	0	-1	0	-1
$x_{154}$	0	0	0	0	0	0	0	0	0	0
$x_{156}$	0	0	0	0	0	0	0	0	0	0
$x_{160}$	-1	-1	1	0	-1	1	-1	0	-1	0
$x_{170}$	0	0	0	0	0	0	0	0	0	0
$x_{173}$	0	-1	1	0	0	1	-1	0	0	0
$x_{178}$	-3	-3	-3	-1	-4	-1	-1	0	-1	0
$x_{179}$	0	0	0	0	0	0	0	0	0	0
$x_{180}$	-2	-2	-5	-1	-3	0	-2	-1	2	-1
$x_{182}$	0	0	0	0	0	0	0	0	0	0
$x_{185}$	-1	-1	0	0	-1	1	-1	0	0	0

	$x_{36}$	$x_{38}$	$x_{41}$	$x_{43}$	$x_{45}$	$x_{48}$	$x_{49}$	$x_{52}$	$x_{56}$
$x_{36}$	0								
$x_{38}$	0	-15							
$x_{41}$	0	-13	2						
$x_{43}$	0	-1	-14	-20					
$x_{45}$	0	-3	0	-6	0				
$x_{48}$	0	-2	3	-4	1	0			
$x_{49}$	0	-5	0	-4	-1	-10	-1		
$x_{52}$	0	1	5	-1	6	0	-7	0	
$x_{56}$	0	-2	0	-2	-1	-4	0	-8	0
$x_{61}$	0	-2	0	-2	0	-4	0	-4	0
$x_{70}$	-5	-3	3	-2	-1	-4	5	-4	0
$x_{78}$	3	1	-1	1	0	2	-1	2	0
$x_{85}$	2	0	0	0	0	0	0	0	0
$x_{89}$	0	0	0	0	0	0	0	-1	0
$x_{97}$	-1	-1	0	-1	0	-2	0	-2	0
$x_{103}$	-1	-1	0	-1	0	-1	0	-1	0
$x_{106}$	-1	-1	1	0	0	-1	2	-1	0
$x_{107}$	0	0	0	0	0	0	0	0	0
$x_{119}$	-1	-1	1	0	0	-1	2	-1	0
$x_{120}$	0	0	0	0	0	0	0	0	0
$x_{126}$	0	0	0	0	0	0	0	0	0
$x_{127}$	0	0	0	0	0	0	0	0	0
$x_{150}$	-2	-1	0	0	0	-1	0	-1	0
$x_{151}$	0	0	0	0	0	-1	0	-1	0
$x_{152}$	0	0	1	0	0	0	1	0	0
$x_{154}$	0	0	0	0	0	0	0	0	0
$x_{156}$	0	0	0	0	0	0	0	-1	0
$x_{160}$	-1	0	1	0	0	-1	1	-1	0
$x_{170}$	0	0	0	0	0	0	0	0	0
$x_{173}$	-1	0	1	0	0	0	1	0	0
$x_{178}$	1	0	0	0	0	-1	0	-1	0
$x_{179}$	0	0	0	0	0	0	0	0	0
$x_{180}$	2	-1	0	0	0	-1	0	-1	0
$x_{182}$	0	0	0	0	0	0	0	0	0
$x_{185}$	-1	0	1	0	0	0	1	-1	0

	$x_{70}$	$x_{78}$	$x_{85}$	$x_{89}$	$x_{97}$	$x_{103}$	$x_{106}$	$x_{107}$	$x_{119}$	$x_{120}$	$x_{126}$	$x_{127}$	$x_{150}$	$x_{151}$	$x_{152}$	$x_{154}$	$x_{156}$	$x_{160}$	$x_{170}$	$x_{173}$	$x_{178}$	$x_{179}$	$x_{180}$	$x_{182}$	$x_{185}$	$x_{61}$	$x_{70}$	$x_{78}$	$x_{85}$	$x_{89}$	$x_{97}$	$x_{103}$	$x_{106}$	$x_{107}$	$x_{119}$
$x_{70}$	-4																																		
$x_{78}$	2	-19																																	
$x_{85}$	0	-1	-8																																
$x_{89}$	0	-1	-1	0																															
$x_{97}$	-2	19	-6	0	0																														
$x_{103}$	-1	-4	-14	-1	0	-1																													
$x_{106}$	-1	15	-4	0	0	0	-1																												
$x_{107}$	0	-1	-1	0	0	0	0	-3	0																										
$x_{119}$	-1	3	-1	0	-4	-1	0	1	0																										
$x_{120}$	0	0	0	0	0	0	0	0	0	-1	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	-1							
$x_{126}$	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	3						
$x_{127}$	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	-1						
$x_{150}$	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0						
$x_{151}$	0	2	0	0	0	-11	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	9						
$x_{152}$	0	0	0	-1	0	0	0	-1	1	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	1						
$x_{154}$	0	0	0	-1	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0						
$x_{156}$	0	0	0	0	-2	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0						
$x_{160}$	-1	1	-1	0	-8	-1	0	1	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	46						
$x_{170}$	0	-1	0	0	0	1	0	0	1	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0						
$x_{173}$	0	1	0	0	-2	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	34						
$x_{178}$	-1	-9	-67	-3	0	-2	0	-2	10	0	-2	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	-1							
$x_{179}$	0	0	12	0	0	-2	-58	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0							
$x_{180}$	-1	-2	-11	-64	-1	-1	-1	-1	-1	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0							
$x_{182}$	0	0	0	0	-16	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	7							
$x_{185}$	0	1	0	0	-4	0	0	0	1	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	39							
	$x_{61}$	$x_{70}$	$x_{78}$	$x_{85}$	$x_{89}$	$x_{97}$	$x_{103}$	$x_{106}$	$x_{107}$	$x_{119}$																									
$x_{126}$	0																																		
$x_{127}$	0	-1																																	
$x_{150}$	0	0	0																																
$x_{151}$	0	2	0	0																															
$x_{152}$	0	0	0	4	0																														
$x_{154}$	0	0	0	1	0	1																													
$x_{156}$	0	0	0	2	-1	1	0	0	0																										
$x_{160}$	-1	3	-1	0	25	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0							
$x_{170}$	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0							
$x_{173}$	-1	1	0	0	4	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	20	0						
$x_{178}$	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	-1	0						
$x_{179}$	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0							
$x_{180}$	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0							
$x_{182}$	0	2	0	0	10	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	20	0						
$x_{185}$	-1	-2	-1	0	17	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	38	0						
	$x_{120}$	$x_{126}$	$x_{127}$	$x_{150}$	$x_{151}$	$x_{152}$	$x_{154}$	$x_{156}$	$x_{160}$	$x_{170}$																									

$x_{178}$	0				
$x_{179}$	0	-14			
$x_{180}$	0	-4	0		
$x_{182}$	3	0	0	0	
$x_{185}$	17	0	0	0	14

$x_{173} \quad x_{178} \quad x_{179} \quad x_{180} \quad x_{182}$

See the related review(s):

[τ Branching Fractions](#)

$$(\Gamma(\tau^+) - \Gamma(\tau^-)) / (\Gamma(\tau^+) + \Gamma(\tau^-))$$

$$\tau^\pm \rightarrow \pi^\pm K_S^0 \nu_\tau \text{ (RATE DIFFERENCE) / (RATE SUM)}$$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b><math>-0.36 \pm 0.23 \pm 0.11</math></b>	LEES	12M BABR	$476 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

### τ<sup>-</sup> BRANCHING RATIOS

$$\Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K^0 \nu_\tau \text{ ("1-prong")}) / \Gamma_{\text{total}} \quad \Gamma_1 / \Gamma$$

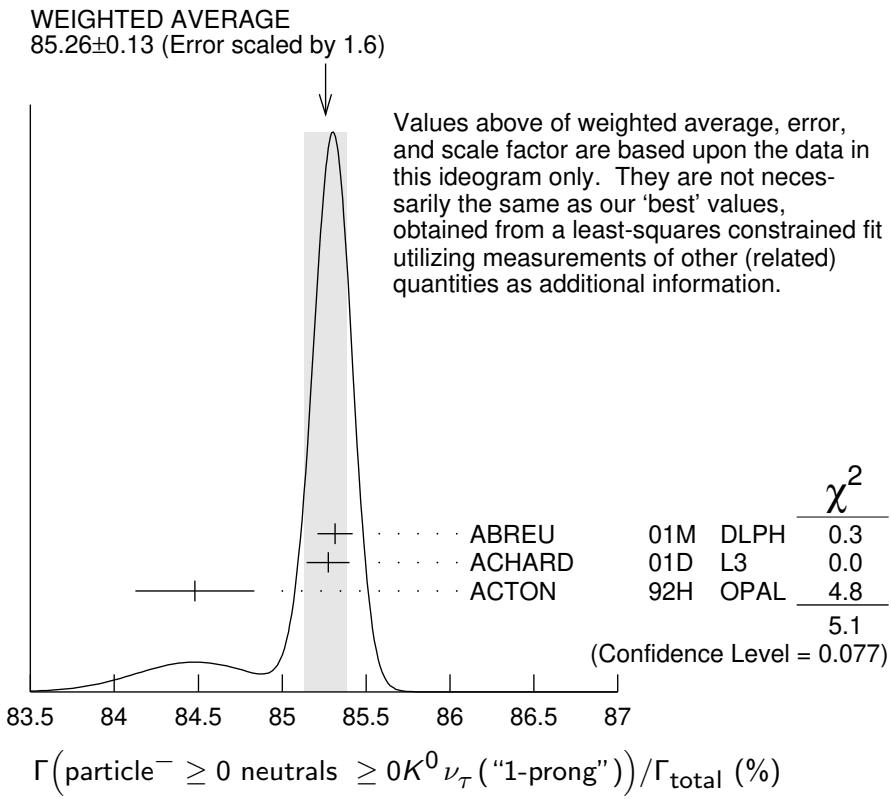
$$\Gamma_1 / \Gamma = (\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10} + \Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + \Gamma_{36} + \Gamma_{38} + \Gamma_{41} + \Gamma_{43} + \Gamma_{45} + \Gamma_{48} + \Gamma_{49} + \Gamma_{50} + \Gamma_{52} + \Gamma_{56} + \Gamma_{57} + 0.7196 \Gamma_{150} + 0.7196 \Gamma_{152} + 0.7196 \Gamma_{154} + 0.7196 \Gamma_{156} + 0.339 \Gamma_{170} + 0.0835 \Gamma_{178} + 0.0835 \Gamma_{179} + 0.0835 \Gamma_{180}) / \Gamma$$

The charged particle here can be  $e$ ,  $\mu$ , or hadron. In many analyses, the sum of the topological branching fractions (1, 3, and 5 prongs) is constrained to be unity. Since the 5-prong fraction is very small, the measured 1-prong and 3-prong fractions are highly correlated and cannot be treated as independent quantities in our overall fit. We arbitrarily choose to use the 3-prong fraction in our fit, and leave the 1-prong fraction out. We do, however, use these 1-prong measurements in our average below.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>85.24 ± 0.06 OUR FIT</b>				
<b>85.26 ± 0.13 OUR AVERAGE</b>	Error includes scale factor of 1.6. See the ideogram below.			
• • • We use the following data for averages but not for fits. • • •				
85.316 ± 0.093 ± 0.049	78k	<sup>1</sup> ABREU	01M DLPH	1992–1995 LEP runs
85.274 ± 0.105 ± 0.073		<sup>2</sup> ACHARD	01D L3	1992–1995 LEP runs
84.48 ± 0.27 ± 0.23		ACTON	92H OPAL	1990–1991 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
85.45 ± 0.69 ± 0.65		DECAMP	92C ALEP	Repl. by SCHAEFEL 05C

<sup>1</sup> The correlation coefficients between this measurement and the ABREU 01M measurements of  $B(\tau \rightarrow \text{3-prong})$  and  $B(\tau \rightarrow \text{5-prong})$  are  $-0.98$  and  $-0.08$  respectively.

<sup>2</sup> The correlation coefficients between this measurement and the ACHARD 01D measurements of  $B(\tau \rightarrow \text{"3-prong"})$  and  $B(\tau \rightarrow \text{"5-prong"})$  are  $-0.978$  and  $-0.082$  respectively.



$$\Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}} = \Gamma_2/\Gamma$$

$$\begin{aligned} \Gamma_2/\Gamma = & (\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10} + \Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.6534\Gamma_{36} + \\ & 0.6534\Gamma_{38} + 0.6534\Gamma_{41} + 0.6534\Gamma_{43} + 0.6534\Gamma_{45} + 0.0942\Gamma_{48} + 0.3069\Gamma_{49} + \Gamma_{50} + \\ & 0.0942\Gamma_{52} + 0.3069\Gamma_{56} + \Gamma_{57} + 0.7196\Gamma_{150} + 0.7196\Gamma_{152} + 0.7196\Gamma_{154} + 0.4702\Gamma_{156} + \\ & 0.1049\Gamma_{170} + 0.0835\Gamma_{178} + 0.0835\Gamma_{179} + 0.0835\Gamma_{180})/\Gamma \end{aligned}$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
-----------	------	-------------	------	---------

**84.58±0.06 OUR FIT**

**85.1 ±0.4 OUR AVERAGE**

• • • We use the following data for averages but not for fits. • • •

85.6 ±0.6 ±0.3	3300	<sup>1</sup> ADEVA	91F L3	$E_{\text{cm}}^{\text{ee}} = 88.3\text{--}94.3 \text{ GeV}$
84.9 ±0.4 ±0.3		BEHREND	89B CELL	$E_{\text{cm}}^{\text{ee}} = 14\text{--}47 \text{ GeV}$
84.7 ±0.8 ±0.6		<sup>2</sup> AIHARA	87B TPC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

86.4 ±0.3 ±0.3		ABACHI	89B HRS	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
87.1 ±1.0 ±0.7		<sup>3</sup> BURCHAT	87 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
87.2 ±0.5 ±0.8		SCHMIDKE	86 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
84.7 ±1.1 ±1.6	169	<sup>4</sup> ALTHOFF	85 TASS	$E_{\text{cm}}^{\text{ee}} = 34.5 \text{ GeV}$
86.1 ±0.5 ±0.9		BARTEL	85F JADE	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$
87.8 ±1.3 ±3.9		<sup>5</sup> BERGER	85 PLUT	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$
86.7 ±0.3 ±0.6		FERNANDEZ	85 MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

<sup>1</sup> Not independent of ADEVA 91F  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$  value.

<sup>2</sup> Not independent of AIHARA 87B  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$ ,  $\Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ , and  $\Gamma(h^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$  values.

- <sup>3</sup> Not independent of SCHMIDKE 86 value (also not independent of BURCHAT 87 value for  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$ ).
- <sup>4</sup> Not independent of ALTHOFF 85  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$ ,  $\Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$ ,  $\Gamma(h^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$ , and  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$  values.
- <sup>5</sup> Not independent of (1-prong +  $0\pi^0$ ) and (1-prong +  $\geq 1\pi^0$ ) values.

 **$\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$**  **$\Gamma_3 / \Gamma$** 

To minimize the effect of experiments with large systematic errors, we exclude experiments which together would contribute 5% of the weight in the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>17.39 ± 0.04 OUR FIT</b>				
<b>17.33 ± 0.05 OUR AVERAGE</b>				
17.319 ± 0.070 ± 0.032	54k	<sup>1</sup> SCHAEL	05C ALEP	1991-1995 LEP runs
17.34 ± 0.09 ± 0.06	31.4k	ABBIENDI	03 OPAL	1990-1995 LEP runs
17.342 ± 0.110 ± 0.067	21.5k	<sup>2</sup> ACCIARRI	01F L3	1991-1995 LEP runs
17.325 ± 0.095 ± 0.077	27.7k	ABREU	99X DLPH	1991-1995 LEP runs
• • • We use the following data for averages but not for fits. • • •				
17.37 ± 0.08 ± 0.18		<sup>3</sup> ANASTASSOV 97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
17.31 ± 0.11 ± 0.05	20.7k	BUSKULIC	96C ALEP	Repl. by SCHAEL 05C
17.02 ± 0.19 ± 0.24	6586	ABREU	95T DLPH	Repl. by ABREU 99X
17.36 ± 0.27	7941	AKERS	95I OPAL	Repl. by ABBIENDI 03
17.6 ± 0.4 ± 0.4	2148	ADRIANI	93M L3	Repl. by ACCIARRI 01F
17.4 ± 0.3 ± 0.5		<sup>4</sup> ALBRECHT	93G ARG	$E_{\text{cm}}^{\text{ee}} = 9.4-10.6 \text{ GeV}$
17.35 ± 0.41 ± 0.37		DECAMP	92C ALEP	1989-1990 LEP runs
17.7 ± 0.8 ± 0.4	568	BEHREND	90 CELL	$E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$
17.4 ± 1.0	2197	ADEVA	88 MRKJ	$E_{\text{cm}}^{\text{ee}} = 14-16 \text{ GeV}$
17.7 ± 1.2 ± 0.7		AIHARA	87B TPC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
18.3 ± 0.9 ± 0.8		BURCHAT	87 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
18.6 ± 0.8 ± 0.7	558	<sup>5</sup> BARTEL	86D JADE	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$
12.9 ± 1.7 ± 0.7		ALTHOFF	85 TASS	$E_{\text{cm}}^{\text{ee}} = 34.5 \text{ GeV}$
18.0 ± 0.9 ± 0.5	473	<sup>5</sup> ASH	85B MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
18.0 ± 1.0 ± 0.6		<sup>6</sup> BALTRUSAIT..85	MRK3	$E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$
19.4 ± 1.6 ± 1.7	153	BERGER	85 PLUT	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$
17.6 ± 2.6 ± 2.1	47	BEHREND	83C CELL	$E_{\text{cm}}^{\text{ee}} = 34 \text{ GeV}$
17.8 ± 2.0 ± 1.8		BERGER	81B PLUT	$E_{\text{cm}}^{\text{ee}} = 9-32 \text{ GeV}$

<sup>1</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> The correlation coefficient between this measurement and the ACCIARRI 01F measurement of  $B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  is 0.08.

<sup>3</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(e\bar{\nu}_e \nu_\tau)$ ,  $B(\mu\bar{\nu}_\mu \nu_\tau) / B(e\bar{\nu}_e \nu_\tau)$ ,  $B(h^- \nu_\tau)$ , and  $B(h^- \nu_\tau) / B(e\bar{\nu}_e \nu_\tau)$  are 0.50, 0.58, 0.50, and 0.08 respectively.

<sup>4</sup> Not independent of ALBRECHT 92D  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma(e^- \bar{\nu}_e \nu_\tau)$  and ALBRECHT 93G  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}^2$  values.

<sup>5</sup> Modified using  $B(e^- \bar{\nu}_e \nu_\tau) / B(\text{"1 prong"})$  and  $B(\text{"1 prong"}) = 0.855$ .

<sup>6</sup> Error correlated with BALTRUSAITIS 85  $e\nu\bar{\nu}$  value.

$\Gamma(\mu^-\bar{\nu}_\mu\nu_\tau\gamma)/\Gamma_{\text{total}}$   $\Gamma_4/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.367 ± 0.008 OUR AVERAGE</b>				
0.363 ± 0.002 ± 0.015	22k	<sup>1</sup> SHIMIZU	18A	BELL $711 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
0.369 ± 0.003 ± 0.010	16k	<sup>2</sup> LEES	15G	BABR $431 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
0.361 ± 0.016 ± 0.035		<sup>3</sup> BERGFELD	00	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.30 ± 0.04 ± 0.05	116	<sup>4</sup> ALEXANDER	96S	OPAL 1991–1994 LEP runs
0.23 ± 0.10	10	<sup>5</sup> WU	90	MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

<sup>1</sup> SHIMIZU 18A impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_{\gamma}^* > 10 \text{ MeV}$ .

<sup>2</sup> LEES 15G impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_{\gamma}^* > 10 \text{ MeV}$ .

<sup>3</sup> BERGFELD 00 impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_{\gamma}^* > 10 \text{ MeV}$ . For  $E_{\gamma}^* > 20 \text{ MeV}$ , they quote  $(3.04 \pm 0.14 \pm 0.30) \times 10^{-3}$ .

<sup>4</sup> ALEXANDER 96S impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_{\gamma}^* > 20 \text{ MeV}$ .

<sup>5</sup> WU 90 reports  $\Gamma(\mu^-\bar{\nu}_\mu\nu_\tau\gamma)/\Gamma(\mu^-\bar{\nu}_\mu\nu_\tau) = 0.013 \pm 0.006$ , which is converted to  $\Gamma(\mu^-\bar{\nu}_\mu\nu_\tau\gamma)/\Gamma_{\text{total}}$  using  $\Gamma(\mu^-\bar{\nu}_\mu\nu_\tau\gamma)/\Gamma_{\text{total}} = 17.35\%$ . Requirements on detected  $\gamma$ 's correspond to a  $\tau$  rest frame energy cutoff  $E_{\gamma} > 37 \text{ MeV}$ .

 $\Gamma(e^-\bar{\nu}_e\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_5/\Gamma$ 

To minimize the effect of experiments with large systematic errors, we exclude experiments which together would contribute 5% of the weight in the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>17.82 ± 0.04 OUR FIT</b>				
<b>17.82 ± 0.05 OUR AVERAGE</b>				
17.837 ± 0.072 ± 0.036	56k	<sup>1</sup> SCHael	05C	ALEP 1991–1995 LEP runs
17.806 ± 0.104 ± 0.076	24.7k	<sup>2</sup> ACCIARRI	01F	L3 1991–1995 LEP runs
17.81 ± 0.09 ± 0.06	33.1k	ABBIENDI	99H	OPAL 1991–1995 LEP runs
17.877 ± 0.109 ± 0.110	23.3k	ABREU	99X	DLPH 1991–1995 LEP runs
17.76 ± 0.06 ± 0.17		<sup>3</sup> ANASTASSOV	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
17.78 ± 0.10 ± 0.09	25.3k	ALEXANDER	96D	OPAL Repl. by ABBIENDI 99H
17.79 ± 0.12 ± 0.06	20.6k	BUSKULIC	96C	ALEP Repl. by SCHael 05C
17.51 ± 0.23 ± 0.31	5059	ABREU	95T	DLPH Repl.. by ABREU 99X
17.9 ± 0.4 ± 0.4	2892	ADRIANI	93M	L3 Repl. by ACCIARRI 01F
17.5 ± 0.3 ± 0.5		<sup>4</sup> ALBRECHT	93G	ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
17.97 ± 0.14 ± 0.23	3970	AKERIB	92	CLEO Repl. by ANASTASSOV 97
19.1 ± 0.4 ± 0.6	2960	<sup>5</sup> AMMAR	92	CLEO $E_{\text{cm}}^{\text{ee}} = 10.5\text{--}10.9 \text{ GeV}$
18.09 ± 0.45 ± 0.45		DECAMP	92C	ALEP Repl. by SCHael 05C
17.0 ± 0.5 ± 0.6	1.7k	ABACHI	90	HRS $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
18.4 ± 0.8 ± 0.4	644	BEHREND	90	CELL $E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$
16.3 ± 0.3 ± 3.2		JANSSEN	89	CBAL $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
18.4 ± 1.2 ± 1.0		AIHARA	87B	TPC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

19.1	$\pm 0.8$	$\pm 1.1$		BURCHAT	87	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
16.8	$\pm 0.7$	$\pm 0.9$	515	<sup>5</sup> BARTEL	86D	JADE	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$
20.4	$\pm 3.0$	$\pm 1.4$ $-0.9$		ALTHOFF	85	TASS	$E_{\text{cm}}^{\text{ee}} = 34.5 \text{ GeV}$
17.8	$\pm 0.9$	$\pm 0.6$	390	<sup>5</sup> ASH	85B	MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
18.2	$\pm 0.7$	$\pm 0.5$		<sup>6</sup> BALTRUSAIT..85	85	MRK3	$E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$
13.0	$\pm 1.9$	$\pm 2.9$		BERGER	85	PLUT	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$
18.3	$\pm 2.4$	$\pm 1.9$	60	BEHREND	83C	CELL	$E_{\text{cm}}^{\text{ee}} = 34 \text{ GeV}$
16.0	$\pm 1.3$		459	<sup>7</sup> BACINO	78B	DLCO	$E_{\text{cm}}^{\text{ee}} = 3.1\text{--}7.4 \text{ GeV}$

<sup>1</sup> Correlation matrix for SCHAEFEL 05C branching fractions, in percent:

- (1)  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$
- (2)  $\Gamma(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$
- (3)  $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau)/\Gamma_{\text{total}}$
- (4)  $\Gamma(\tau^- \rightarrow \pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$
- (5)  $\Gamma(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (6)  $\Gamma(\tau^- \rightarrow \pi^- 3\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (7)  $\Gamma(\tau^- \rightarrow h^- 4\pi^0 \nu_\tau (\text{ex. } K^0, \eta))/\Gamma_{\text{total}}$
- (8)  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0, \omega))/\Gamma_{\text{total}}$
- (9)  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (10)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (11)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ 3\pi^0 \nu_\tau)/\Gamma_{\text{total}}$
- (12)  $\Gamma(\tau^- \rightarrow 3h^- 2h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (13)  $\Gamma(\tau^- \rightarrow 3h^- 2h^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$

	(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)	(11)	(12)
(2)	-20											
(3)	-9	-6										
(4)	-16	-12	2									
(5)	-5	-5	-17	-37								
(6)	0	-4	-15	2	-27							
(7)	-2	-4	-24	-15	20	-47						
(8)	-14	-9	15	-5	-17	-14	-8					
(9)	-13	-12	-25	-30	4	-2	16	-15				
(10)	0	-2	-23	-14	4	10	13	-6	-17			
(11)	1	0	-5	1	4	6	0	-9	-2	-11		
(12)	0	1	9	4	-8	-4	-6	9	-5	-4	-2	
(13)	1	-4	-3	-5	3	2	-4	-3	-1	4	1	-24

<sup>2</sup> The correlation coefficient between this measurement and the ACCIARRI 01F measurement of  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$  is 0.08.

<sup>3</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu \bar{\nu}_\mu \nu_\tau)$ ,  $B(\mu \bar{\nu}_\mu \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$ ,  $B(h^- \nu_\tau)$ , and  $B(h^- \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$  are 0.50, -0.42, 0.48, and -0.39 respectively.

<sup>4</sup> Not independent of ALBRECHT 92D  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$  and ALBRECHT 93G  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}^2$  values.

<sup>5</sup> Modified using  $B(e^- \bar{\nu}_e \nu_\tau)/B(\text{"1 prong"})$  and  $B(\text{"1 prong"}) = 0.855$ .

<sup>6</sup> Error correlated with BALTRUSAITIS 85  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$ .

<sup>7</sup> BACINO 78B value comes from fit to events with  $e^\pm$  and one other nonelectron charged prong.

### $\Gamma(\mu^-\bar{\nu}_\mu\nu_\tau)/\Gamma(e^-\bar{\nu}_e\nu_\tau)$

### $\Gamma_3/\Gamma_5$

Standard Model prediction including mass effects is 0.9726.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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### **97.62 ± 0.28 OUR FIT**

### **97.9 ± 0.4 OUR AVERAGE**

$97.96 \pm 0.16 \pm 0.36$	731k	<sup>1</sup> AUBERT	10F	BABR $467 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$97.77 \pm 0.63 \pm 0.87$		<sup>2</sup> ANASTASSOV	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$99.7 \pm 3.5 \pm 4.0$		ALBRECHT	92D	ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$

<sup>1</sup> Correlation matrix for AUBERT 10F branching fractions:

- (1)  $\Gamma(\tau^- \rightarrow \mu^-\bar{\nu}_\mu\nu_\tau) / \Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)$
- (2)  $\Gamma(\tau^- \rightarrow \pi^-\nu_\tau) / \Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)$
- (3)  $\Gamma(\tau^- \rightarrow K^-\nu_\tau) / \Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)$

(1)	(2)
(2)	0.25
(3)	0.12    0.33

<sup>2</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu\bar{\nu}_\mu\nu_\tau)$ ,  $B(e\bar{\nu}_e\nu_\tau)$ ,  $B(h^-\nu_\tau)$ , and  $B(h^-\nu_\tau)/B(e\bar{\nu}_e\nu_\tau)$  are 0.58, -0.42, 0.07, and 0.45 respectively.

### $\Gamma(e^-\bar{\nu}_e\nu_\tau\gamma)/\Gamma_{\text{total}}$

### $\Gamma_6/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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### **1.83 ± 0.05 OUR AVERAGE**

$1.79 \pm 0.02 \pm 0.10$	12k	<sup>1</sup> SHIMIZU	18A	BELL $711 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$1.847 \pm 0.015 \pm 0.052$	18k	<sup>2</sup> LEES	15G	BABR $431 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$1.75 \pm 0.06 \pm 0.17$		<sup>3</sup> BERGFELD	00	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> SHIMIZU 18A impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10 \text{ MeV}$ .

<sup>2</sup> LEES 15G impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10 \text{ MeV}$ .

<sup>3</sup> BERGFELD 00 impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10 \text{ MeV}$ .

### $\Gamma(h^- \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$

### $\Gamma_7/\Gamma$

$$\Gamma_7/\Gamma = (\Gamma_9 + \Gamma_{10} + \frac{1}{2}\Gamma_{36} + \frac{1}{2}\Gamma_{38} + \Gamma_{50})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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### **12.03 ± 0.05 OUR FIT**

### **12.2 ± 0.4 OUR AVERAGE**

$12.47 \pm 0.26 \pm 0.43$	2967	<sup>1</sup> ACCIARRI	95	L3    1992 LEP run
$12.4 \pm 0.7 \pm 0.7$	283	<sup>2</sup> ABREU	92N	DLPH    1990 LEP run
$12.1 \pm 0.7 \pm 0.5$	309	ALEXANDER	91D	OPAL    1990 LEP run
$\bullet \bullet \bullet$ We use the following data for averages but not for fits. $\bullet \bullet \bullet$				
$11.3 \pm 0.5 \pm 0.8$	798	<sup>3</sup> FORD	87	MAC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$12.44 \pm 0.11 \pm 0.11$	15k	<sup>4</sup> BUSKULIC	96	ALEP	Repl. by SCHael 05C
$11.7 \pm 0.6 \pm 0.8$		<sup>5</sup> ALBRECHT	92D	ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
$12.98 \pm 0.44 \pm 0.33$		<sup>6</sup> DECOMP	92C	ALEP	Repl. by SCHael 05C
$12.3 \pm 0.9 \pm 0.5$	1338	BEHREND	90	CELL	$E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$
$11.1 \pm 1.1 \pm 1.4$		<sup>7</sup> BURCHAT	87	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
$12.3 \pm 0.6 \pm 1.1$	328	<sup>8</sup> BARTEL	86D	JADE	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$
$13.0 \pm 2.0 \pm 4.0$		BERGER	85	PLUT	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$
$11.2 \pm 1.7 \pm 1.2$	34	<sup>9</sup> BEHREND	83C	CELL	$E_{\text{cm}}^{\text{ee}} = 34 \text{ GeV}$

<sup>1</sup> ACCIARRI 95 with 0.65% added to remove their correction for  $\pi^- K_L^0$  backgrounds.

<sup>2</sup> ABREU 92N with 0.5% added to remove their correction for  $K^*(892)^-$  backgrounds.

<sup>3</sup> FORD 87 result for  $B(\pi^- \nu_\tau)$  with 0.67% added to remove their  $K^-$  correction and adjusted for 1992  $B$  ("1 prong").

<sup>4</sup> BUSKULIC 96 quote  $11.78 \pm 0.11 \pm 0.13$  We add 0.66 to undo their correction for unseen  $K_L^0$  and modify the systematic error accordingly.

<sup>5</sup> Not independent of ALBRECHT 92D  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$ ,  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau)$ , and  $\Gamma(h^- \geq 0 K_L^0 \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$  values.

<sup>6</sup> DECOMP 92C quote  $B(h^- \geq 0 K_L^0 \geq 0 (K_S^0 \rightarrow \pi^+ \pi^-) \nu_\tau) = 13.32 \pm 0.44 \pm 0.33$ .

We subtract 0.35 to correct for their inclusion of the  $K_S^0$  decays.

<sup>7</sup> BURCHAT 87 with 1.1% added to remove their correction for  $K^-$  and  $K^*(892)^-$  backgrounds.

<sup>8</sup> BARTEL 86D result for  $B(\pi^- \nu_\tau)$  with 0.59% added to remove their  $K^-$  correction and adjusted for 1992  $B$  ("1 prong").

<sup>9</sup> BEHREND 83C quote  $B(\pi^- \nu_\tau) = 9.9 \pm 1.7 \pm 1.3$  after subtracting  $1.3 \pm 0.5$  to correct for  $B(K^- \nu_\tau)$ .

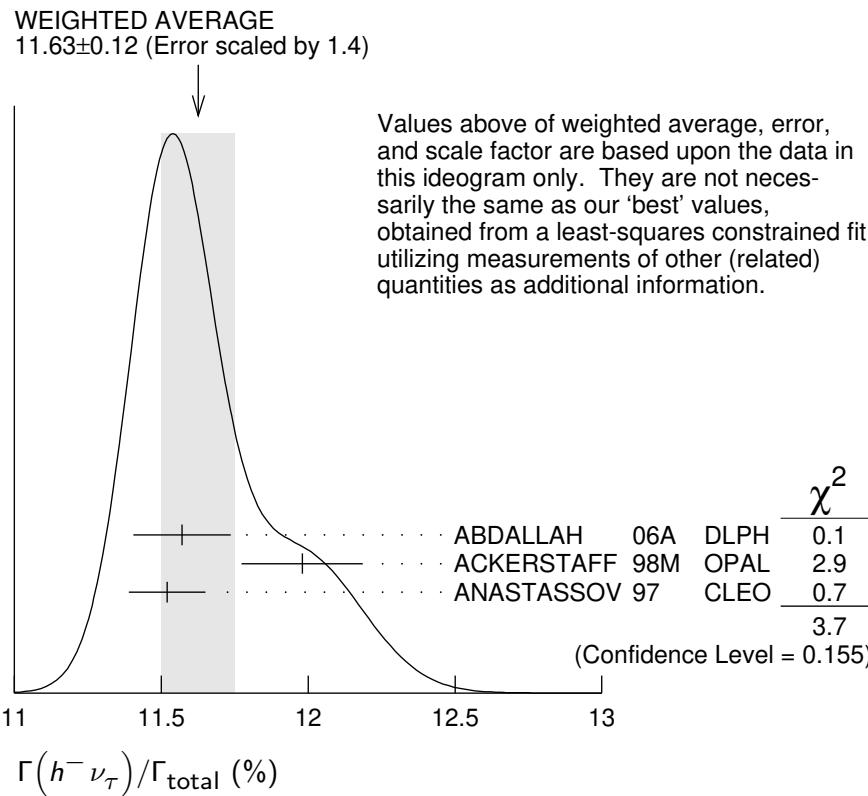
$\Gamma(h^- \nu_\tau)/\Gamma_{\text{total}}$		$\Gamma_8/\Gamma = (\Gamma_9 + \Gamma_{10})/\Gamma$		
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>11.51 <math>\pm 0.05</math> OUR FIT</b>				
<b>11.63 <math>\pm 0.12</math> OUR AVERAGE</b>		Error includes scale factor of 1.4. See the ideogram below.		
11.571 $\pm 0.120 \pm 0.114$	19k	<sup>1</sup> ABDALLAH 06A	DLPH	1992–1995 LEP runs
11.98 $\pm 0.13 \pm 0.16$		ACKERSTAFF 98M	OPAL	1991–1995 LEP runs
11.52 $\pm 0.05 \pm 0.12$		<sup>2</sup> ANASTASSOV 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> Correlation matrix for ABDALLAH 06A branching fractions, in percent:

- (1)  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$
- (2)  $\Gamma(\tau^- \rightarrow h^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$
- (3)  $\Gamma(\tau^- \rightarrow h^- \geq 1 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (4)  $\Gamma(\tau^- \rightarrow h^- 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (5)  $\Gamma(\tau^- \rightarrow h^- \geq 3\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (6)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (7)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (8)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \geq 1 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (9)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \geq 2 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (10)  $\Gamma(\tau^- \rightarrow 3h^- 2h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (11)  $\Gamma(\tau^- \rightarrow 3h^- 2h^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$

	(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)
(2)	-34									
(3)	-47	56								
(4)	6	-66	15							
(5)	-6	38	11	-86						
(6)	-7	-8	15	0	-2					
(7)	-2	-1	-5	-3	3	-53				
(8)	-4	-4	-13	-4	-2	-56	75			
(9)	-1	-1	-4	3	-6	26	-78	-16		
(10)	-1	-1	1	0	0	-2	-3	-1	3	
(11)	0	0	0	0	0	1	0	-5	5	-57

<sup>2</sup>The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu\bar{\nu}_\mu\nu_\tau)$ ,  $B(e\bar{\nu}_e\nu_\tau)$ ,  $B(\mu\bar{\nu}_\mu\nu_\tau)/B(e\bar{\nu}_e\nu_\tau)$ , and  $B(h^-\nu_\tau)/B(e\bar{\nu}_e\nu_\tau)$  are 0.50, 0.48, 0.07, and 0.63 respectively.



### $\Gamma(h^-\nu_\tau)/\Gamma(e^-\bar{\nu}_e\nu_\tau)$

VALUE (units  $10^{-2}$ )    EVTS    DOCUMENT ID    TECN    COMMENT

**64.62±0.33 OUR FIT**

**64.0 ±0.7 OUR AVERAGE** Error includes scale factor of 1.6.

• • • We use the following data for averages but not for fits. • • •

63.33±0.14±0.61    394k    <sup>1</sup>AUBERT    10F BABR     $467 \text{ fb}^{-1}$   $E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$   
64.84±0.41±0.60    <sup>2</sup>ANASTASSOV 97 CLEO     $E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$

### $\Gamma_8/\Gamma_5 = (\Gamma_9 + \Gamma_{10})/\Gamma_5$

<sup>1</sup> Not independent of AUBERT 10F  $\Gamma(\tau^- \rightarrow \pi^-\nu_\tau)/\Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)$  and  $\Gamma(\tau^- \rightarrow K^-\nu_\tau)/\Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)$ .

<sup>2</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu\bar{\nu}_\mu\nu_\tau)$ ,  $B(e\bar{\nu}_e\nu_\tau)$ ,  $B(\mu\bar{\nu}_\mu\nu_\tau)/B(e\bar{\nu}_e\nu_\tau)$ , and  $B(h^-\nu_\tau)$  are 0.08, -0.39, 0.45, and 0.63 respectively.

$\Gamma(\pi^-\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma$				
<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>10.82 ± 0.05 OUR FIT</b>					
<b>10.828 ± 0.070 ± 0.078</b>	38k	<sup>1</sup> SCHAEL	05C ALEP	1991-1995 LEP runs	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
11.06 ± 0.11 ± 0.14		<sup>2</sup> BUSKULIC	96 ALEP	Repl. by SCHAEL 05C	
11.7 ± 0.4 ± 1.8	1138	BLOCKER	82D MRK2	$E_{\text{cm}}^{ee} = 3.5\text{--}6.7 \text{ GeV}$	
<sup>1</sup> See footnote to SCHAEL 05C $\Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)/\Gamma_{\text{total}}$ measurement for correlations with other measurements.					
<sup>2</sup> Not independent of BUSKULIC 96 $B(h^-\nu_\tau)$ and $B(K^-\nu_\tau)$ values.					

$\Gamma(\pi^-\nu_\tau)/\Gamma(e^-\bar{\nu}_e\nu_\tau)$	$\Gamma_9/\Gamma_5$				
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>60.71 ± 0.32 OUR FIT</b>					
<b>59.45 ± 0.14 ± 0.61</b>	369k	<sup>1</sup> AUBERT	10F BABR	$467 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$	
<sup>1</sup> See footnote to AUBERT 10F $\Gamma(\tau^- \rightarrow \mu^-\bar{\nu}_\mu\nu_\tau)/\Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)$ for correlations with other measurements.					

$\Gamma(K^-\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{10}/\Gamma$				
<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>0.696 ± 0.010 OUR FIT</b>					
<b>0.685 ± 0.023 OUR AVERAGE</b>					
0.658 ± 0.027 ± 0.029		<sup>1</sup> ABBIENDI	01J OPAL	1990–1995 LEP runs	
0.696 ± 0.025 ± 0.014	2032	BARATE	99K ALEP	1991–1995 LEP runs	
0.85 ± 0.18	27	ABREU	94K DLPH	LEP 1992 $Z$ data	
0.66 ± 0.07 ± 0.09	99	BATTLE	94 CLEO	$E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.72 ± 0.04 ± 0.04	728	BUSKULIC	96 ALEP	Repl. by BARATE 99K	
0.59 ± 0.18	16	MILLS	84 DLCO	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$	
1.3 ± 0.5	15	BLOCKER	82B MRK2	$E_{\text{cm}}^{ee} = 3.9\text{--}6.7 \text{ GeV}$	

<sup>1</sup> The correlation coefficient between this measurement and the ABBIENDI 01J  $B(\tau^- \rightarrow K^- \geq 0\pi^0 \geq 0K^0 \geq 0\gamma \nu_\tau)$  is 0.60.

$\Gamma(K^-\nu_\tau)/\Gamma(e^-\bar{\nu}_e\nu_\tau)$	$\Gamma_{10}/\Gamma_5$				
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>3.91 ± 0.05 OUR FIT</b>					
<b>3.882 ± 0.032 ± 0.057</b>	25k	<sup>1</sup> AUBERT	10F BABR	$467 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$	
<sup>1</sup> See footnote to AUBERT 10F $\Gamma(\tau^- \rightarrow \mu^-\bar{\nu}_\mu\nu_\tau)/\Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)$ for correlations with other measurements.					

$\Gamma(K^-\nu_\tau)/\Gamma(\pi^-\nu_\tau)$	$\Gamma_{10}/\Gamma_9$				
<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>6.44 ± 0.09 OUR FIT</b>					
• • • We use the following data for averages but not for fits. • • •					
<b>6.531 ± 0.056 ± 0.093</b>		<sup>1</sup> AUBERT	10F BABR	$467 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$	

<sup>1</sup> Not independent of AUBERT 10F  $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  and  $\Gamma(\tau^- \rightarrow K^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$ .

### $\Gamma(h^- \geq 1 \text{ neutrals} \nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_{11}/\Gamma$

$$\Gamma_{11}/\Gamma = (\Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.15344\Gamma_{36} + 0.15344\Gamma_{38} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.0942\Gamma_{48} + 0.0942\Gamma_{52} + 0.7196\Gamma_{150} + 0.7196\Gamma_{152} + 0.7196\Gamma_{154} + 0.1107\Gamma_{156} + 0.0835\Gamma_{178} + 0.0835\Gamma_{179} + 0.0835\Gamma_{180})/\Gamma$$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>37.01 ± 0.09 OUR FIT</b>			

• • • We do not use the following data for averages, fits, limits, etc. • • •

$36.14 \pm 0.33 \pm 0.58$	1 AKERS	94E	OPAL	1991–1992 LEP runs
$38.4 \pm 1.2 \pm 1.0$	2 BURCHAT	87	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
$42.7 \pm 2.0 \pm 2.9$	BERGER	85	PLUT	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$

<sup>1</sup> Not independent of ACKERSTAFF 98M  $B(h^- \pi^0 \nu_\tau)$  and  $B(h^- \geq 2\pi^0 \nu_\tau)$  values.

<sup>2</sup> BURCHAT 87 quote for  $B(\pi^\pm \geq 1 \text{ neutral} \nu_\tau) = 0.378 \pm 0.012 \pm 0.010$ . We add 0.006 to account for contribution from  $(K^{*-} \nu_\tau)$  which they fixed at BR = 0.013.

### $\Gamma(h^- \geq 1\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$

$\Gamma_{12}/\Gamma$

$$\Gamma_{12}/\Gamma = (\Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.3257\Gamma_{150} + 0.3257\Gamma_{152} + 0.3257\Gamma_{154})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>36.51 ± 0.09 OUR FIT</b>				

• • • We use the following data for averages but not for fits. • • •

<b>36.641 ± 0.155 ± 0.127</b>	45k	1 ABDALLAH	06A	DLPH	1992–1995 LEP runs
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<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

### $\Gamma(h^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_{13}/\Gamma = (\Gamma_{14} + \Gamma_{16})/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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#### **25.93 ± 0.09 OUR FIT**

#### **25.73 ± 0.16 OUR AVERAGE**

$25.67 \pm 0.01 \pm 0.39$	5.4M	FUJIKAWA	08	BELL	$72 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$25.740 \pm 0.201 \pm 0.138$	35k	1 ABDALLAH	06A	DLPH	1992–1995 LEP runs
$25.89 \pm 0.17 \pm 0.29$		ACKERSTAFF	98M	OPAL	1991–1995 LEP runs
$25.05 \pm 0.35 \pm 0.50$	6613	ACCIARRI	95	L3	1992 LEP run
$25.87 \pm 0.12 \pm 0.42$	51k	2 ARTUSO	94	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$25.76 \pm 0.15 \pm 0.13$	31k	BUSKULIC	96	ALEP	Repl. by SCHAEEL 05C
$25.98 \pm 0.36 \pm 0.52$		3 AKERS	94E	OPAL	Repl. by ACKER-STAFF 98M

$22.9 \pm 0.8 \pm 1.3$	283	4 ABREU	92N	DLPH	$E_{\text{cm}}^{\text{ee}} = 88.2\text{--}94.2 \text{ GeV}$
$23.1 \pm 0.4 \pm 0.9$	1249	5 ALBRECHT	92Q	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$

$25.02 \pm 0.64 \pm 0.88$	1849	DECAMP	92C	ALEP	1989–1990 LEP runs
$22.0 \pm 0.8 \pm 1.9$	779	ANTREASYAN	91	CBAL	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$

$22.6 \pm 1.5 \pm 0.7$	1101	BEHREND	90	CELL	$E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$
$23.1 \pm 1.9 \pm 1.6$		BEHREND	84	CELL	$E_{\text{cm}}^{\text{ee}} = 14,22 \text{ GeV}$

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> ARTUSO 94 reports the combined result from three independent methods, one of which (23% of the  $\tau^- \rightarrow h^- \pi^0 \nu_\tau$ ) is normalized to the inclusive one-prong branching fraction, taken as  $0.854 \pm 0.004$ . Renormalization to the present value causes negligible change.

<sup>3</sup> AKERS 94E quote  $(26.25 \pm 0.36 \pm 0.52) \times 10^{-2}$ ; we subtract 0.27% from their number to correct for  $\tau^- \rightarrow h^- K_L^0 \nu_\tau$ .

<sup>4</sup> ABREU 92N with 0.5% added to remove their correction for  $K^*(892)^-$  backgrounds.

<sup>5</sup> ALBRECHT 92Q with 0.5% added to remove their correction for  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  background.

### $\Gamma(\pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$

### $\Gamma_{14}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>25.49 <math>\pm 0.09</math> OUR FIT</b>				
<b>25.46 <math>\pm 0.12</math> OUR AVERAGE</b>				
$25.471 \pm 0.097 \pm 0.085$	81k	<sup>1</sup> SCHAEL	05C ALEP	1991-1995 LEP runs
$\bullet \bullet \bullet$ We use the following data for averages but not for fits. $\bullet \bullet \bullet$				
25.36 $\pm 0.44$		<sup>2</sup> ARTUSO	94 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
25.30 $\pm 0.15$ $\pm 0.13$		<sup>3</sup> BUSKULIC	96 ALEP	Repl. by SCHAEL 05C
21.5 $\pm 0.4$ $\pm 1.9$	4400	<sup>4,5</sup> ALBRECHT	88L ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
23.0 $\pm 1.3$ $\pm 1.7$	582	ADLER	87B MRK3	$E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$
25.8 $\pm 1.7$ $\pm 2.5$		<sup>6</sup> BURCHAT	87 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
22.3 $\pm 0.6$ $\pm 1.4$	629	<sup>5</sup> YELTON	86 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

<sup>1</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> Not independent of ARTUSO 94 B( $h^- \pi^0 \nu_\tau$ ) and BATTLE 94 B( $K^- \pi^0 \nu_\tau$ ) values.

<sup>3</sup> Not independent of BUSKULIC 96 B( $h^- \pi^0 \nu_\tau$ ) and B( $K^- \pi^0 \nu_\tau$ ) values.

<sup>4</sup> The authors divide by  $(\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10})/\Gamma = 0.467$  to obtain this result.

<sup>5</sup> Experiment had no hadron identification. Kaon corrections were made, but insufficient information is given to permit their removal.

<sup>6</sup> BURCHAT 87 value is not independent of YELTON 86 value. Nonresonant decays included.

### $\Gamma(\pi^- \pi^0 \text{non-}\rho(770) \nu_\tau)/\Gamma_{\text{total}}$

### $\Gamma_{15}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.3 <math>\pm 0.1 \pm 0.3</math></b>		<sup>1</sup> BEHREND	84 CELL	$E_{\text{cm}}^{\text{ee}} = 14,22 \text{ GeV}$

<sup>1</sup> BEHREND 84 assume a flat nonresonant mass distribution down to the  $\rho(770)$  mass, using events with mass above 1300 to set the level.

### $\Gamma(K^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$

### $\Gamma_{16}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.433 <math>\pm 0.015</math> OUR FIT</b>				
<b>0.426 <math>\pm 0.016</math> OUR AVERAGE</b>				
0.416 $\pm 0.003 \pm 0.018$	78k	AUBERT	07AP BABR	$230 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
0.471 $\pm 0.059 \pm 0.023$	360	ABBIENDI	04J OPAL	1991-1995 LEP runs
0.444 $\pm 0.026 \pm 0.024$	923	BARATE	99K ALEP	1991-1995 LEP runs
0.51 $\pm 0.10 \pm 0.07$	37	BATTLE	94 CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
0.52 $\pm 0.04 \pm 0.05$	395	BUSKULIC	96 ALEP	Repl. by BARATE 99K

$\Gamma(h^- \geq 2\pi^0\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{17}/\Gamma$ 

$$\Gamma_{17}/\Gamma = (\Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.15344\Gamma_{36} + 0.15344\Gamma_{38} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.09419\Gamma_{48} + 0.0942\Gamma_{52} + 0.3257\Gamma_{150} + 0.3257\Gamma_{152} + 0.3257\Gamma_{154})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**10.81±0.09 OUR FIT****9.91±0.31±0.27**

ACKERSTAFF 98M OPAL 1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

9.89 ± 0.34 ± 0.55		<sup>1</sup> AKERS 94E	OPAL	Repl. by ACKER-STAFF 98M
14.0 ± 1.2 ± 0.6	938	<sup>2</sup> BEHREND 90	CELL	$E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$
12.0 ± 1.4 ± 2.5		<sup>3</sup> BURCHAT 87	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
13.9 ± 2.0 ± 1.9		<sup>4</sup> AIHARA 86E	TPC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

<sup>1</sup> AKERS 94E not independent of AKERS 94E  $B(h^- \geq 1\pi^0\nu_\tau)$  and  $B(h^- \pi^0\nu_\tau)$  measurements.

<sup>2</sup> No independent of BEHREND 90  $\Gamma(h^- 2\pi^0\nu_\tau \text{ (exp. } K^0))$  and  $\Gamma(h^- \geq 3\pi^0\nu_\tau)$ .

<sup>3</sup> Error correlated with BURCHAT 87  $\Gamma(\rho^-\nu_e)/\Gamma(\text{total})$  value.

<sup>4</sup> AIHARA 86E (TPC) quote  $B(2\pi^0\pi^-\nu_\tau) + 1.6B(3\pi^0\pi^-\nu_\tau) + 1.1B(\pi^0\eta\pi^-\nu_\tau)$ .

 $\Gamma(h^- 2\pi^0\nu_\tau)/\Gamma_{\text{total}}$ 
 $\Gamma_{18}/\Gamma$ 

$$\Gamma_{18}/\Gamma = (\Gamma_{20} + \Gamma_{23} + 0.15344\Gamma_{36} + 0.15344\Gamma_{38})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**9.48±0.10 OUR FIT**

• • • We do not use the following data for averages, fits, limits, etc. • • •

9.48 ± 0.13 ± 0.10	12k	<sup>1</sup> BUSKULIC 96	ALEP	Repl. by SCHAEEL 05C
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<sup>1</sup> BUSKULIC 96 quote  $9.29 \pm 0.13 \pm 0.10$ . We add 0.19 to undo their correction for  $\tau^- \rightarrow h^- K^0\nu_\tau$ .

 $\Gamma(h^- 2\pi^0\nu_\tau \text{ (ex. } K^0))/\Gamma_{\text{total}}$ 
 $\Gamma_{19}/\Gamma$ 

$$\Gamma_{19}/\Gamma = (\Gamma_{20} + \Gamma_{23})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**9.32 ±0.10 OUR FIT****9.17 ±0.27 OUR AVERAGE**

9.498 ± 0.320 ± 0.275	9.5k	<sup>1</sup> ABDALLAH 06A	DLPH	1992–1995 LEP runs
8.88 ± 0.37 ± 0.42	1060	ACCIARRI	L3	1992 LEP run

• • • We use the following data for averages but not for fits. • • •

8.96 ± 0.16 ± 0.44		<sup>2</sup> PROCARIO 93	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

10.38 ± 0.66 ± 0.82	809	<sup>3</sup> DECAMP 92C	ALEP	Repl. by SCHAEEL 05C
5.7 ± 0.5 ± 1.7	133	<sup>4</sup> ANTREASYAN 91	CBAL	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
10.0 ± 1.5 ± 1.1	333	<sup>5</sup> BEHREND 90	CELL	$E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$
8.7 ± 0.4 ± 1.1	815	<sup>6</sup> BAND 87	MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
6.2 ± 0.6 ± 1.2		<sup>7</sup> GAN 87	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
6.0 ± 3.0 ± 1.8		BEHREND 84	CELL	$E_{\text{cm}}^{\text{ee}} = 14,22 \text{ GeV}$

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> PROCARIO 93 entry is obtained from  $B(h^- 2\pi^0\nu_\tau)/B(h^- \pi^0\nu_\tau)$  using ARTUSO 94 result for  $B(h^- \pi^0\nu_\tau)$ .

<sup>3</sup> We subtract 0.0015 to account for  $\tau^- \rightarrow K^*(892)^-\nu_\tau$  contribution.

<sup>4</sup> ANTREASYAN 91 subtract 0.001 to account for the  $\tau^- \rightarrow K^*(892)^-\nu_\tau$  contribution.

<sup>5</sup> BEHREND 90 subtract 0.002 to account for the  $\tau^- \rightarrow K^*(892)^-\nu_\tau$  contribution.

<sup>6</sup> BAND 87 assume  $B(\pi^- 3\pi^0\nu_\tau) = 0.01$  and  $B(\pi^- \pi^0\eta\nu_\tau) = 0.005$ .

<sup>7</sup> GAN 87 analysis use photon multiplicity distribution.

$$\frac{\Gamma(h^- 2\pi^0\nu_\tau(\text{ex. } K^0))}{\Gamma(h^- \pi^0\nu_\tau)} / \frac{\Gamma_1}{\Gamma_{13}} = \frac{(\Gamma_{20} + \Gamma_{23})}{(\Gamma_{14} + \Gamma_{16})}$$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>36.0 \pm 0.4</math> OUR FIT</b>			
<b><math>34.2 \pm 0.6 \pm 1.6</math></b>	<sup>1</sup> PROCARIO 93 CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$	

<sup>1</sup> PROCARIO 93 quote  $0.345 \pm 0.006 \pm 0.016$  after correction for 2 kaon backgrounds assuming  $B(K^*-\nu_\tau) = 1.42 \pm 0.18\%$  and  $B(h^- K^0\pi^0\nu_\tau) = 0.48 \pm 0.48\%$ . We multiply by  $0.990 \pm 0.010$  to remove these corrections to  $B(h^- \pi^0\nu_\tau)$ .

$$\frac{\Gamma(\pi^- 2\pi^0\nu_\tau(\text{ex. } K^0))}{\Gamma_{\text{total}}} / \frac{\Gamma_2}{\Gamma} = \frac{(\Gamma_{20} + \Gamma_{23})}{(\Gamma_{14} + \Gamma_{16})}$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>9.26 \pm 0.10</math> OUR FIT</b>				
<b><math>9.239 \pm 0.086 \pm 0.090</math></b>	31k	<sup>1</sup> SCHAEL 05C ALEP	1991-1995 LEP runs	

• • • We do not use the following data for averages, fits, limits, etc. • • •

$9.21 \pm 0.13 \pm 0.11$	<sup>2</sup> BUSKULIC 96 ALEP	Repl. by SCHAEL 05C
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<sup>1</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> Not independent of BUSKULIC 96  $B(h^- 2\pi^0\nu_\tau(\text{ex. } K^0))$  and  $B(K^- 2\pi^0\nu_\tau(\text{ex. } K^0))$  values.

$$\frac{\Gamma(\pi^- 2\pi^0\nu_\tau(\text{ex. } K^0), \text{scalar})}{\Gamma(\pi^- 2\pi^0\nu_\tau(\text{ex. } K^0))} / \frac{\Gamma_2}{\Gamma_{20}} = \frac{(\Gamma_{21} + \Gamma_{23})}{(\Gamma_{14} + \Gamma_{16})}$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.094</math></b>	95	<sup>1</sup> BROWDER 00 CLEO	$4.7 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	

<sup>1</sup> Model-independent limit from structure function analysis on contribution to  $B(\tau^- \rightarrow \pi^- 2\pi^0\nu_\tau(\text{ex. } K^0))$  from scalars.

$$\frac{\Gamma(\pi^- 2\pi^0\nu_\tau(\text{ex. } K^0), \text{vector})}{\Gamma(\pi^- 2\pi^0\nu_\tau(\text{ex. } K^0))} / \frac{\Gamma_2}{\Gamma_{20}} = \frac{(\Gamma_{22} + \Gamma_{23})}{(\Gamma_{14} + \Gamma_{16})}$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;0.073</math></b>	95	<sup>1</sup> BROWDER 00 CLEO	$4.7 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	

<sup>1</sup> Model-independent limit from structure function analysis on contribution to  $B(\tau^- \rightarrow \pi^- 2\pi^0\nu_\tau(\text{ex. } K^0))$  from vectors.

$$\frac{\Gamma(K^- 2\pi^0\nu_\tau(\text{ex. } K^0))}{\Gamma_{\text{total}}} / \frac{\Gamma_3}{\Gamma} = \frac{(\Gamma_{23} + \Gamma_{24})}{(\Gamma_{14} + \Gamma_{16})}$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>6.5 \pm 2.2</math> OUR FIT</b>				
<b><math>5.8 \pm 2.4</math> OUR AVERAGE</b>				

$5.6 \pm 2.0 \pm 1.5$	131	BARATE 99K ALEP	1991–1995 LEP runs
$9 \pm 10 \pm 3$	3	<sup>1</sup> BATTLE 94 CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$8 \pm 2 \pm 2$	59	BUSKULIC 96 ALEP	Repl. by BARATE 99K
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<sup>1</sup> BATTLE 94 quote  $(14 \pm 10 \pm 3) \times 10^{-4}$  or  $< 30 \times 10^{-4}$  at 90% CL. We subtract  $(5 \pm 2) \times 10^{-4}$  to account for  $\tau^- \rightarrow K^-(K^0 \rightarrow \pi^0\pi^0)\nu_\tau$  background.

$\Gamma(h^- \geq 3\pi^0\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{24}/\Gamma$ 

$$\Gamma_{24}/\Gamma = (\Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.0942\Gamma_{48} + 0.0942\Gamma_{52} + 0.3257\Gamma_{150} + 0.3257\Gamma_{152} + 0.3257\Gamma_{154} + 0.0501\Gamma_{156})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.34 ± 0.07 OUR FIT**

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.53 ± 0.40 ± 0.46	186	DECAMP	92C	ALEP Repl. by SCHAEEL 05C
3.2 ± 1.0 ± 1.0		BEHREND	90	CELL $E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$

 $\Gamma(h^- \geq 3\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$   $\Gamma_{25}/\Gamma$ 

$$\Gamma_{25}/\Gamma = (\Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.3257\Gamma_{150} + 0.3257\Gamma_{152} + 0.3257\Gamma_{154})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.25 ± 0.07 OUR FIT**

1.403 ± 0.214 ± 0.224	1.1k	1 ABDALLAH	06A	DLEPH 1992–1995 LEP runs
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<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

 $\Gamma(h^- 3\pi^0\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{26}/\Gamma$ 

$$\Gamma_{26}/\Gamma = (\Gamma_{27} + \Gamma_{28} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.3257\Gamma_{152})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.18 ± 0.07 OUR FIT**

**1.21 ± 0.17 OUR AVERAGE** Error includes scale factor of 1.2.

1.70 ± 0.24 ± 0.38	293	ACCIARRI	95	L3 1992 LEP run
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• • • We use the following data for averages but not for fits. • • •

1.15 ± 0.08 ± 0.13	1 PROCARIO	93	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1.24 ± 0.09 ± 0.11	2.3k	2 BUSKULIC	96	ALEP Repl. by SCHAEEL 05C
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0.0 +1.4 -0.1	+1.1 -0.1	3 GAN	87	MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
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<sup>1</sup> PROCARIO 93 entry is obtained from  $B(h^- 3\pi^0\nu_\tau)/B(h^- \pi^0\nu_\tau)$  using ARTUSO 94 result for  $B(h^- \pi^0\nu_\tau)$ .

<sup>2</sup> BUSKULIC 96 quote  $B(h^- 3\pi^0\nu_\tau(\text{ex. } K^0)) = 1.17 \pm 0.09 \pm 0.11$ . We add 0.07 to remove their correction for  $K^0$  backgrounds.

<sup>3</sup> Highly correlated with GAN 87  $\Gamma(\eta\pi^- \pi^0\nu_\tau)/\Gamma_{\text{total}}$  value. Authors quote  $B(\pi^\pm 3\pi^0\nu_\tau) + 0.67B(\pi^\pm \eta\pi^0\nu_\tau) = 0.047 \pm 0.010 \pm 0.011$ .

 $\Gamma(h^- 3\pi^0\nu_\tau)/\Gamma(h^- \pi^0\nu_\tau)$   $\Gamma_{26}/\Gamma_{13}$ 

$$\Gamma_{26}/\Gamma_{13} = (\Gamma_{27} + \Gamma_{28} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.3257\Gamma_{152})/(\Gamma_{14} + \Gamma_{16})$$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
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**4.54 ± 0.28 OUR FIT**

4.4 ± 0.3 ± 0.5	1 PROCARIO	93	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
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<sup>1</sup> PROCARIO 93 quote  $0.041 \pm 0.003 \pm 0.005$  after correction for 2 kaon backgrounds assuming  $B(K^{*-}\nu_\tau) = 1.42 \pm 0.18\%$  and  $B(h^- K^0 \pi^0 \nu_\tau) = 0.48 \pm 0.48\%$ . We add  $0.003 \pm 0.003$  and multiply the sum by  $0.990 \pm 0.010$  to remove these corrections.

$\Gamma(\pi^- 3\pi^0 \nu_\tau (\text{ex.} K^0)) / \Gamma_{\text{total}}$  $\Gamma_{27}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.04 ± 0.07 OUR FIT</b>				
<b>0.977 ± 0.069 ± 0.058</b>	6.1k	<sup>1</sup> SCHAEL	05C ALEP	1991-1995 LEP runs

<sup>1</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

 $\Gamma(K^- 3\pi^0 \nu_\tau (\text{ex.} K^0, \eta)) / \Gamma_{\text{total}}$  $\Gamma_{28}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.8 ± 2.1 OUR FIT</b>				
<b>3.7 ± 2.1 ± 1.1</b>	22	BARATE	99K ALEP	1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

5 ± 13		<sup>1</sup> BUSKULIC	94E ALEP	Repl. by BARATE 99K
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<sup>1</sup> BUSKULIC 94E quote  $B(K^- \geq 0\pi^0 \geq 0K^0 \nu_\tau) = [B(K^- \nu_\tau) + B(K^- \pi^0 \nu_\tau) + B(K^- K^0 \nu_\tau) + B(K^- \pi^0 \pi^0 \nu_\tau) + B(K^- \pi^0 K^0 \nu_\tau)] = (5 \pm 13) \times 10^{-4}$  accounting for common systematic errors in BUSKULIC 94E and BUSKULIC 94F measurements of these modes. We assume  $B(K^- \geq 2K^0 \nu_\tau)$  and  $B(K^- \geq 4\pi^0 \nu_\tau)$  are negligible.

 $\Gamma(h^- 4\pi^0 \nu_\tau (\text{ex.} K^0)) / \Gamma_{\text{total}}$  $\Gamma_{29}/\Gamma$ 

$$\Gamma_{29}/\Gamma = (\Gamma_{30} + 0.3257\Gamma_{150} + 0.3257\Gamma_{154})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.16 ± 0.04 OUR FIT</b>				

<b>0.16 ± 0.05 ± 0.05</b>		<sup>1</sup> PROCARIO	93 CLEO	$E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.16 ± 0.04 ± 0.09	232	<sup>2</sup> BUSKULIC	96 ALEP	Repl. by SCHAEL 05C
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<sup>1</sup> PROCARIO 93 quotes  $B(h^- 4\pi^0 \nu_\tau) / B(h^- \pi^0 \nu_\tau) = 0.006 \pm 0.002 \pm 0.002$ . We multiply by the ARTUSO 94 result for  $B(h^- \pi^0 \nu_\tau)$  to obtain  $B(h^- 4\pi^0 \nu_\tau)$ . PROCARIO 93 assume  $B(h^- \geq 5\pi^0 \nu_\tau)$  is small and do not correct for it.

<sup>2</sup> BUSKULIC 96 quote result for  $\tau^- \rightarrow h^- \geq 4\pi^0 \nu_\tau$ . We assume  $B(h^- \geq 5\pi^0 \nu_\tau)$  is negligible.

 $\Gamma(h^- 4\pi^0 \nu_\tau (\text{ex.} K^0, \eta)) / \Gamma_{\text{total}}$  $\Gamma_{30}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.11 ± 0.04 OUR FIT</b>				
<b>0.112 ± 0.037 ± 0.035</b>	957	<sup>1</sup> SCHAEL	05C ALEP	1991-1995 LEP runs

<sup>1</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

 $\Gamma(a_1(1260)\nu_\tau \rightarrow \pi^- \gamma \nu_\tau) / \Gamma_{\text{total}}$  $\Gamma_{31}/\Gamma = (0.0021\Gamma_{20} + 0.0021\Gamma_{70})/\Gamma$ 

The uncertainty on  $\Gamma(\tau^- \rightarrow a_1(1260)\nu_\tau \rightarrow \pi^- \gamma \nu_\tau) / \Gamma_{\text{total}}$  is the sum in quadrature of the uncertainty on the fit result for  $\Gamma(\tau^- \rightarrow a_1(1260)\nu_\tau \rightarrow \pi^- \gamma \nu_\tau) / \Gamma_{\text{total}}$  and of the uncertainty on  $\Gamma(a_1(1260) \rightarrow \pi\gamma) / \Gamma_{\text{total}} = ((2.1 \pm 0.8) \times 10^{-3})$  as reported in SCHAEL 05C, which is the coefficient of the relationship that defines  $\Gamma(\tau^- \rightarrow a_1(1260)\nu_\tau \rightarrow \pi^- \gamma \nu_\tau) / \Gamma_{\text{total}}$  in terms of  $\Gamma(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau (\text{ex.} K^0)) / \Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0, \omega)) / \Gamma_{\text{total}}$ .

VALUE (units $10^{-4}$ )	DOCUMENT ID
<b>3.8 ± 1.5 OUR FIT</b>	

$\Gamma(K^- \geq 0 K^0 \geq 0 \gamma \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{32}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.552 ± 0.029 OUR FIT****1.53 ± 0.04 OUR AVERAGE**

1.528 ± 0.039 ± 0.040

<sup>1</sup> ABBIENDI 01J OPAL 1990–1995 LEP runs

1.54 ± 0.24

ABREU 94K DLPH LEP 1992  $Z$  data

1.70 ± 0.12 ± 0.19 202

<sup>2</sup> BATTLE 94 CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6$  GeV

• • • We use the following data for averages but not for fits. • • •

1.520 ± 0.040 ± 0.041 4006

<sup>3</sup> BARATE 99K ALEP 1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.70 ± 0.05 ± 0.06 1610

<sup>4</sup> BUSKULIC 96 ALEP Repl. by BARATE 99K

1.6 ± 0.4 ± 0.2 35

AIHARA 87B TPC  $E_{\text{cm}}^{\text{ee}} = 29$  GeV

1.71 ± 0.29 53

MILLS 84 DLCO  $E_{\text{cm}}^{\text{ee}} = 29$  GeV

<sup>1</sup> The correlation coefficient between this measurement and the ABBIENDI 01J  $B(\tau^- \rightarrow K^- \nu_\tau)$  is 0.60.

<sup>2</sup> BATTLE 94 quote  $1.60 \pm 0.12 \pm 0.19$ . We add  $0.10 \pm 0.02$  to correct for their rejection of  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

<sup>3</sup> Not independent of BARATE 99K  $B(K^- \nu_\tau)$ ,  $B(K^- \pi^0 \nu_\tau)$ ,  $B(K^- 2\pi^0 \nu_\tau$  (ex.  $K^0$ )),  $B(K^- 3\pi^0 \nu_\tau$  (ex.  $K^0$ )),  $B(K^- K^0 \nu_\tau)$ , and  $B(K^- K^0 \pi^0 \nu_\tau)$  values.

<sup>4</sup> Not independent of BUSKULIC 96  $B(K^- \nu_\tau)$ ,  $B(K^- \pi^0 \nu_\tau)$ ,  $B(K^- 2\pi^0 \nu_\tau)$ ,  $B(K^- K^0 \nu_\tau)$ , and  $B(K^- K^0 \pi^0 \nu_\tau)$  values.

 $\Gamma(K^- \geq 1(\pi^0 \text{ or } K^0 \text{ or } \gamma) \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{33}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.859 ± 0.028 OUR FIT****0.86 ± 0.05 OUR AVERAGE**

• • • We use the following data for averages but not for fits. • • •

0.869 ± 0.031 ± 0.034

<sup>1</sup> ABBIENDI 01J OPAL 1990–1995 LEP runs

0.69 ± 0.25

<sup>2</sup> ABREU 94K DLPH LEP 1992  $Z$  data

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.2 ± 0.5 +0.2 -0.4 9 AIHARA 87B TPC  $E_{\text{cm}}^{\text{ee}} = 29$  GeV

<sup>1</sup> Not independent of ABBIENDI 01J  $B(\tau^- \rightarrow K^- \nu_\tau)$  and  $B(\tau^- \rightarrow K^- \geq 0 \pi^0 \geq 0 K^0 \geq 0 \gamma \nu_\tau)$  values.

<sup>2</sup> Not independent of ABREU 94K  $B(K^- \nu_\tau)$  and  $B(K^- \geq 0 \text{ neutrals} \nu_\tau)$  measurements.

 $\Gamma(K_S^0(\text{particles})^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{34}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.943 ± 0.028 OUR FIT****0.918 ± 0.015 OUR AVERAGE**

0.970 ± 0.058 ± 0.062 929

BARATE 98E ALEP 1991–1995 LEP runs

0.97 ± 0.09 ± 0.06 141

AKERS 94G OPAL  $E_{\text{cm}}^{\text{ee}} = 88–94$  GeV

• • • We use the following data for averages but not for fits. • • •

0.915 ± 0.001 ± 0.015 398k

<sup>1</sup> RYU 14 BELL  $669 \text{ fb}^{-1}$   $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

<sup>1</sup> Not independent of RYU 14 measurements of  $B(\tau^- \rightarrow \pi^-\bar{K}^0\nu_\tau)$ ,  $B(\tau^- \rightarrow K^-\bar{K}^0\nu_\tau)$ ,  $B(\tau^- \rightarrow \pi^-\bar{K}^0\pi^0\nu_\tau)$ ,  $B(\tau^- \rightarrow K^-\bar{K}^0\pi^0\nu_\tau)$ ,  $B(\tau^- \rightarrow \pi^-\bar{K}_S^0\bar{K}_S^0\nu_\tau)$ , and  $B(\tau^- \rightarrow \pi^-\bar{K}_S^0\bar{K}_S^0\pi^0\nu_\tau)$ .

### $\Gamma(h^-\bar{K}^0\nu_\tau)/\Gamma_{\text{total}}$

### $\Gamma_{35}/\Gamma = (\Gamma_{36} + \Gamma_{38})/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.987 ± 0.014 OUR FIT**

**0.90 ± 0.07 OUR AVERAGE**

$0.855 \pm 0.036 \pm 0.073$     1242    COAN    96    CLEO     $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

$1.01 \pm 0.11 \pm 0.07$     555    <sup>1</sup> BARATE    98E    ALEP    1991–1995 LEP runs

<sup>1</sup> Not independent of BARATE 98E  $B(\tau^- \rightarrow \pi^-\bar{K}^0\nu_\tau)$  and  $B(\tau^- \rightarrow K^-\bar{K}^0\nu_\tau)$  values.

### $\Gamma(\pi^-\bar{K}^0\nu_\tau)/\Gamma_{\text{total}}$

### $\Gamma_{36}/\Gamma$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**8.38 ± 0.14 OUR FIT**

**8.39 ± 0.22 OUR AVERAGE** Error includes scale factor of 1.5. See the ideogram below.

$8.32 \pm 0.02 \pm 0.16$     158k    <sup>1</sup> RYU    14    BELL     $669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$9.33 \pm 0.68 \pm 0.49$     377    ABBIENDI    00C    OPAL    1991–1995 LEP runs

$9.28 \pm 0.45 \pm 0.34$     937    <sup>2</sup> BARATE    99K    ALEP    1991–1995 LEP runs

$9.5 \pm 1.5 \pm 0.6$     <sup>3</sup> ACCIARRI    95F    L3    1991–1993 LEP runs

• • • We use the following data for averages but not for fits. • • •

$8.55 \pm 1.17 \pm 0.66$     509    <sup>4</sup> BARATE    98E    ALEP    1991–1995 LEP runs

$7.04 \pm 0.41 \pm 0.72$     <sup>5</sup> COAN    96    CLEO     $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

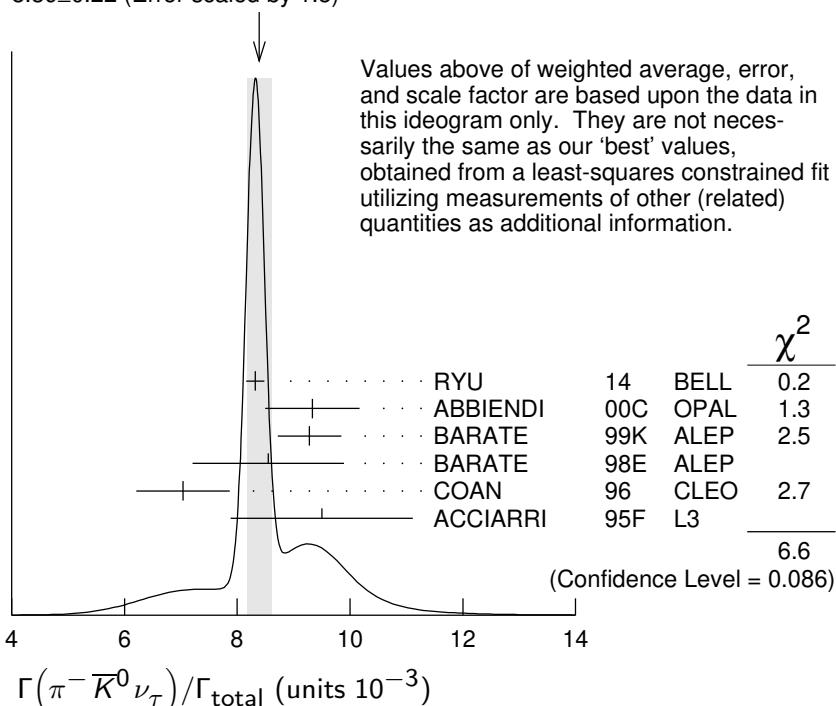
• • • We do not use the following data for averages, fits, limits, etc. • • •

$8.08 \pm 0.04 \pm 0.26$     53k    EPIFANOV    07    BELL    Repl. by RYU 14

$7.9 \pm 1.0 \pm 0.9$     98    <sup>6</sup> BUSKULIC    96    ALEP    Repl. by BARATE 99K

#### WEIGHTED AVERAGE

$8.39 \pm 0.22$  (Error scaled by 1.5)



<sup>1</sup> RYU 14 reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

<sup>2</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

<sup>3</sup> ACCIARRI 95F do not identify  $\pi^-/K^-$  and assume  $B(K^- K^0 \nu_\tau) = (0.29 \pm 0.12)\%$ .

<sup>4</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays. Not independent of BARATE 98E  $B(K^0 \text{ particles}^- \nu_\tau)$  value.

<sup>5</sup> Not independent of COAN 96  $B(h^- K^0 \nu_\tau)$  and  $B(K^- K^0 \nu_\tau)$  measurements.

<sup>6</sup> BUSKULIC 96 measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

### $\Gamma(\pi^- \bar{K}^0 (\text{non-}K^*(892)^-) \nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_{37}/\Gamma$

VALUE (units $10^{-4}$ )	CL %	DOCUMENT ID	TECN	COMMENT
<b>5.4±2.1</b>		1 EPIFANOV 07	BELL 351 fb $^{-1}$	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<17 95 ACCIARRI 95F L3 1991–1993 LEP runs

<sup>1</sup> EPIFANOV 07 quote  $B(\tau^- \rightarrow K^*(892)^- \nu_\tau) B(K^*(892)^- \rightarrow K_S^0 \pi^-) / B(\tau^- \rightarrow K_S^0 \pi^- \nu_\tau) = 0.933 \pm 0.027$ . We multiply their  $B(\tau^- \rightarrow \bar{K}^0 \pi^- \nu_\tau)$  by  $[1 - (0.933 \pm 0.027)]$  to obtain this result.

### $\Gamma(K^- K^0 \nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_{38}/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>14.86±0.34 OUR FIT</b>				
<b>14.83±0.35 OUR AVERAGE</b>				

14.78±0.22±0.40 29k 1 LEES 18B BABR 468 fb $^{-1}$   $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

14.80±0.14±0.54 33k 2 RYU 14 BELL 669 fb $^{-1}$   $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

16.2 ± 2.1 ± 1.1 150 3 BARATE 99K ALEP 1991–1995 LEP runs

15.8 ± 4.2 ± 1.7 46 4 BARATE 98E ALEP 1991–1995 LEP runs

15.1 ± 2.1 ± 2.2 111 COAN 96 CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

26 ± 9 ± 2 13 <sup>5</sup> BUSKULIC 96 ALEP Repl. by BARATE 99K

<sup>1</sup> LEES 18B reconstructs  $K_S^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

<sup>2</sup> RYU 14 reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

<sup>3</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

<sup>4</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

<sup>5</sup> BUSKULIC 96 measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

### $\Gamma(K^- K^0 \geq 0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_{39}/\Gamma = (\Gamma_{38} + \Gamma_{43})/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.299±0.007 OUR FIT</b>				
<b>0.330±0.055±0.039</b>	124	ABBIENDI	00C OPAL	1991–1995 LEP runs

### $\Gamma(h^- \bar{K}^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_{40}/\Gamma = (\Gamma_{41} + \Gamma_{43})/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.532±0.013 OUR FIT</b>				
<b>0.50 ± 0.06 OUR AVERAGE</b>				Error includes scale factor of 1.2.

0.562±0.050±0.048 264 COAN 96 CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

0.446±0.052±0.046 157 <sup>1</sup> BARATE 98E ALEP 1991–1995 LEP runs

<sup>1</sup> Not independent of BARATE 98E  $B(\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau)$  and  $B(\tau^- \rightarrow K^- K^0 \pi^0 \nu_\tau)$  values.

$\Gamma(\pi^-\bar{K}^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$		$\Gamma_{41}/\Gamma$			
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.382 <math>\pm</math> 0.013 OUR FIT</b>					
<b>0.383 <math>\pm</math> 0.014 OUR AVERAGE</b>					
0.386 $\pm$ 0.004 $\pm$ 0.014	27k	<sup>1</sup> RYU	14	BELL	$669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
0.347 $\pm$ 0.053 $\pm$ 0.037	299	<sup>2</sup> BARATE	99k	ALEP	1991–1995 LEP runs
0.294 $\pm$ 0.073 $\pm$ 0.037	142	<sup>3</sup> BARATE	98E	ALEP	1991–1995 LEP runs
0.41 $\pm$ 0.12 $\pm$ 0.03		<sup>4</sup> ACCIARRI	95F	L3	1991–1993 LEP runs
$\bullet \bullet \bullet$ We use the following data for averages but not for fits. $\bullet \bullet \bullet$					
0.417 $\pm$ 0.058 $\pm$ 0.044		<sup>5</sup> COAN	96	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
0.32 $\pm$ 0.11 $\pm$ 0.05	23	<sup>6</sup> BUSKULIC	96	ALEP	Repl. by BARATE 99K
<sup>1</sup> RYU 14 reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+\pi^-$ decays.					
<sup>2</sup> BARATE 99K measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.					
<sup>3</sup> BARATE 98E reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+\pi^-$ decays.					
<sup>4</sup> ACCIARRI 95F do not identify $\pi^-/K^-$ and assume $B(K^- K^0 \pi^0 \nu_\tau) = (0.05 \pm 0.05)\%$ .					
<sup>5</sup> Not independent of COAN 96 $B(h^- K^0 \pi^0 \nu_\tau)$ and $B(K^- K^0 \pi^0 \nu_\tau)$ measurements.					
<sup>6</sup> BUSKULIC 96 measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.					

$\Gamma(\bar{K}^0\rho^-\nu_\tau)/\Gamma_{\text{total}}$		$\Gamma_{42}/\Gamma$			
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.22 <math>\pm</math> 0.05 OUR AVERAGE</b>					
<b>0.22 <math>\pm</math> 0.05 OUR AVERAGE</b>					
0.250 $\pm$ 0.057 $\pm$ 0.044		<sup>1</sup> BARATE	99K	ALEP	1991–1995 LEP runs
0.188 $\pm$ 0.054 $\pm$ 0.038		<sup>2</sup> BARATE	98E	ALEP	1991–1995 LEP runs
<sup>1</sup> BARATE 99K measure $K^0$ 's by detecting $K_L^0$ 's in hadron calorimeter. They determine the $\bar{K}^0\rho^-$ fraction in $\tau^- \rightarrow \pi^-\bar{K}^0\pi^0\nu_\tau$ decays to be $(0.72 \pm 0.12 \pm 0.10)$ and multiply their $B(\pi^-\bar{K}^0\pi^0\nu_\tau)$ measurement by this fraction to obtain the quoted result.					
<sup>2</sup> BARATE 98E reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+\pi^-$ decays. They determine the $\bar{K}^0\rho^-$ fraction in $\tau^- \rightarrow \pi^-\bar{K}^0\pi^0\nu_\tau$ decays to be $(0.64 \pm 0.09 \pm 0.10)$ and multiply their $B(\pi^-\bar{K}^0\pi^0\nu_\tau)$ measurement by this fraction to obtain the quoted result.					

$\Gamma(K^-K^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$		$\Gamma_{43}/\Gamma$			
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>15.0 <math>\pm</math> 0.7 OUR FIT</b>					
<b>14.9 <math>\pm</math> 0.7 OUR AVERAGE</b>					
14.96 $\pm$ 0.20 $\pm$ 0.74	8.3k	<sup>1</sup> RYU	14	BELL	$669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
14.3 $\pm$ 2.5 $\pm$ 1.5	78	<sup>2</sup> BARATE	99K	ALEP	1991–1995 LEP runs
15.2 $\pm$ 7.6 $\pm$ 2.1	15	<sup>3</sup> BARATE	98E	ALEP	1991–1995 LEP runs
14.5 $\pm$ 3.6 $\pm$ 2.0	32	COAN	96	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
10 $\pm$ 5 $\pm$ 3	5	<sup>4</sup> BUSKULIC	96	ALEP	Repl. by BARATE 99K
<sup>1</sup> RYU 14 reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+\pi^-$ decays.					
<sup>2</sup> BARATE 99K measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.					
<sup>3</sup> BARATE 98E reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+\pi^-$ decays.					
<sup>4</sup> BUSKULIC 96 measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.					

$\Gamma(\pi^- \bar{K}^0 \geq 1\pi^0 \nu_\tau)/\Gamma_{\text{total}}$ 

<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{44}/\Gamma = (\Gamma_{41} + \Gamma_{45})/\Gamma$
<b><math>0.408 \pm 0.025</math> OUR FIT</b>					
<b><math>0.324 \pm 0.074 \pm 0.066</math></b>	148	ABBIENDI	00C	OPAL	1991–1995 LEP runs

 $\Gamma(\pi^- \bar{K}^0 \pi^0 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ 

<u>VALUE (units <math>10^{-3}</math>)</u>	<u>CL %</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{45}/\Gamma$
<b><math>0.26 \pm 0.23</math> OUR FIT</b>						

**$0.26 \pm 0.24$**  <sup>1</sup> BARATE 99R ALEP 1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<0.66$	95	17	<sup>2</sup> BARATE	99K	ALEP	1991–1995 LEP runs
$0.58 \pm 0.33 \pm 0.14$		5	<sup>3</sup> BARATE	98E	ALEP	1991–1995 LEP runs

<sup>1</sup> BARATE 99R combine the BARATE 98E and BARATE 99K measurements to obtain this value.

<sup>2</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

<sup>3</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

 $\Gamma(K^- K^0 \pi^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ 

<u>VALUE</u>	<u>CL %</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{46}/\Gamma$
<b><math>&lt;0.16 \times 10^{-3}</math></b>	95	<sup>1</sup> BARATE	99R	ALEP	1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<0.18 \times 10^{-3}$	95	<sup>2</sup> BARATE	99K	ALEP	1991–1995 LEP runs
$<0.39 \times 10^{-3}$	95	<sup>3</sup> BARATE	98E	ALEP	1991–1995 LEP runs

<sup>1</sup> BARATE 99R combine the BARATE 98E and BARATE 99K bounds to obtain this value.

<sup>2</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in hadron calorimeter.

<sup>3</sup> BARATE 98E reconstruct  $K^0$ 's by using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

 $\Gamma(\pi^- K^0 \bar{K}^0 \nu_\tau)/\Gamma_{\text{total}}$ 

<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{47}/\Gamma = (\Gamma_{48} + \Gamma_{49} + \Gamma_{50})/\Gamma$
<b><math>0.155 \pm 0.024</math> OUR FIT</b>					

• • • We use the following data for averages but not for fits. • • •

<b><math>0.153 \pm 0.030 \pm 0.016</math></b>	74	<sup>1</sup> BARATE	98E	ALEP	1991–1995 LEP runs
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.31 $\pm 0.12 \pm 0.04$		<sup>2</sup> ACCIARRI	95F	L3	1991–1993 LEP runs
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<sup>1</sup> BARATE 98E obtain this value by adding twice their  $B(\pi^- K_S^0 K_S^0 \nu_\tau)$  value to their  $B(\pi^- K_S^0 K_L^0 \nu_\tau)$  value.

<sup>2</sup> ACCIARRI 95F assume  $B(\pi^- K_S^0 K_S^0 \nu) = B(\pi^- K_S^0 K_L^0 \nu) = 1/2B(\pi^- K_S^0 K_L^0 \nu)$ .

 $\Gamma(\pi^- K_S^0 K_S^0 \nu_\tau)/\Gamma_{\text{total}}$ 

Bose-Einstein correlations might make the mixing fraction different than 1/4.

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{48}/\Gamma$
<b><math>2.35 \pm 0.06</math> OUR FIT</b>					

 **$2.32 \pm 0.06$  OUR AVERAGE**

2.33 $\pm 0.03 \pm 0.09$	6.7k	RYU	14	BELL	$669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
2.31 $\pm 0.04 \pm 0.08$	5.0k	<sup>1</sup> LEES	12Y	BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
2.6 $\pm 1.0 \pm 0.5$	6	BARATE	98E	ALEP	1991–1995 LEP runs
2.3 $\pm 0.5 \pm 0.3$	42	COAN	96	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

<sup>1</sup> The correlation coefficient between this measurement and the LEES 12Y  $\Gamma(\tau^- \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$  one is 0.0828.

$\Gamma(\pi^- K_S^0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{49}/\Gamma$ 

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>10.8 \pm 2.4</math> OUR FIT</b>				
<b><math>10.1 \pm 2.3 \pm 1.3</math></b>	68	BARATE	98E	ALEP 1991–1995 LEP runs

 $\Gamma(\pi^- K_L^0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{50}/\Gamma = \Gamma_{48}/\Gamma$ 

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>DOCUMENT ID</u>
<b><math>2.35 \pm 0.06</math> OUR FIT</b>	

 $\Gamma(\pi^- K^0 \bar{K}^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{51}/\Gamma = (\Gamma_{52} + \Gamma_{56} + \Gamma_{57})/\Gamma$ 

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>3.6 \pm 1.2</math> OUR FIT</b>			

• • • We use the following data for averages but not for fits. • • •

 **$3.1 \pm 2.3$** 

<sup>1</sup> BARATE 99R combine BARATE 98E  $\Gamma(\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$  and

$\Gamma(\pi^- K_S^0 K_L^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$  measurements to obtain this value.

 $\Gamma(\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{52}/\Gamma$ 

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>1.82 \pm 0.21</math> OUR FIT</b>					
<b><math>1.80 \pm 0.21</math> OUR AVERAGE</b>					

2.00  $\pm 0.22 \pm 0.20$  303 RYU 14 BELL 669 fb $^{-1}$   $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV  
 1.60  $\pm 0.20 \pm 0.22$  409 LEES 12Y BABR 468 fb $^{-1}$   $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<20 95 BARATE 98E ALEP 1991–1995 LEP runs

<sup>1</sup> The correlation coefficient between this measurement and the LEES 12Y  $\Gamma(\tau^- \rightarrow \pi^- K_S^0 K_S^0 \nu_\tau)/\Gamma_{\text{total}}$  one is 0.0828.

 $\Gamma(K^{*-} K^0 \pi^0 \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{53}/\Gamma$ 

<u>VALUE (units <math>10^{-6}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>10.8 \pm 1.4 \pm 1.5</math></b>	RYU	14	BELL 669 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

 $\Gamma(f_1(1285) \pi^- \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{54}/\Gamma$ 

<u>VALUE (units <math>10^{-6}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>6.8 \pm 1.3 \pm 0.7</math></b>	RYU	14	BELL 669 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

 $\Gamma(f_1(1420) \pi^- \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{55}/\Gamma$ 

<u>VALUE (units <math>10^{-6}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>2.4 \pm 0.5 \pm 0.6</math></b>	RYU	14	BELL 669 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV

 $\Gamma(\pi^- K_S^0 K_L^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{56}/\Gamma$ 

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>3.2 \pm 1.2</math> OUR FIT</b>				
<b><math>3.1 \pm 1.1 \pm 0.5</math></b>	11	BARATE	98E	ALEP 1991–1995 LEP runs

$\Gamma(\pi^- K_L^0 K_L^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ VALUE (units  $10^{-5}$ ) **$1.82 \pm 0.21$  OUR FIT** $\Gamma_{57}/\Gamma = \Gamma_{52}/\Gamma$ DOCUMENT ID $\Gamma(K^- K_S^0 K_S^0 \nu_\tau)/\Gamma_{\text{total}}$ VALUE **$<6.3 \times 10^{-7}$** CL%

90

 $\Gamma_{58}/\Gamma$ DOCUMENT IDTECNCOMMENTLEES 12Y BABR  $468 \text{ fb}^{-1}$   $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$  $\Gamma(K^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ VALUE **$<4.0 \times 10^{-7}$** CL%

90

 $\Gamma_{59}/\Gamma$ DOCUMENT IDTECNCOMMENTLEES 12Y BABR  $468 \text{ fb}^{-1}$   $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$  $\Gamma(K^0 h^+ h^- h^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$  $\Gamma_{60}/\Gamma$ VALUE (%)CL%DOCUMENT IDTECNCOMMENT **$<0.17$** 

95

TSCHIRHART

88

HRS

 $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

&lt;0.27

90

BELTRAMI

85

HRS

 $E_{\text{cm}}^{ee} = 29 \text{ GeV}$  $\Gamma(K^0 h^+ h^- h^- \nu_\tau)/\Gamma_{\text{total}}$  $\Gamma_{61}/\Gamma$ VALUE (units  $10^{-4}$ )EVTSDOCUMENT IDTECNCOMMENT **$2.5 \pm 2.0$  OUR FIT** **$2.3 \pm 1.9 \pm 0.7$** 

6

<sup>1</sup> BARATE

98E

ALEP

1991–1995 LEP runs

<sup>1</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays. $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$  $\Gamma_{62}/\Gamma$ 

$$\begin{aligned} \Gamma_{62}/\Gamma = & (0.34598\Gamma_{36} + 0.34598\Gamma_{38} + 0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + \\ & 0.6920\Gamma_{49} + 0.4247\Gamma_{52} + 0.6920\Gamma_{56} + 0.6534\Gamma_{61} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{97} + \Gamma_{103} + \\ & \Gamma_{106} + \Gamma_{107} + 0.2804\Gamma_{150} + 0.2804\Gamma_{152} + 0.2804\Gamma_{154} + 0.2628\Gamma_{156} + 0.7259\Gamma_{170} + \\ & 0.9078\Gamma_{178} + 0.9078\Gamma_{179} + 0.9078\Gamma_{180})/\Gamma \end{aligned}$$

VALUE (%)EVTSDOCUMENT IDTECNCOMMENT **$15.20 \pm 0.06$  OUR FIT** **$14.8 \pm 0.4$  OUR AVERAGE**

14.4 ± 0.6 ± 0.3

ADEVA

91F

L3

 $E_{\text{cm}}^{ee} = 88.3\text{--}94.3 \text{ GeV}$ 

15.0 ± 0.4 ± 0.3

BEHREND

89B

CELL

 $E_{\text{cm}}^{ee} = 14\text{--}47 \text{ GeV}$ 

15.1 ± 0.8 ± 0.6

AIHARA

87B

TPC

 $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

13.5 ± 0.3 ± 0.3

ABACHI

89B

HRS

 $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ 

12.8 ± 1.0 ± 0.7

<sup>1</sup> BURCHAT

87

MRK2

 $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ 

12.1 ± 0.5 ± 1.2

RUCKSTUHL

86

DLCO

 $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ 

12.8 ± 0.5 ± 0.8

SCHMIDKE

86

MRK2

 $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ 15.3 ± 1.1 <sup>+1.3</sup> <sub>-1.6</sub>

ALTHOFF

85

TASS

 $E_{\text{cm}}^{ee} = 34.5 \text{ GeV}$ 

13.6 ± 0.5 ± 0.8

BARTEL

85F

JADE

 $E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$ 

12.2 ± 1.3 ± 3.9

<sup>2</sup> BERGER

85

PLUT

 $E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$ 

13.3 ± 0.3 ± 0.6

FERNANDEZ

85

MAC

 $E_{\text{cm}}^{ee} = 29 \text{ GeV}$ 

24 ± 6

BRANDELIK

80

TASS

 $E_{\text{cm}}^{ee} = 30 \text{ GeV}$ 

32 ± 5

<sup>3</sup> BACINO

78B

DLCO

 $E_{\text{cm}}^{ee} = 3.1\text{--}7.4 \text{ GeV}$

35	$\pm 11$	<sup>3</sup>	BRANDELIK	78	DASP	Assumes $V-A$ decay
18	$\pm 6.5$	33	<sup>3</sup> JAROS	78	LGW	$E_{\text{cm}}^{\text{ee}} > 6 \text{ GeV}$

<sup>1</sup> BURCHAT 87 value is not independent of SCHMIDKE 86 value.

<sup>2</sup> Not independent of BERGER 85  $\Gamma(\mu^-\bar{\nu}_\mu\nu_\tau)/\Gamma_{\text{total}}$ ,  $\Gamma(e^-\bar{\nu}_e\nu_\tau)/\Gamma_{\text{total}}$ ,  $\Gamma(h^- \geq 1 \text{ neutrals} \nu_\tau)/\Gamma_{\text{total}}$ , and  $\Gamma(h^- \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$ , and therefore not used in the fit.

<sup>3</sup> Low energy experiments are not in average or fit because the systematic errors in background subtraction are judged to be large.

$$\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau (\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-) (\text{"3-prong"}) ) / \Gamma_{\text{total}} \quad \Gamma_{63}/\Gamma$$

$$\Gamma_{63}/\Gamma = (\Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2804\Gamma_{150} + 0.2804\Gamma_{152} + 0.2804\Gamma_{154} + 0.491\Gamma_{170} + 0.9078\Gamma_{178} + 0.9078\Gamma_{179} + 0.9078\Gamma_{180}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**14.55  $\pm 0.06$  OUR FIT**

**14.61  $\pm 0.06$  OUR AVERAGE**

$14.556 \pm 0.105 \pm 0.076$	<sup>1</sup> ACHARD	01D	L3	1992–1995 LEP runs
$14.96 \pm 0.09 \pm 0.22$	10.4k	AKERS	95Y	OPAL 1991–1994 LEP runs

• • • We use the following data for averages but not for fits. • • •

$14.652 \pm 0.067 \pm 0.086$	SCHAEL	05C	ALEP	1991–1995 LEP runs
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$14.569 \pm 0.093 \pm 0.048$	23k	<sup>2</sup> ABREU	01M	DLEP 1992–1995 LEP runs
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$14.22 \pm 0.10 \pm 0.37$	<sup>3</sup> BALEST	95C	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$15.26 \pm 0.26 \pm 0.22$	ACTON	92H	OPAL	Repl. by AKERS 95Y
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$13.3 \pm 0.3 \pm 0.8$	<sup>4</sup> ALBRECHT	92D	ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
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$14.35 \begin{array}{l} +0.40 \\ -0.45 \end{array} \pm 0.24$	DECAMP	92C	ALEP	1989–1990 LEP runs
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<sup>1</sup> The correlation coefficients between this measurement and the ACHARD 01D measurements of  $B(\tau \rightarrow \text{"1-prong"})$  and  $B(\tau \rightarrow \text{"5-prong"})$  are  $-0.978$  and  $-0.19$  respectively.

<sup>2</sup> The correlation coefficients between this measurement and the ABREU 01M measurements of  $B(\tau \rightarrow \text{1-prong})$  and  $B(\tau \rightarrow \text{5-prong})$  are  $-0.98$  and  $-0.08$  respectively.

<sup>3</sup> Not independent of BALEST 95C  $B(h^- h^- h^+ \nu_\tau)$  and  $B(h^- h^- h^+ \pi^0 \nu_\tau)$  values, and BORTOLETTO 93  $B(h^- h^- h^+ 2\pi^0 \nu_\tau) / B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau)$  value.

<sup>4</sup> This ALBRECHT 92D value is not independent of their  $\Gamma(\mu^-\bar{\nu}_\mu\nu_\tau)\Gamma(e^-\bar{\nu}_e\nu_\tau)/\Gamma_{\text{total}}^2$  value.

$$\Gamma(h^- h^- h^+ \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{64}/\Gamma$$

$$\Gamma_{64}/\Gamma = (0.34598\Gamma_{36} + 0.34598\Gamma_{38} + \Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.491\Gamma_{170} + 0.0153\Gamma_{178} + 0.0153\Gamma_{179}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**9.80  $\pm 0.05$  OUR FIT**

• • • We use the following data for averages but not for fits. • • •

<b>7.6 <math>\pm 0.1 \pm 0.5</math></b>	7.5k	<sup>1</sup> ALBRECHT	96E	ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$9.92 \pm 0.10 \pm 0.09$	11.2k	<sup>2</sup> BUSKULIC	96	ALEP Repl. by SCHAEL 05C
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$9.49 \pm 0.36 \pm 0.63$		DECAMP	92C	ALEP Repl. by SCHAEL 05C
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$8.7 \pm 0.7 \pm 0.3$	694	<sup>3</sup> BEHREND	90	CELL $E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$
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$7.0 \pm 0.3 \pm 0.7$	1566	<sup>4</sup> BAND	87	MAC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
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$6.7 \pm 0.8 \pm 0.9$		<sup>5</sup> BURCHAT	87	MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
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$6.4 \pm 0.4 \pm 0.9$		<sup>6</sup> RUCKSTUHL	86	DLCO $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
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$7.8 \pm 0.5 \pm 0.8$	890	SCHMIDKE	86	MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
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$8.4 \pm 0.4 \pm 0.7$	1255	<sup>6</sup> FERNANDEZ	85	MAC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
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$9.7 \pm 2.0 \pm 1.3$		BEHREND	84	CELL $E_{\text{cm}}^{\text{ee}} = 14,22 \text{ GeV}$
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- <sup>1</sup> ALBRECHT 96E not independent of ALBRECHT 93c  $\Gamma(h^- h^- h^+ \nu_\tau \text{ (ex. } K^0) \times \Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)) / \Gamma_{\text{total}}^2$  value.
- <sup>2</sup> BUSKULIC 96 quote  $B(h^- h^- h^+ \nu_\tau \text{ (ex. } K^0)) = 9.50 \pm 0.10 \pm 0.11$ . We add 0.42 to remove their  $K^0$  correction and reduce the systematic error accordingly.
- <sup>3</sup> BEHREND 90 subtract 0.3% to account for the  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  contribution to measured events.
- <sup>4</sup> BAND 87 subtract for charged kaon modes; not independent of FERNANDEZ 85 value.
- <sup>5</sup> BURCHAT 87 value is not independent of SCHMIDKE 86 value.
- <sup>6</sup> Value obtained by multiplying paper's  $R = B(h^- h^- h^+ \nu_\tau) / B(3\text{-prong})$  by  $B(3\text{-prong}) = 0.143$  and subtracting 0.3% for  $K^*(892)$  background.

$$\Gamma(h^- h^- h^+ \nu_\tau \text{ (ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{65} / \Gamma$$

$$\Gamma_{65} / \Gamma = (\Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.491 \Gamma_{170} + 0.0153 \Gamma_{178} + 0.0153 \Gamma_{179}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.46 ± 0.05 OUR FIT</b>				
<b>9.44 ± 0.14 OUR AVERAGE</b>				Error includes scale factor of 1.4. See the ideogram below.

9.317 ± 0.090 ± 0.082      12.2k      <sup>1</sup> ABDALLAH      06A      DLPH      1992–1995 LEP runs  
 9.51 ± 0.07 ± 0.20      37.7k      BALEST      95C      CLEO       $E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

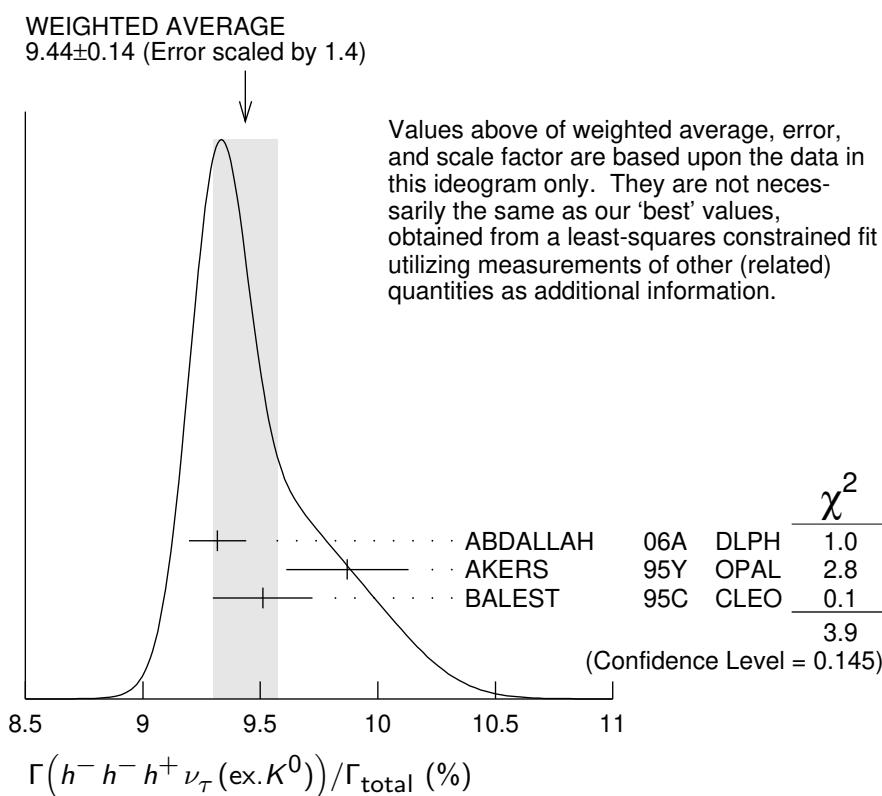
9.87 ± 0.10 ± 0.24      <sup>2</sup> AKERS      95Y      OPAL      1991–1994 LEP runs  
 • • • We do not use the following data for averages, fits, limits, etc. • • •

9.50 ± 0.10 ± 0.11      11.2k      <sup>3</sup> BUSKULIC      96      ALEP      Repl. by SCHAEEL 05C

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> Not independent of AKERS 95Y  $B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  and  $B(h^- h^- h^+ \nu_\tau \text{ (ex. } K^0)) / B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  values.

<sup>3</sup> Not independent of BUSKULIC 96  $B(h^- h^- h^+ \nu_\tau)$  value.



$$\Gamma(h^- h^- h^+ \nu_\tau(\text{ex. } K^0)) / \Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau(\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-))$$

("3-prong")  $\Gamma_{65}/\Gamma_{63}$

$$\Gamma_{65}/\Gamma_{63} = (\Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.491\Gamma_{170} + 0.0153\Gamma_{178} + 0.0153\Gamma_{179}) / (0.4247\Gamma_{52} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{89} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2804\Gamma_{150} + 0.2302\Gamma_{151} + 0.2804\Gamma_{152} + 0.2804\Gamma_{154} + 0.1131\Gamma_{156} + 0.3257\Gamma_{160} + 0.491\Gamma_{170} + 0.9078\Gamma_{178} + 0.9078\Gamma_{179} + 0.9078\Gamma_{180} + 0.892\Gamma_{182})$$

<u>VALUE (units <math>10^{-2}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>64.98 \pm 0.31</math> OUR FIT</b>			
<b><math>66.0 \pm 0.4 \pm 1.4</math></b>	AKERS	95Y OPAL	1991–1994 LEP runs

$$\Gamma(h^- h^- h^+ \nu_\tau(\text{ex. } K^0, \omega)) / \Gamma_{\text{total}}$$

$\Gamma_{66}/\Gamma = (\Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.491\Gamma_{170}) / \Gamma$

<u>VALUE (%)</u>	<u>DOCUMENT ID</u>
<b><math>9.43 \pm 0.05</math> OUR FIT</b>	

$$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$$

$\Gamma_{67}/\Gamma = (0.34598\Gamma_{36} + \Gamma_{70} + 0.0153\Gamma_{178}) / \Gamma$

VALUE (%) DOCUMENT ID

<u>9.31 <math>\pm 0.05</math> OUR FIT</u>	<u>DOCUMENT ID</u>

$$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}}$$

$\Gamma_{68}/\Gamma = (\Gamma_{70} + 0.0153\Gamma_{178}) / \Gamma$

VALUE (%) EVTS DOCUMENT ID TECN COMMENT

<u>9.02 <math>\pm 0.05</math> OUR FIT</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>8.77 \pm 0.13</math> OUR AVERAGE</b>				Error includes scale factor of 1.1.

<sup>1</sup> Quoted statistical error is 0.003%. Correlation matrix for LEE 10 branching fractions:

- (1)  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (2)  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (3)  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$
- (4)  $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau) / \Gamma_{\text{total}}$

(1)	(2)	(3)	
(2)	0.175		
(3)	0.049	0.080	
(4)	-0.053	0.035	-0.008

<sup>2</sup> Correlation matrix for AUBERT 08 branching fractions:

- (1)  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (2)  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}}$
- (3)  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$
- (4)  $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau) / \Gamma_{\text{total}}$

(1)	(2)	(3)	
(2)	0.544		
(3)	0.390	0.177	
(4)	0.031	0.093	0.087

<sup>3</sup> 47% correlated with BRIERE 03  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  and 71% correlated with  $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$  because of a common 5% normalization error.

$$\Gamma(\pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0), \text{non-axial vector})/\Gamma(\pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0)) \quad \Gamma_{69}/\Gamma_{68}$$

$$\Gamma_{69}/\Gamma_{68} = \Gamma_{69}/(\Gamma_{70} + 0.0153\Gamma_{177})$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.261	95	<sup>1</sup> ACKERSTAFF 97R	OPAL	1992–1994 LEP runs

<sup>1</sup> Model-independent limit from structure function analysis on contribution to  $B(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0))$  from non-axial vectors.

$$\Gamma(\pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0, \omega))/\Gamma_{\text{total}} \quad \Gamma_{70}/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.99 ±0.05 OUR FIT</b>				
<b>9.041±0.060±0.076</b>	29k	<sup>1</sup> SCHAEL	05C ALEP	1991–1995 LEP runs

<sup>1</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

$$\Gamma(h^- h^- h^+ \geq 1 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{71}/\Gamma$$

$$\begin{aligned} \Gamma_{71}/\Gamma = & (0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.4247\Gamma_{52} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \\ & \Gamma_{103} + \Gamma_{107} + 0.2804\Gamma_{150} + 0.2804\Gamma_{152} + 0.2804\Gamma_{154} + 0.2926\Gamma_{156} + 0.892\Gamma_{178} + \\ & 0.892\Gamma_{179} + 0.9078\Gamma_{180})/\Gamma \end{aligned}$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.29±0.05 OUR FIT</b>				

• • • We do not use the following data for averages, fits, limits, etc. • • •

5.6 ± 0.7 ± 0.3	352	<sup>1</sup> BEHREND	90	CELL $E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$
4.2 ± 0.5 ± 0.9	203	<sup>2</sup> ALBRECHT	87L	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
6.1 ± 0.8 ± 0.9		<sup>3</sup> BURCHAT	87	MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
7.6 ± 0.4 ± 0.9		<sup>4,5</sup> RUCKSTUHL	86	DLCO $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
4.7 ± 0.5 ± 0.8	530	<sup>6</sup> SCHMIDKE	86	MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
5.6 ± 0.4 ± 0.7		<sup>5</sup> FERNANDEZ	85	MAC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
6.2 ± 2.3 ± 1.7		BEHREND	84	CELL $E_{\text{cm}}^{\text{ee}} = 14,22 \text{ GeV}$

<sup>1</sup> BEHREND 90 value is not independent of BEHREND 90  $B(3h\nu_\tau \geq 1 \text{ neutrals}) + B(5\text{-prong})$ .

<sup>2</sup> ALBRECHT 87L measure the product of branching ratios  $B(3\pi^\pm \pi^0 \nu_\tau) B((e\bar{\nu} \text{ or } \mu\bar{\nu} \text{ or } \pi \text{ or } K \text{ or } \rho)\nu_\tau) = 0.029$  and use the PDG 86 values for the second branching ratio which sum to  $0.69 \pm 0.03$  to get the quoted value.

<sup>3</sup> BURCHAT 87 value is not independent of SCHMIDKE 86 value.

<sup>4</sup> Contributions from kaons and from  $>1\pi^0$  are subtracted. Not independent of (3-prong +  $0\pi^0$ ) and (3-prong +  $\geq 0\pi^0$ ) values.

<sup>5</sup> Value obtained using paper's  $R = B(h^- h^- h^+ \nu_\tau)/B(3\text{-prong})$  and current  $B(3\text{-prong}) = 0.143$ .

<sup>6</sup> Not independent of SCHMIDKE 86  $h^- h^- h^+ \nu_\tau$  and  $h^- h^- h^+(\geq 0\pi^0)\nu_\tau$  values.

$$\Gamma(h^- h^- h^+ \geq 1\pi^0 \nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{72}/\Gamma$$

$$\begin{aligned} \Gamma_{72}/\Gamma = & (\Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{150} + 0.2302\Gamma_{152} + 0.2302\Gamma_{154} + \\ & 0.892\Gamma_{178} + 0.892\Gamma_{179} + 0.9078\Gamma_{180})/\Gamma \end{aligned}$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.09 ±0.05 OUR FIT</b>				
<b>5.10 ±0.12 OUR AVERAGE</b>				

• • • We use the following data for averages but not for fits. • • •

5.106±0.083±0.103	10.1k	<sup>1</sup> ABDALLAH	06A DLPH	1992–1995 LEP runs
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$5.09 \pm 0.10 \pm 0.23$       <sup>2</sup> AKERS      95Y OPAL 1991–1994 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

$4.95 \pm 0.29 \pm 0.65$       570      DECOMP      92C ALEP Repl. by SCHAEEL 05C

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> Not independent of AKERS 95Y  $B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  and  $B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau \text{ (ex. } K^0)/B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  values.

### $\Gamma(h^- h^- h^+ \pi^0 \nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_{73}/\Gamma$

$$\Gamma_{73}/\Gamma = (0.34598\Gamma_{41} + 0.34598\Gamma_{43} + \Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{178} + 0.892\Gamma_{179} + 0.0153\Gamma_{180})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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#### **4.76 ± 0.05 OUR FIT**

• • • We do not use the following data for averages, fits, limits, etc. • • •

$4.45 \pm 0.09 \pm 0.07$       6.1k      <sup>1</sup> BUSKULIC      96      ALEP Repl. by SCHAEEL 05C

<sup>1</sup> BUSKULIC 96 quote  $B(h^- h^- h^+ \pi^0 \nu_\tau \text{ (ex. } K^0)) = 4.30 \pm 0.09 \pm 0.09$ . We add 0.15 to remove their  $K^0$  correction and reduce the systematic error accordingly.

### $\Gamma(h^- h^- h^+ \pi^0 \nu_\tau \text{ (ex. } K^0))/\Gamma_{\text{total}}$

$\Gamma_{74}/\Gamma$

$$\Gamma_{74}/\Gamma = (\Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{178} + 0.892\Gamma_{179} + 0.0153\Gamma_{180})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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#### **4.57 ± 0.05 OUR FIT**

**4.45 ± 0.14 OUR AVERAGE** Error includes scale factor of 1.2.

$4.545 \pm 0.106 \pm 0.103$       8.9k      <sup>1</sup> ABDALLAH      06A DLPH 1992–1995 LEP runs

$4.23 \pm 0.06 \pm 0.22$       7.2k      BAILEST      95C CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

### $\Gamma(h^- h^- h^+ \pi^0 \nu_\tau \text{ (ex. } K^0, \omega))/\Gamma_{\text{total}}$

$$\Gamma_{75}/\Gamma = (\Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152})/\Gamma$$

VALUE (%)	DOCUMENT ID
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#### **2.79 ± 0.07 OUR FIT**

### $\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_{76}/\Gamma$

$$\Gamma_{76}/\Gamma = (0.34598\Gamma_{41} + \Gamma_{78} + 0.892\Gamma_{178} + 0.0153\Gamma_{180})/\Gamma$$

VALUE (%)	DOCUMENT ID
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#### **4.62 ± 0.05 OUR FIT**

### $\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau \text{ (ex. } K^0))/\Gamma_{\text{total}}$

$\Gamma_{77}/\Gamma$

$$\Gamma_{77}/\Gamma = (\Gamma_{78} + 0.892\Gamma_{178} + 0.0153\Gamma_{180})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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#### **4.49 ± 0.05 OUR FIT**

**4.55 ± 0.13 OUR AVERAGE** Error includes scale factor of 1.6.

$4.598 \pm 0.057 \pm 0.064$  16k      <sup>1</sup> SCHAEEL      05C ALEP 1991–1995 LEP runs

$4.19 \pm 0.10 \pm 0.21$       <sup>2</sup> EDWARDS      00A CLEO  $4.7 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup>SCHAEL 05C quote  $(4.590 \pm 0.057 \pm 0.064)\%$ . We add 0.008% to remove their correction for  $\tau^- \rightarrow \pi^- \pi^0 \omega \nu_\tau \rightarrow \pi^- \pi^0 \pi^+ \pi^- \nu_\tau$  decays. See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.  
<sup>2</sup>EDWARDS 00A quote  $(4.19 \pm 0.10) \times 10^{-2}$  with a 5% systematic error.

$\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0, \omega))/\Gamma_{\text{total}}$	$\Gamma_{78}/\Gamma$
VALUE (%)	DOCUMENT ID
<b><math>2.74 \pm 0.07</math> OUR FIT</b>	

$\Gamma(h^- \rho \pi^0 \nu_\tau)/\Gamma(h^- h^- h^+ \pi^0 \nu_\tau)$	$\Gamma_{79}/\Gamma_{73}$			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$0.30 \pm 0.04 \pm 0.02$	393	ALBRECHT	91D ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$

$\Gamma(h^- \rho^+ h^- \nu_\tau)/\Gamma(h^- h^- h^+ \pi^0 \nu_\tau)$	$\Gamma_{80}/\Gamma_{73}$			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$0.10 \pm 0.03 \pm 0.04$	142	ALBRECHT	91D ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$

$\Gamma(h^- \rho^- h^+ \nu_\tau)/\Gamma(h^- h^- h^+ \pi^0 \nu_\tau)$	$\Gamma_{81}/\Gamma_{73}$			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$0.26 \pm 0.05 \pm 0.01$	370	ALBRECHT	91D ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$

$\Gamma(h^- h^- h^+ \geq 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$	$\Gamma_{82}/\Gamma$			
$\Gamma_{82}/\Gamma = (\Gamma_{85} + \Gamma_{86} + 0.2302\Gamma_{150} + 0.2302\Gamma_{154} + 0.892\Gamma_{180})/\Gamma$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.517 \pm 0.031</math> OUR FIT</b>				
<b><math>0.561 \pm 0.068 \pm 0.095</math></b>	1.3k	<sup>1</sup> ABDALLAH	06A DLPH	1992–1995 LEP runs

<sup>1</sup>See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{83}/\Gamma$
$\Gamma_{83}/\Gamma = (0.4247\Gamma_{48} + \Gamma_{85} + 0.2302\Gamma_{150} + 0.2302\Gamma_{154} + 0.892\Gamma_{180})/\Gamma$	
VALUE (%)	DOCUMENT ID
<b><math>0.505 \pm 0.031</math> OUR FIT</b>	

$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$	$\Gamma_{84}/\Gamma$			
$\Gamma_{84}/\Gamma = (\Gamma_{85} + 0.2302\Gamma_{150} + 0.2302\Gamma_{154} + 0.892\Gamma_{180})/\Gamma$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.495 \pm 0.031</math> OUR FIT</b>				
<b><math>0.435 \pm 0.030 \pm 0.035</math></b>	2.6k	<sup>1</sup> SCHAEL	05C ALEP	1991–1995 LEP runs

$\bullet \bullet \bullet$  We do not use the following data for averages, fits, limits, etc.  $\bullet \bullet \bullet$

$0.50 \pm 0.07 \pm 0.07$       1.8k      BUSKULIC      96      ALEP      Repl. by SCHAEL 05C

<sup>1</sup>SCHAEL 05C quote  $(0.392 \pm 0.030 \pm 0.035)\%$ . We add 0.043% to remove their correction for  $\tau^- \rightarrow \pi^- \eta \pi^0 \nu_\tau \rightarrow \pi^- \pi^+ \pi^- 2\pi^0 \nu_\tau$  and  $\tau^- \rightarrow K^*(892)^- \eta \nu_\tau \rightarrow K^- \pi^+ \pi^- 2\pi^0 \nu_\tau$  decays. See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

$$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) \quad \Gamma_{84}/\Gamma_{62}$$

$$\begin{aligned} \Gamma_{84}/\Gamma_{62} = & (\Gamma_{85} + 0.2302\Gamma_{150} + 0.2302\Gamma_{154} + 0.892\Gamma_{180}) / (0.34598\Gamma_{36} + \\ & 0.34598\Gamma_{38} + 0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.6920\Gamma_{49} + 0.8494\Gamma_{52} + \\ & 0.6920\Gamma_{56} + 0.6534\Gamma_{61} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{89} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + \\ & 0.2804\Gamma_{150} + 0.2302\Gamma_{151} + 0.2804\Gamma_{152} + 0.2804\Gamma_{154} + 0.3759\Gamma_{156} + 0.3257\Gamma_{160} + \\ & 0.7259\Gamma_{170} + 0.9078\Gamma_{178} + 0.9078\Gamma_{179} + 0.9078\Gamma_{180} + 0.892\Gamma_{182}) \end{aligned}$$

<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>3.26 \pm 0.20</math> OUR FIT</b>				
<b><math>3.4 \pm 0.2 \pm 0.3</math></b>	668	BORTOLETTO93	CLEO	$E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$

$$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0, \omega, \eta)) / \Gamma_{\text{total}} \quad \Gamma_{85}/\Gamma$$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>DOCUMENT ID</u>
<b><math>10 \pm 4</math> OUR FIT</b>	

$$\Gamma(h^- h^- h^+ 3\pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{86}/\Gamma = (0.4247\Gamma_{52} + \Gamma_{87} + 0.1131\Gamma_{156}) / \Gamma$$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>2.13 \pm 0.30</math> OUR FIT</b>					
<b><math>2.2 \pm 0.3 \pm 0.4</math></b>	139	ANASTASSOV 01	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 4.9	95	SCHAEL	05C	ALEP	1991-1995 LEP runs
$2.85 \pm 0.56 \pm 0.51$	57	ANDERSON	97	CLEO	Repl. by ANAS-TASSOV 01
11 $\pm 4 \pm 5$	440	<sup>1</sup> BUSKULIC	96	ALEP	Repl. by SCHAEL 05C

<sup>1</sup> BUSKULIC 96 state their measurement is for  $B(h^- h^- h^+ \geq 3\pi^0 \nu_\tau)$ . We assume that  $B(h^- h^- h^+ \geq 4\pi^0 \nu_\tau)$  is very small.

$$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{87}/\Gamma$$

$$\Gamma_{87}/\Gamma = (\Gamma_{89} + 0.2302\Gamma_{151} + 0.3257\Gamma_{160} + 0.892\Gamma_{182}) / \Gamma$$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>1.95 \pm 0.30</math> OUR FIT</b>			

• • • We use the following data for averages but not for fits. • • •

<b><math>2.07 \pm 0.18 \pm 0.37</math></b>	<sup>1</sup> LEES	12X	BABR	$468 \text{ fb}^{-1}$	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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<sup>1</sup> Not independent of LEES 12X  $\Gamma(\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma$ ,  $\Gamma(\tau^- \rightarrow \eta \pi^- \pi^0 \nu_\tau) / \Gamma$ ,  $\Gamma(\tau^- \rightarrow \pi^- \omega 2\pi^0 \nu_\tau) / \Gamma$ , and  $\Gamma(\tau^- \rightarrow f_1(1285) \pi^- \nu_\tau \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau) / \Gamma$  values.

$$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta, f_1(1285))) / \Gamma_{\text{total}} \quad \Gamma_{88}/\Gamma$$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>1.69 \pm 0.08 \pm 0.43</math></b>	LEES	12X	BABR

$$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285))) / \Gamma_{\text{total}} \quad \Gamma_{89}/\Gamma$$

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>1.4 \pm 2.7</math> OUR FIT</b>			

<b><math>1.0 \pm 0.8 \pm 3.0</math></b>	<sup>1</sup> LEES	12X	BABR	$468 \text{ fb}^{-1}$	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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<sup>1</sup> LEES 12X measurement corresponds to the lower limit of  $< 5.8 \times 10^{-5}$  at 90% CL.

$$\Gamma(K^- h^+ h^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{90}/\Gamma$$

$$\Gamma_{90}/\Gamma = (0.34598\Gamma_{38} + 0.34598\Gamma_{43} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2804\Gamma_{152} + 0.491\Gamma_{170} + 0.9078\Gamma_{179})/\Gamma$$

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.629±0.014 OUR FIT</b>				
<0.6	90	AIHARA	84C TPC	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$

$$\Gamma(K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{91}/\Gamma = (\Gamma_{97} + \Gamma_{106} + 0.0153\Gamma_{179})/\Gamma$$

VALUE (%)	DOCUMENT ID
<b>0.437±0.007 OUR FIT</b>	

$$\Gamma(K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{91}/\Gamma_{68}$$

$$\Gamma_{91}/\Gamma_{68} = (\Gamma_{97} + \Gamma_{106} + 0.0153\Gamma_{179})/(\Gamma_{70} + 0.0153\Gamma_{178})$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.85±0.08 OUR FIT</b>				
<b>5.44±0.21±0.53</b>	7.9k	RICHICHI	99	CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$$\Gamma(K^- h^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{92}/\Gamma$$

$$\Gamma_{92}/\Gamma = (\Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{179})/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID
<b>8.6±1.2 OUR FIT</b>	

$$\Gamma(K^- h^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{92}/\Gamma_{77}$$

$$\Gamma_{92}/\Gamma_{77} = (\Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{179})/(\Gamma_{78} + 0.892\Gamma_{178} + 0.0153\Gamma_{180})$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.91±0.26 OUR FIT</b>				
<b>2.61±0.45±0.42</b>	719	RICHICHI	99	CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$$\Gamma(K^- \pi^+ \pi^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{93}/\Gamma$$

$$\Gamma_{93}/\Gamma = (0.34598\Gamma_{38} + 0.34598\Gamma_{43} + \Gamma_{97} + \Gamma_{103} + 0.2804\Gamma_{152} + 0.9078\Gamma_{179})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.477±0.014 OUR FIT</b>				
<b>0.58 ±0.15 ±0.12</b>	20	<sup>1</sup> BAUER	94	TPC $E_{\text{cm}}^{ee} = 29 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.22	+0.16 -0.13	±0.05	9	<sup>2</sup> MILLS	85	DLCO $E_{\text{cm}}^{ee} = 29 \text{ GeV}$
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<sup>1</sup> We multiply 0.58% by 0.20, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

<sup>2</sup> Error correlated with MILLS 85 ( $KK\pi\nu$ ) value. We multiply 0.22% by 0.23, the relative systematic error quoted by MILLS 85, to obtain the systematic error.

$$\Gamma(K^- \pi^+ \pi^- \geq 0 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{94}/\Gamma$$

$$\Gamma_{94}/\Gamma = (\Gamma_{97} + \Gamma_{103} + 0.2302\Gamma_{152} + 0.9078\Gamma_{179})/\Gamma$$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>0.373±0.013 OUR FIT</b>			
<b>0.30 ±0.05 OUR AVERAGE</b>			

• • • We use the following data for averages but not for fits. • • •

0.343±0.073±0.031	ABBIENDI	00D	OPAL	1990–1995 LEP runs
0.275±0.064	<sup>1</sup> BARATE	98	ALEP	1991–1995 LEP runs

<sup>1</sup> Not independent of BARATE 98  $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau)/\Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$  values.

$$\Gamma(K^-\pi^+\pi^-\nu_\tau)/\Gamma_{\text{total}}$$

VALUE (%)

**0.345±0.007 OUR FIT**

$$\Gamma_{95}/\Gamma = (0.34598\Gamma_{38} + \Gamma_{97} + 0.0153\Gamma_{179})/\Gamma$$

DOCUMENT ID

$$\Gamma(K^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$$

VALUE (%)

**0.293±0.007 OUR FIT**

$$0.290\pm 0.018 \text{ OUR AVERAGE}$$

EVTS

DOCUMENT ID

$$\Gamma_{96}/\Gamma = (\Gamma_{97} + 0.0153\Gamma_{179})/\Gamma$$

TECN

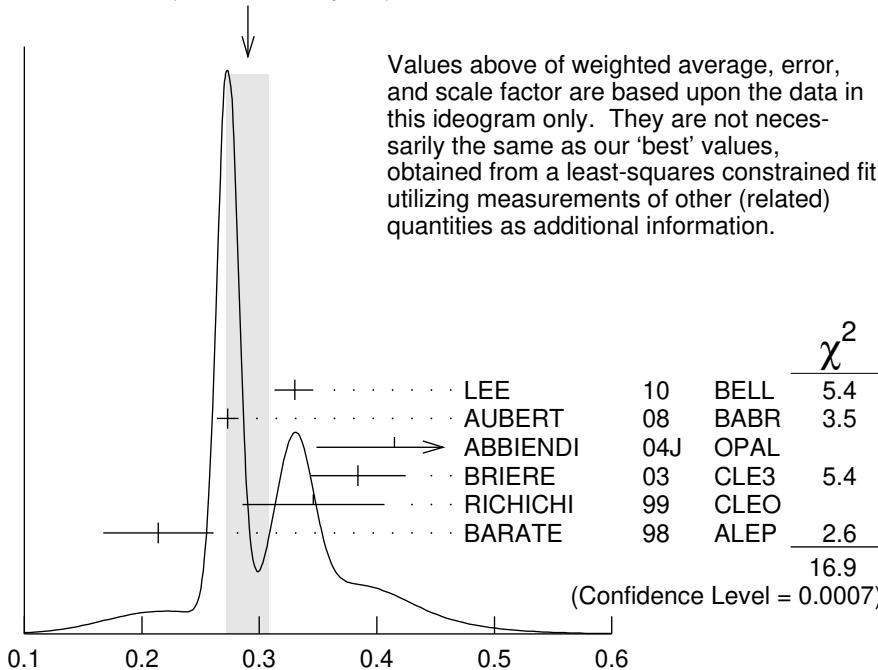
COMMENT

Error includes scale factor of 2.4. See the ideogram below.

$0.330\pm 0.001^{+0.016}_{-0.017}$	794k	<sup>1</sup> LEE	10	BELL	$666 \text{ fb}^{-1}$	$E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$
$0.273\pm 0.002\pm 0.009$	70k	<sup>2</sup> AUBERT	08	BABR	$342 \text{ fb}^{-1}$	$E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$
$0.415\pm 0.053\pm 0.040$	269	ABBIENDI	04J	OPAL	1991–1995 LEP runs	
$0.384\pm 0.014\pm 0.038$	3.5k	<sup>3</sup> BRIERE	03	CLE3	$E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$	
$0.214\pm 0.037\pm 0.029$		BARATE	98	ALEP	1991–1995 LEP runs	
<b>• • • We use the following data for averages but not for fits. • • •</b>						
$0.346\pm 0.023\pm 0.056$	158	<sup>4</sup> RICHICHI	99	CLEO	$E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$	
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>						
$0.360\pm 0.082\pm 0.048$		ABBIENDI	00D	OPAL	1990–1995 LEP runs	

## WEIGHTED AVERAGE

0.290±0.018 (Error scaled by 2.4)



$$\Gamma(K^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}} (\%)$$

<sup>1</sup> See footnote to LEE 10  $\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$  measurement for correlations with other measurements. Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0))/\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex. } K^0))$  value.

<sup>2</sup> See footnote to AUBERT 08  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>3</sup> 47% correlated with BRIERE 03  $\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau$  and 34% correlated with  $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$  because of a common 5% normalization error.

<sup>4</sup> Not independent of RICHICHI 99

$\Gamma(\tau^- \rightarrow K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$ ,  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  and BALEST 95c  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  values.

$$\Gamma(K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{96}/\Gamma_{68}$$

$$\Gamma_{96}/\Gamma_{68} = (\Gamma_{97} + 0.0153\Gamma_{179}) / (\Gamma_{70} + 0.0153\Gamma_{178})$$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.25 \pm 0.07</math> OUR FIT</b>				

• • • We use the following data for averages but not for fits. • • •

$$3.92 \pm 0.02^{+0.15}_{-0.16} \quad 794k \quad ^1 \text{LEE} \quad 10 \quad \text{BELL} \quad 666 \text{ fb}^{-1} \quad E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$$

<sup>1</sup> Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  values.

$$\Gamma(K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0, \omega)) / \Gamma_{\text{total}} \quad \Gamma_{97}/\Gamma$$

VALUE (units $10^{-3}$ )	DOCUMENT ID
<b><math>2.93 \pm 0.07</math> OUR FIT</b>	

$$\Gamma(K^- \rho^0 \nu_\tau \rightarrow K^- \pi^+ \pi^- \nu_\tau) / \Gamma(K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{98}/\Gamma_{96}$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.48 \pm 0.14 \pm 0.10</math></b>	<sup>1</sup> ASNER	00B CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$$0.39 \pm 0.14 \quad ^2 \text{BARATE} \quad 99R \quad \text{ALEP} \quad 1991\text{--}1995 \text{ LEP runs}$$

<sup>1</sup> ASNER 00B assume  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)$  decays proceed only through  $K\rho$  and  $K^*\pi$  intermediate states. They assume the resonance structure of  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)$  decays is dominated by  $K_1(1270)^-$  and  $K_1(1400)^-$  resonances, and assume  $B(K_1(1270) \rightarrow K^*(892)\pi) = (16 \pm 5)\%$ ,  $B(K_1(1270) \rightarrow K\rho) = (42 \pm 6)\%$ , and  $B(K_1(1400) \rightarrow K\rho) = 0$ .

<sup>2</sup> BARATE 99R assume  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)$  decays proceed only through  $K\rho$  and  $K^*\pi$  intermediate states. The quoted error is statistical only.

$$\Gamma(K^- \pi^+ \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{99}/\Gamma$$

$$\Gamma_{99}/\Gamma = (0.34598\Gamma_{43} + \Gamma_{103} + 0.2302\Gamma_{152} + 0.892\Gamma_{179}) / \Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID
<b><math>13.1 \pm 1.2</math> OUR FIT</b>	

$$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}} \quad \Gamma_{100}/\Gamma$$

$$\Gamma_{100}/\Gamma = (\Gamma_{103} + 0.2302\Gamma_{152} + 0.892\Gamma_{179})/\Gamma$$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**7.9±1.2 OUR FIT****7.3±1.2 OUR AVERAGE**

$7.4 \pm 0.8 \pm 1.1$  <sup>1</sup> ARMS 05 CLE3  $7.6 \text{ fb}^{-1}$ ,  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$6.1 \pm 3.9 \pm 1.8$  BARATE 98 ALEP 1991–1995 LEP runs

• • • We use the following data for averages but not for fits. • • •

$7.5 \pm 2.6 \pm 1.8$  <sup>2</sup> RICHICHI 99 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<17$  95 ABBIENDI 00D OPAL 1990–1995 LEP runs

<sup>1</sup> Not independent of ARMS 05  $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0,\omega)) / \Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow K^-\omega\nu_\tau) / \Gamma_{\text{total}}$  values.

<sup>2</sup> Not independent of RICHICHI 99

$\Gamma(\tau^- \rightarrow K^-h^+\pi^-\nu_\tau(\text{ex.}K^0)) / \Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))$ ,  $\Gamma(\tau^- \rightarrow K^-K^+\pi^-\nu_\tau) / \Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))$  and BAEST 95c  $\Gamma(\tau^- \rightarrow h^-h^-h^+\nu_\tau(\text{ex.}K^0)) / \Gamma_{\text{total}}$  values.

$$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0,\eta))/\Gamma_{\text{total}} \quad \Gamma_{101}/\Gamma = (\Gamma_{103} + 0.892\Gamma_{179})/\Gamma$$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>DOCUMENT ID</u>
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**7.6±1.2 OUR FIT**

$$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0,\omega)) / \Gamma_{\text{total}} \quad \Gamma_{102}/\Gamma$$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>3.7 \pm 0.5 \pm 0.8</math></b>	833	ARMS	05	CLE3 $7.6 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0,\omega,\eta))/\Gamma_{\text{total}} \quad \Gamma_{103}/\Gamma$$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>DOCUMENT ID</u>
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**3.9±1.4 OUR FIT**

$$\Gamma(K^-\pi^+K^-\geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{104}/\Gamma$$

<u>VALUE (%)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt;0.09</math></b>	95	BAUER	94	TPC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

$$\Gamma(K^-K^+\pi^-\geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{105}/\Gamma = (\Gamma_{106} + \Gamma_{107})/\Gamma$$

<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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 **$0.1496 \pm 0.0033$  OUR FIT****0.203 ± 0.031 OUR AVERAGE**

$0.159 \pm 0.053 \pm 0.020$  ABBIENDI 00D OPAL 1990–1995 LEP runs

$0.15 \begin{array}{l} +0.09 \\ -0.07 \end{array} \pm 0.03$  4 <sup>1</sup> BAUER 94 TPC  $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

$0.238 \pm 0.042$  <sup>2</sup> BARATE 98 ALEP 1991–1995 LEP runs

<sup>1</sup> We multiply 0.15% by 0.20, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

<sup>2</sup> Not independent of BARATE 98  $\Gamma(\tau^- \rightarrow K^-K^+\pi^-\nu_\tau) / \Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow K^-K^+\pi^-\pi^0\nu_\tau) / \Gamma_{\text{total}}$  values.

$\Gamma(K^- K^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{106}/\Gamma$																																				
<u>VALUE (units <math>10^{-3}</math>)</u>	<u>EVTS</u>																																				
<b><math>1.435 \pm 0.027</math> OUR FIT</b>	<u>DOCUMENT ID</u>																																				
<b><math>1.43 \pm 0.07</math> OUR AVERAGE</b>	Error includes scale factor of 2.4. See the ideogram below.																																				
1.55 $\pm 0.01$ $\pm 0.06$ 1.346 $\pm 0.010$ $\pm 0.036$ 1.55 $\pm 0.06$ $\pm 0.09$ 1.63 $\pm 0.21$ $\pm 0.17$ $\bullet \bullet \bullet$ We use the following data for averages but not for fits. • • • 0.87 $\pm 0.56$ $\pm 0.40$ 1.45 $\pm 0.13$ $\pm 0.28$ $\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. • • • 2.2 $\pm 1.7$ $\pm 0.5$	1 LEE 10 BELL 666 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV 2 AUBERT 08 BABR 342 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV 3 BRIERE 03 CLE3 $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV BARATE 98 ALEP 1991–1995 LEP runs ABBIENDI 00D OPAL 1990–1995 LEP runs 4 RICHICHI 99 CLEO $E_{\text{cm}}^{\text{ee}} = 10.6$ GeV MILLS 85 DLCO $E_{\text{cm}}^{\text{ee}} = 29$ GeV																																				
<sup>1</sup> See footnote to LEE 10 $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ measurement for correlations with other measurements. Not independent of LEE 10 $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$ value.																																					
<sup>2</sup> See footnote to AUBERT 08 $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ measurement for correlations with other measurements.																																					
<sup>3</sup> 71% correlated with BRIERE 03 $\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau$ and 34% correlated with $\tau \rightarrow K^- \pi^+ \pi^- \nu_\tau$ because of a common 5% normalization error.																																					
<sup>4</sup> Not independent of RICHICHI 99 $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$ and BAEST 95C $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ values.																																					
<sup>5</sup> Error correlated with MILLS 85 ( $K \pi \pi \pi^0 \nu$ ) value. We multiply 0.22% by 0.23, the relative systematic error quoted by MILLS 85, to obtain the systematic error.																																					
<p style="text-align: center;">WEIGHTED AVERAGE <math>1.43 \pm 0.07</math> (Error scaled by 2.4)</p> <p>Values above of weighted average, error, and scale factor are based upon the data in this ideogram only. They are not necessarily the same as our 'best' values, obtained from a least-squares constrained fit utilizing measurements of other (related) quantities as additional information.</p>																																					
<table style="margin-left: auto; margin-right: auto;"> <thead> <tr> <th></th> <th></th> <th style="text-align: right;"><math>\chi^2</math></th> </tr> </thead> <tbody> <tr> <td>LEE</td> <td>10</td> <td style="text-align: right;">BELL</td> <td style="text-align: right;">5.7</td> </tr> <tr> <td>AUBERT</td> <td>08</td> <td style="text-align: right;">BABR</td> <td style="text-align: right;">4.9</td> </tr> <tr> <td>BRIERE</td> <td>03</td> <td style="text-align: right;">CLE3</td> <td style="text-align: right;">1.3</td> </tr> <tr> <td>ABBIENDI</td> <td>00D</td> <td style="text-align: right;">OPAL</td> <td></td> </tr> <tr> <td>RICHICHI</td> <td>99</td> <td style="text-align: right;">CLEO</td> <td></td> </tr> <tr> <td>BARATE</td> <td>98</td> <td style="text-align: right;">ALEP</td> <td></td> </tr> <tr> <td colspan="3" style="text-align: right;"><math>\chi^2</math></td></tr> <tr> <td colspan="3" style="text-align: right;">11.8</td></tr> <tr> <td colspan="3" style="text-align: right;">(Confidence Level = 0.0027)</td></tr> </tbody> </table> <p style="text-align: center;"><math>\Gamma(K^- K^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}</math> (units <math>10^{-3}</math>)</p>				$\chi^2$	LEE	10	BELL	5.7	AUBERT	08	BABR	4.9	BRIERE	03	CLE3	1.3	ABBIENDI	00D	OPAL		RICHICHI	99	CLEO		BARATE	98	ALEP		$\chi^2$			11.8			(Confidence Level = 0.0027)		
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$$\Gamma(K^- K^+ \pi^- \nu_\tau) / \Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{106}/\Gamma_{68}$$

$$\Gamma_{106}/\Gamma_{68} = \Gamma_{106}/(\Gamma_{70} + 0.0153\Gamma_{178})$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.592 ± 0.030 OUR FIT****1.83 ± 0.05 OUR AVERAGE**1.60 ± 0.15 ± 0.30 2.3k RICHICHI 99 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ 

• • • We use the following data for averages but not for fits. • • •

1.84 ± 0.01 ± 0.05 108k <sup>1</sup> LEE 10 BELL 666 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ <sup>1</sup> Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  values.

$$\Gamma(K^- K^+ \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{107}/\Gamma$$

VALUE (units 10 <sup>-4</sup> )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.61 ± 0.18 OUR FIT****0.60 ± 0.18 OUR AVERAGE**0.55 ± 0.14 ± 0.12 48 ARMS 05 CLE3 7.6 fb<sup>-1</sup>,  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ 

7.5 ± 2.9 ± 1.5 BARATE 98 ALEP 1991–1995 LEP runs

• • • We use the following data for averages but not for fits. • • •

3.3 ± 1.8 ± 0.7 158 <sup>1</sup> RICHICHI 99 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

&lt;27 95 ABBIENDI 00D OPAL 1990–1995 LEP runs

<sup>1</sup> Not independent of RICHICHI 99 $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  and BAEST 95C  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  values.

$$\Gamma(K^- K^+ \pi^- \pi^0 \nu_\tau) / \Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{107}/\Gamma_{77}$$

$$\Gamma_{107}/\Gamma_{77} = \Gamma_{107}/(\Gamma_{78} + 0.892\Gamma_{178} + 0.0153\Gamma_{180})$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.14 ± 0.04 OUR FIT****0.79 ± 0.44 ± 0.16** 158 <sup>1</sup> RICHICHI 99 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ <sup>1</sup> RICHICHI 99 also quote a 95%CL upper limit of 0.0157 for this measurement.

$$\Gamma(K^- K^+ K^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{108}/\Gamma = 0.491 \Gamma_{170}/\Gamma$$

VALUE (units 10 <sup>-5</sup> )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
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**2.2 ± 0.8 OUR FIT** Error includes scale factor of 5.4.**2.1 ± 0.8 OUR AVERAGE** Error includes scale factor of 5.4.3.29 ± 0.17<sup>+0.19</sup><sub>-0.20</sub> 3.2k <sup>1</sup> LEE 10 BELL 666 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ 1.58 ± 0.13 ± 0.12 275 <sup>2</sup> AUBERT 08 BABR 342 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 3.7 90 BRIERE 03 CLE3  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ 

&lt; 19 90 BARATE 98 ALEP 1991–1995 LEP runs

<sup>1</sup> See footnote to LEE 10  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  measurement for correlations with other measurements. Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  value.<sup>2</sup> See footnote to AUBERT 08  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

$\Gamma(K^- K^+ K^- \nu_\tau)/\Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$   $\Gamma_{108}/\Gamma_{68}$ 

<u>VALUE</u> (units $10^{-4}$ )	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$3.90 \pm 0.02^{+0.22}_{-0.23}$	3.2k	<sup>1</sup> LEE	10	BELL $666 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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<sup>1</sup> Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau)/\Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  values.

 $\Gamma(K^- K^+ K^- \nu_\tau (\text{ex. } \phi))/\Gamma_{\text{total}}$   $\Gamma_{109}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<2.5 \times 10^{-6}$	90	AUBERT	08	BABR $342 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(K^- K^+ K^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{110}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<4.8 \times 10^{-6}$	90	ARMS	05	CLE3 $7.6 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\pi^- K^+ \pi^- \geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{111}/\Gamma$ 

<u>VALUE (%)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<0.25$	95	BAUER	94	TPC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

 $\Gamma(e^- e^- e^+ \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{112}/\Gamma$ 

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$2.8 \pm 1.4 \pm 0.4$	5	ALAM	96	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\mu^- e^- e^+ \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{113}/\Gamma$ 

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<3.2$	90	ALAM	96	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\pi^- e^- e^+ \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{114}/\Gamma$ 

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
seen	400	<sup>1</sup> JIN	19	BELL $562 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.46 \pm 0.13 \pm 0.21$	400	<sup>1</sup> JIN	19	BELL axial-vector, $562 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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$3.01 \pm 0.27 \pm 0.43$	400	<sup>1</sup> JIN	19	BELL vector, $562 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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<sup>1</sup> JIN 19 measures  $B(\tau^- \rightarrow \pi^- e^- e^+ \nu_\tau (m_{\pi^- e^- e^+} > 1.05 \text{ GeV}/c^2)) = (5.90 \pm 0.53 \pm 0.86) \times 10^{-6}$ , which is only sensitive to the structure-dependent contribution, and assumes that the decay proceeds with either a pure axial-vector current or a pure vector current to obtain the two respective branching fraction measurements for this mode, which are 100% correlated.

 $\Gamma(\pi^- \mu^- \mu^+ \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{115}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<1.14 \times 10^{-5}$	90	JIN	19	BELL $562 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$$\Gamma(3h^- 2h^+ \geq 0 \text{ neutrals } \nu_\tau (\text{ex. } K_S^0 \rightarrow \pi^- \pi^+) \text{ ("5-prong")}) / \Gamma_{\text{total}} \quad \Gamma_{116} / \Gamma$$

$\Gamma_{116} / \Gamma = (\Gamma_{117} + \Gamma_{123}) / \Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.099 ± 0.004 OUR FIT</b>				
<b>0.107 ± 0.007 OUR AVERAGE</b> Error includes scale factor of 1.1.				
0.170 ± 0.022 ± 0.026		<sup>1</sup> ACHARD 01D L3		1992–1995 LEP runs
0.097 ± 0.005 ± 0.011	419	GIBAUT 94B CLEO		$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
0.102 ± 0.029	13	BYLSMA 87 HRS		$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
• • • We use the following data for averages but not for fits. • • •				
0.093 ± 0.009 ± 0.012		SCHAEL 05C ALEP		1991–1995 LEP runs
0.115 ± 0.013 ± 0.006	112	<sup>2</sup> ABREU 01M DLPH		1992–1995 LEP runs
0.119 ± 0.013 ± 0.008	119	<sup>3</sup> ACKERSTAFF 99E OPAL		1991–1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.26 ± 0.06 ± 0.05		ACTON 92H OPAL		$E_{\text{cm}}^{\text{ee}} = 88.2\text{--}94.2 \text{ GeV}$
0.10 <sup>+0.05</sup> <sub>-0.04</sub> ± 0.03		DECAMP 92C ALEP		1989–1990 LEP runs
0.16 ± 0.13 ± 0.04		BEHREND 89B CELL		$E_{\text{cm}}^{\text{ee}} = 14\text{--}47 \text{ GeV}$
0.3 ± 0.1 ± 0.2		BARTEL 85F JADE		$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$
0.13 ± 0.04	10	BELTRAMI 85 HRS		Repl. by BYLSMA 87
0.16 ± 0.08 ± 0.04	4	BURCHAT 85 MRK2		$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
1.0 ± 0.4	10	BEHREND 82 CELL		Repl. by BEHREND 89B

<sup>1</sup> The correlation coefficients between this measurement and the ACHARD 01D measurements of  $B(\tau \rightarrow \text{"1-prong"})$  and  $B(\tau \rightarrow \text{"3-prong"})$  are  $-0.082$  and  $-0.19$  respectively.

<sup>2</sup> The correlation coefficients between this measurement and the ABREU 01M measurements of  $B(\tau \rightarrow \text{1-prong})$  and  $B(\tau \rightarrow \text{3-prong})$  are  $-0.08$  and  $-0.08$  respectively.

<sup>3</sup> Not independent of ACKERSTAFF 99E  $B(\tau^- \rightarrow 3h^- 2h^+ \nu_\tau (\text{ex. } K^0))$  and  $B(\tau^- \rightarrow 3h^- 2h^+ \pi^0 \nu_\tau (\text{ex. } K^0))$  measurements.

$$\Gamma(3h^- 2h^+ \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{117} / \Gamma = (\Gamma_{118} + \Gamma_{120} + 0.0153 \Gamma_{185}) / \Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.29 ± 0.31 OUR FIT</b>				
<b>8.32 ± 0.35 OUR AVERAGE</b>				
9.7 ± 1.5 ± 0.5	96	<sup>1</sup> ABDALLAH 06A DLPH		1992–1995 LEP runs
7.2 ± 0.9 ± 1.2	165	<sup>2</sup> SCHAEL 05C ALEP		1991–1995 LEP runs
9.1 ± 1.4 ± 0.6	97	ACKERSTAFF 99E OPAL		1991–1995 LEP runs
7.7 ± 0.5 ± 0.9	295	GIBAUT 94B CLEO		$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
6.4 ± 2.3 ± 1.0	12	ALBRECHT 88B ARG		$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
5.1 ± 2.0	7	BYLSMA 87 HRS		$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

8.56 ± 0.05 ± 0.42    34k    AUBERT,B 05W BABR  $232 \text{ fb}^{-1}$ ,  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

8.0 ± 1.1 ± 1.3    58    BUSKULIC 96 ALEP    Repl. by SCHAEL 05C

6.7 ± 3.0    5    <sup>3</sup> BELTRAMI 85 HRS    Repl. by BYLSMA 87

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>3</sup> The error quoted is statistical only.

$\Gamma(3\pi^- 2\pi^+ \nu_\tau (\text{ex. } K^0, \omega)) / \Gamma_{\text{total}}$   $\Gamma_{118}/\Gamma = (\Gamma_{119} + \Gamma_{173})/\Gamma$ 

<u>VALUE</u> (units $10^{-4}$ )	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>8.27 \pm 0.31</math> OUR FIT</b>			

• • • We use the following data for averages but not for fits. • • •

<b><math>8.33 \pm 0.04 \pm 0.43</math></b>	<sup>1</sup> LEES	12X BABR	$468 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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<sup>1</sup> Not independent of LEES 12X  $\Gamma(\tau^- \rightarrow f_1(1285) \pi^- \nu_\tau \rightarrow 3\pi^- 2\pi^+ \nu_\tau) / \Gamma$  and  $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \nu_\tau (\text{ex. } K^0, \omega, f_1(1285))) / \Gamma$  values.

 $\Gamma(3\pi^- 2\pi^+ \nu_\tau (\text{ex. } K^0, \omega, f_1(1285))) / \Gamma_{\text{total}}$   $\Gamma_{119}/\Gamma$ 

<u>VALUE</u> (units $10^{-4}$ )	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>7.75 \pm 0.30</math> OUR FIT</b>				

<b><math>7.68 \pm 0.04 \pm 0.40</math></b>	69k	LEES	12X BABR	$468 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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 $\Gamma(K^- 2\pi^- 2\pi^+ \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$   $\Gamma_{120}/\Gamma$ 

<u>VALUE</u> (units $10^{-6}$ )	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>0.6 \pm 1.2</math> OUR FIT</b>			

<b><math>0.6 \pm 0.5 \pm 1.1</math></b>	<sup>1</sup> LEES	12X BABR	$468 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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<sup>1</sup> LEES 12X measurement corresponds to the lower limit of  $< 2.4 \times 10^{-6}$  at 90% CL.

 $\Gamma(K^+ 3\pi^- \pi^+ \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{121}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt;5.0 \times 10^{-6}</math></b>	90	LEES	12X BABR	$468 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(K^+ K^- 2\pi^- \pi^+ \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{122}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt;4.5 \times 10^{-7}</math></b>	90	LEES	12X BABR	$468 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(3h^- 2h^+ \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$   $\Gamma_{123}/\Gamma = (\Gamma_{124} + \Gamma_{127})/\Gamma$ 

<u>VALUE</u> (units $10^{-4}$ )	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>1.65 \pm 0.11</math> OUR FIT</b>				

 **$1.74 \pm 0.27$  OUR AVERAGE**

$1.6 \pm 1.2 \pm 0.6$	13	<sup>1</sup> ABDALLAH	06A	DLPH	1992–1995 LEP runs
$2.1 \pm 0.7 \pm 0.9$	95	<sup>2</sup> SCHAEL	05C	ALEP	1991–1995 LEP runs
$1.7 \pm 0.2 \pm 0.2$	231	ANASTASSOV	01	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$2.7 \pm 1.8 \pm 0.9$	23	ACKERSTAFF	99E	OPAL	1991–1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$1.8 \pm 0.7 \pm 1.2$	18	BUSKULIC	96	ALEP	Repl. by SCHAEL 05C
$1.9 \pm 0.4 \pm 0.4$	31	GIBAUT	94B	CLEO	Repl. by ANASTASSOV 01
$5.1 \pm 2.2$	6	BYLSMA	87	HRS	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
$6.7 \pm 3.0$	5	<sup>3</sup> BELTRAMI	85	HRS	Repl. by BYLSMA 87

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> SCHAEL 05C quote  $(1.4 \pm 0.7 \pm 0.9) \times 10^{-4}$ . We add  $0.7 \times 10^{-4}$  to remove their correction for  $\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau$  and  $\tau^- \rightarrow K^*(892)^- \eta \nu_\tau \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau$  decays. See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>3</sup> The error quoted is statistical only.

$$\Gamma(3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex.} K^0)) / \Gamma_{\text{total}} \quad \Gamma_{124}/\Gamma$$

$$\Gamma_{124}/\Gamma = (\Gamma_{126} + 0.2302\Gamma_{160} + 0.892\Gamma_{185})/\Gamma$$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.63 ± 0.11 OUR FIT</b>			

• • • We use the following data for averages but not for fits. • • •

$$\mathbf{1.65 \pm 0.05 \pm 0.09} \quad {}^1 \text{LEES} \quad 12X \text{ BABR} \quad 468 \text{ fb}^{-1} \quad E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$$

<sup>1</sup> Not independent of LEES 12X measurements of  $\Gamma(\tau^- \rightarrow 2\pi^- \pi^+ \omega \nu_\tau (\text{ex.} K^0)) / \Gamma$ ,  $\Gamma(\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0)) / \Gamma$ , and  $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex.} K^0, \eta, \omega, f_1(1285))) / \Gamma$ .

$$\Gamma(3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex.} K^0, \eta, f_1(1285))) / \Gamma_{\text{total}} \quad \Gamma_{125}/\Gamma$$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.11 ± 0.04 ± 0.09</b>	<sup>1</sup> LEES	12X	BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> Not independent of LEES 12X  $\Gamma(\tau^- \rightarrow 2\pi^- \pi^+ \omega \nu_\tau (\text{ex.} K^0)) / \Gamma$  and  $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex.} K^0, \eta, \omega, f_1(1285))) / \Gamma$  values.

$$\Gamma(3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex.} K^0, \eta, \omega, f_1(1285))) / \Gamma_{\text{total}} \quad \Gamma_{126}/\Gamma$$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.38 ± 0.09 OUR FIT</b>				

$$\mathbf{0.36 \pm 0.03 \pm 0.09} \quad 7.3k \quad \text{LEES} \quad 12X \text{ BABR} \quad 468 \text{ fb}^{-1} \quad E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$$

$$\Gamma(K^- 2\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex.} K^0)) / \Gamma_{\text{total}} \quad \Gamma_{127}/\Gamma$$

<u>VALUE (units <math>10^{-6}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.1 ± 0.6 OUR FIT</b>			

$$\mathbf{1.1 \pm 0.4 \pm 0.4} \quad {}^1 \text{LEES} \quad 12X \text{ BABR} \quad 468 \text{ fb}^{-1} \quad E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$$

<sup>1</sup> LEES 12X measurement corresponds to the lower limit of  $< 1.9 \times 10^{-6}$  at 90% CL.

$$\Gamma(K^+ 3\pi^- \pi^+ \pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{128}/\Gamma$$

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt; 8 \times 10^{-7}</math></b>	90	LEES	12X	BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$$\Gamma(3h^- 2h^+ 2\pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{129}/\Gamma$$

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt; 3.4 \times 10^{-6}</math></b>	90	AUBERT,B	06	BABR $232 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$$< 1.1 \times 10^{-4} \quad 90 \quad \text{GIBAUT} \quad 94B \text{ CLEO} \quad E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$$

$$\Gamma((5\pi)^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{130}/\Gamma$$

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.78 ± 0.05 OUR FIT</b>				

• • • We use the following data for averages but not for fits. • • •

$$\mathbf{0.61 \pm 0.06 \pm 0.08} \quad {}^1 \text{GIBAUT} \quad 94B \text{ CLEO} \quad E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$$

<sup>1</sup> Not independent of GIBAUT 94B  $B(3h^- 2h^+ \nu_\tau)$ , PROCARIO 93  $B(h^- 4\pi^0 \nu_\tau)$ , and BORTOLETTO 93  $B(2h^- h^+ 2\pi^0 \nu_\tau) / B(\text{"3prong"})$  measurements. Result is corrected for  $\eta$  contributions.

$\Gamma(4h^- 3h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ ("7-prong")})/\Gamma_{\text{total}}$   $\Gamma_{131}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.0 \times 10^{-7}$	90	AUBERT,B	05F	BABR $232 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$<1.8 \times 10^{-5}$	95	ACKERSTAFF	97J	OPAL 1990–1995 LEP runs
$<2.4 \times 10^{-6}$	90	EDWARDS	97B	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.9 \times 10^{-4}$	90	BYLSMA	87	HRS $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

 $\Gamma(4h^- 3h^+ \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{132}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<4.3 \times 10^{-7}$	90	AUBERT,B	05F	BABR $232 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(4h^- 3h^+ \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{133}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.5 \times 10^{-7}$	90	AUBERT,B	05F	BABR $232 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(X^-(S=-1)\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{134}/\Gamma$ 

$$\Gamma_{134}/\Gamma = (\Gamma_{10} + \Gamma_{16} + \Gamma_{23} + \Gamma_{28} + \Gamma_{36} + \Gamma_{41} + \Gamma_{45} + \Gamma_{61} + \Gamma_{97} + \Gamma_{103} + \Gamma_{120} + \Gamma_{127} + \Gamma_{152} + \Gamma_{154} + \Gamma_{156} + 0.8312\Gamma_{170} + \Gamma_{179})/\Gamma$$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>2.92 ± 0.04 OUR FIT</b>			

$\bullet \bullet \bullet$  We use the following data for averages but not for fits.  $\bullet \bullet \bullet$

**2.87 ± 0.12** <sup>1</sup> BARATE 99R ALEP 1991–1995 LEP runs

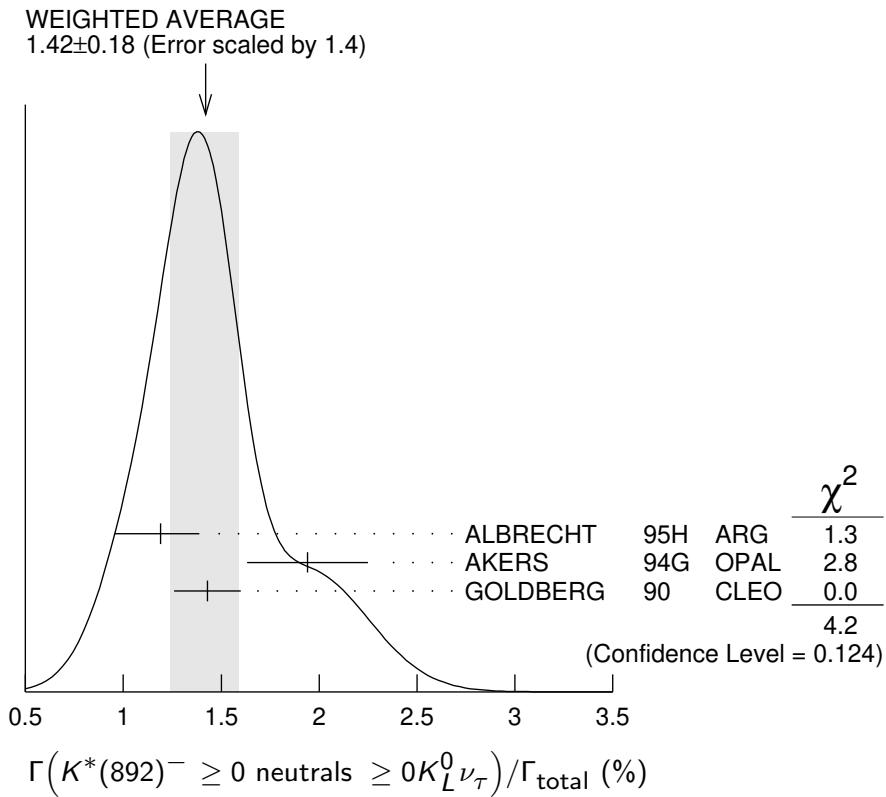
<sup>1</sup> BARATE 99R perform a combined analysis of all ALEPH LEP 1 data on  $\tau$  branching fraction measurements for decay modes having total strangeness equal to  $-1$ .

 $\Gamma(K^*(892)^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{135}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.42 ± 0.18 OUR AVERAGE</b>				Error includes scale factor of 1.4. See the ideogram below.
$1.19 \pm 0.15^{+0.13}_{-0.18}$	104	ALBRECHT	95H	ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
$1.94 \pm 0.27 \pm 0.15$	74	<sup>1</sup> AKERS	94G	OPAL $E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$

<sup>1</sup> AKERS 94G reject events in which a  $K_S^0$  accompanies the  $K^*(892)^-$ . We do not correct for them.

<sup>2</sup> GOLDBERG 90 estimates that 10% of observed  $K^*(892)$  are accompanied by a  $\pi^0$ .



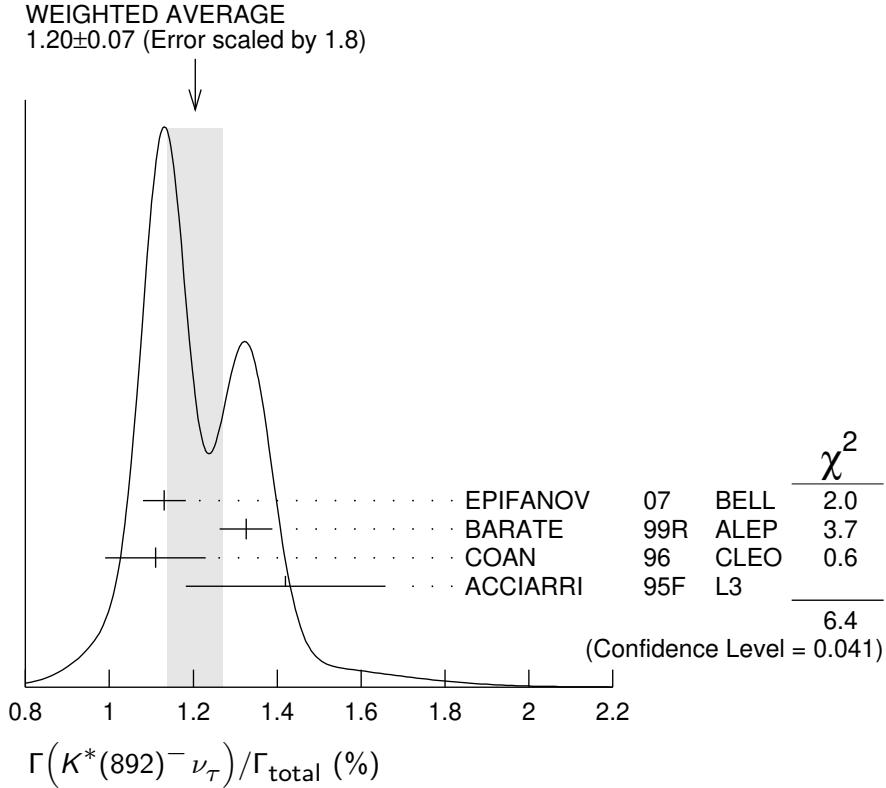
$\Gamma(K^*(892)^- \nu_\tau) / \Gamma_{\text{total}}$	$\Gamma_{136} / \Gamma$
<b>1.20 ±0.07 OUR AVERAGE</b>	Error includes scale factor of 1.8. See the ideogram below.
1.131±0.006±0.051	49k
1.326±0.063	
1.11 ±0.12	
1.42 ±0.22 ±0.09	
• • • We do not use the following data for averages, fits, limits, etc. • • •	
1.39 ±0.09 ±0.10	
1.45 ±0.13 ±0.11	273
1.23 ±0.21 +0.11 -0.21	54
1.9 ±0.3 ±0.4	44
1.5 ±0.4 ±0.4	15
1.3 ±0.3 ±0.3	31
1.7 ±0.7	11
4 EPIFANOV 07	BELL 351 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
5 BUSKULIC 94F	ALEP 1991–1995 LEP runs
6 ALBRECHT 88L	CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6$ GeV
7 TSCHIRHART 88	HRS $E_{\text{cm}}^{\text{ee}}=29$ GeV
8 AIHARA 87C	TPC $E_{\text{cm}}^{\text{ee}}=29$ GeV
YELTON 86	MRK2 $E_{\text{cm}}^{\text{ee}}=29$ GeV
DORFAN 81	MRK2 $E_{\text{cm}}^{\text{ee}}=4.2\text{--}6.7$ GeV

<sup>1</sup> EPIFANOV 07 quote  $B(\tau^- \rightarrow K^*(892)^- \nu_\tau) B(K^*(892)^- \rightarrow K_S^0 \pi^-) = (3.77 \pm 0.02(\text{stat}) \pm 0.12(\text{syst}) \pm 0.12(\text{mod})) \times 10^{-3}$ . We add the systematic and model uncertainties in quadrature and divide by  $B(K^*(892)^- \rightarrow K_S^0 \pi^-) = 0.3333$ .

<sup>2</sup> Not independent of COAN 96  $B(\pi^- \bar{K}^0 \nu_\tau)$  and BATTLE 94  $B(K^- \pi^0 \nu_\tau)$  measurements.  $K\pi$  final states are consistent with and assumed to originate from  $K^*(892)^-$  production.

<sup>3</sup> This result is obtained from their  $B(\pi^- \bar{K}^0 \nu_\tau)$  assuming all those decays originate in  $K^*(892)^-$  decays.

- <sup>4</sup> Not independent of BUSKULIC 96  $B(\pi^- \bar{K}^0 \nu_\tau)$  and  $B(K^- \pi^0 \nu_\tau)$  measurements.  
<sup>5</sup> BUSKULIC 94F obtain this result from BUSKULIC 94F  $B(\bar{K}^0 \pi^- \nu_\tau)$  and BUSKULIC 94E  $B(K^- \pi^0 \nu_\tau)$  assuming all of those decays originate in  $K^*(892)^-$  decays.  
<sup>6</sup> The authors divide by  $\Gamma_2/\Gamma = 0.865$  to obtain this result.  
<sup>7</sup> Not independent of TSCHIRHART 88  $\Gamma(\tau^- \rightarrow h^- \bar{K}^0 \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma$ .  
<sup>8</sup> Decay  $\pi^-$  identified in this experiment, is assumed in the others.



### $\Gamma(K^*(892)^- \nu_\tau) / \Gamma(\pi^- \pi^0 \nu_\tau)$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.075±0.027</b>	<sup>1</sup> ABREU 94K	DLPH	LEP 1992 Z data

<sup>1</sup> ABREU 94K quote  $B(\tau^- \rightarrow K^*(892)^- \nu_\tau)B(K^*(892)^- \rightarrow K^- \pi^0)/B(\tau^- \rightarrow \rho^- \nu_\tau) = 0.025 \pm 0.009$ . We divide by  $B(K^*(892)^- \rightarrow K^- \pi^0) = 0.333$  to obtain this result.

### $\Gamma(K^*(892)^- \nu_\tau \rightarrow \pi^- \bar{K}^0 \nu_\tau) / \Gamma(\pi^- \bar{K}^0 \nu_\tau)$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.933±0.027</b>	49k	EPIFANOV 07	BELL	$351 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

### $\Gamma(K^*(892)^0 K^- \geq 0 \text{ neutrals} \nu_\tau) / \Gamma_{\text{total}}$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.32±0.08±0.12</b>	119	GOLDBERG 90	CLEO	$E_{\text{cm}}^{ee} = 9.4\text{--}10.9 \text{ GeV}$

### $\Gamma(K^*(892)^0 K^- \nu_\tau) / \Gamma_{\text{total}}$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.21 ± 0.04 OUR AVERAGE</b>				
0.213 ± 0.048		<sup>1</sup> BARATE 98	ALEP	1991–1995 LEP runs
0.20 ± 0.05 ± 0.04	47	ALBRECHT 95H	ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$

<sup>1</sup> BARATE 98 measure the  $K^- (\rho^0 \rightarrow \pi^+ \pi^-)$  fraction in  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  decays to be  $(35 \pm 11)\%$  and derive this result from their measurement of  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$  assuming the intermediate states are all  $K^- \rho$  and  $K^- K^*(892)^0$ .

$\Gamma(\bar{K}^*(892)^0 \pi^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{140}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.38 ± 0.11 ± 0.13</b>	105	GOLDBERG	90	CLEO $E_{\text{cm}}^{ee} = 9.4\text{--}10.9 \text{ GeV}$

$\Gamma(\bar{K}^*(892)^0 \pi^- \nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{141}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.22 ± 0.05 OUR AVERAGE</b>				
0.209 ± 0.058		1 BARATE	98	ALEP    1991–1995 LEP runs
0.25 ± 0.10 ± 0.05	27	ALBRECHT	95H	ARG $E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$

<sup>1</sup> BARATE 98 measure the  $K^- K^*(892)^0$  fraction in  $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$  decays to be  $(87 \pm 13)\%$  and derive this result from their measurement of  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$ .

$\Gamma((\bar{K}^*(892) \pi)^- \nu_\tau \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{142}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.10 ± 0.04 OUR AVERAGE</b>				
0.097 ± 0.044 ± 0.036		1 BARATE	99K	ALEP    1991–1995 LEP runs
0.106 ± 0.037 ± 0.032		2 BARATE	98E	ALEP    1991–1995 LEP runs

<sup>1</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter. They determine the  $\bar{K}^0 \rho^-$  fraction in  $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$  decays to be  $(0.72 \pm 0.12 \pm 0.10)$  and multiply their  $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$  measurement by one minus this fraction to obtain the quoted result.

<sup>2</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays. They determine the  $\bar{K}^0 \rho^-$  fraction in  $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$  decays to be  $(0.64 \pm 0.09 \pm 0.10)$  and multiply their  $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$  measurement by one minus this fraction to obtain the quoted result.

$\Gamma(K_1(1270)^- \nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{143}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.47 ± 0.11 OUR AVERAGE</b>				
0.48 ± 0.11		BARATE	99R	ALEP    1991–1995 LEP runs
$0.41^{+0.41}_{-0.35} \pm 0.10$	5	<sup>1</sup> BAUER	94	TPC $E_{\text{cm}}^{ee} = 29 \text{ GeV}$

<sup>1</sup> We multiply 0.41% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

$\Gamma(K_1(1400)^- \nu_\tau)/\Gamma_{\text{total}}$				$\Gamma_{144}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.17 ± 0.26 OUR AVERAGE</b>				Error includes scale factor of 1.7.
0.05 ± 0.17		BARATE	99R	ALEP    1991–1995 LEP runs
$0.76^{+0.40}_{-0.33} \pm 0.20$	11	<sup>1</sup> BAUER	94	TPC $E_{\text{cm}}^{ee} = 29 \text{ GeV}$

<sup>1</sup> We multiply 0.76% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

$[\Gamma(K_1(1270)^-\nu_\tau) + \Gamma(K_1(1400)^-\nu_\tau)]/\Gamma_{\text{total}}$	$(\Gamma_{143} + \Gamma_{144})/\Gamma$			
<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>1.17^{+0.41}_{-0.37} \pm 0.29</math></b>	16	<sup>1</sup> BAUER	94	TPC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

<sup>1</sup> We multiply 1.17% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error. Not independent of BAUER 94  $B(K_1(1270)^-\nu_\tau)$  and BAUER 94  $B(K_1(1400)^-\nu_\tau)$  measurements.

$\Gamma(K_1(1270)^-\nu_\tau)/[\Gamma(K_1(1270)^-\nu_\tau) + \Gamma(K_1(1400)^-\nu_\tau)]$	$\Gamma_{143}/(\Gamma_{143} + \Gamma_{144})$		
<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.69 ± 0.15 OUR AVERAGE</b>			

$0.71 \pm 0.16 \pm 0.11$	<sup>1</sup> ABBIENDI	00D	OPAL	1990–1995 LEP runs
$0.66 \pm 0.19 \pm 0.13$	<sup>2</sup> ASNER	00B	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> ABBIENDI 00D assume the resonance structure of  $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau$  decays is dominated by the  $K_1(1270)^-$  and  $K_1(1400)^-$  resonances.

<sup>2</sup> ASNER 00B assume the resonance structure of  $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau$  (ex.  $K^0$ ) decays is dominated by  $K_1(1270)^-$  and  $K_1(1400)^-$  resonances.

$\Gamma(K^*(1410)^-\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{145}/\Gamma$			
<u>VALUE (units <math>10^{-3}</math>)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>			
<b><math>1.5^{+1.4}_{-1.0}</math></b>	BARATE	99R	ALEP	1991–1995 LEP runs

$\Gamma(K_0^*(1430)^-\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{146}/\Gamma$			
<u>VALUE (units <math>10^{-3}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;0.5</b>	95	BARATE	99R	ALEP    1991–1995 LEP runs

$\Gamma(K_2^*(1430)^-\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{147}/\Gamma$			
<u>VALUE (%)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
<b>&lt;0.3</b>	95		TSCHIRHART 88	HRS $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<0.33$	95	<sup>1</sup> ACCIARRI	95F	L3    1991–1993 LEP runs
$<0.9$	95	0	DORFAN	81    MRK2 $E_{\text{cm}}^{\text{ee}} = 4.2\text{--}6.7 \text{ GeV}$

<sup>1</sup> ACCIARRI 95F quote  $B(\tau^- \rightarrow K^*(1430)^- \rightarrow \pi^-\bar{K}^0\nu_\tau) < 0.11\%$ . We divide by  $B(K^*(1430)^- \rightarrow \pi^-\bar{K}^0) = 0.33$  to obtain the limit shown.

$\Gamma(a_0(980)^-\geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}} \times B(a_0(980) \rightarrow K^0\bar{K}^-)$	$\Gamma_{148}/\Gamma \times B$			
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;2.8</b>	90	GOLDBERG	90	CLEO $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.9 \text{ GeV}$

$\Gamma(\eta\pi^-\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{149}/\Gamma$			
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
<b>&lt; 0.99</b>	95		<sup>1</sup> DEL-AMO-SA..11E	BABR $470 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 6.2$	95	BUSKULIC	97C	ALEP    1991–1994 LEP runs
$< 1.4$	95	0	BARTEL	96    CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

< 3.4	95	ARTUSO	92	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
< 90	95	ALBRECHT	88M	ARG	$E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$
<140	90	BEHREND	88	CELL	$E_{\text{cm}}^{\text{ee}} = 14\text{--}46.8 \text{ GeV}$
<180	95	BARINGER	87	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.5 \text{ GeV}$
<250	90	COFFMAN	87	MRK3	$E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$
510 $\pm 100 \pm 120$	65	DERRICK	87	HRS	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
<100	95	GAN	87B	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

<sup>1</sup> DEL-AMO-SANCHEZ 11E also quote  $B(\tau^- \rightarrow \eta\pi^-\nu_\tau) = (3.4 \pm 3.4 \pm 2.1) \times 10^{-5}$ .

### $\Gamma(\eta\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_{150}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.39 <math>\pm 0.07</math> OUR FIT</b>					

**1.38  $\pm 0.09$  OUR AVERAGE** Error includes scale factor of 1.2.

1.35 $\pm 0.03 \pm 0.07$	6.0k	INAMI	09	BELL	$490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
1.8 $\pm 0.4 \pm 0.2$		BUSKULIC	97C	ALEP	1991–1994 LEP runs
1.7 $\pm 0.2 \pm 0.2$	125	ARTUSO	92	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>					
< 11.0	95	ALBRECHT	88M	ARG	$E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$
< 21.0	95	BARINGER	87	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.5 \text{ GeV}$
42.0 $\pm 7.0 \pm 16.0$		<sup>1</sup> GAN	87	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

<sup>1</sup> Highly correlated with GAN 87  $\Gamma(\pi^-\pi^0\nu_\tau)/\Gamma(\text{total})$  value.

### $\Gamma(\eta\pi^-\pi^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_{151}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.0 <math>\pm 0.4</math> OUR FIT</b>					

**1.81  $\pm 0.31$  OUR AVERAGE**

2.01 $\pm 0.34 \pm 0.22$	381	LEES	12X	BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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**• • • We use the following data for averages but not for fits. • • •**

1.5  $\pm 0.5$  <sup>1</sup> ANASTASSOV 01 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

**• • • We do not use the following data for averages, fits, limits, etc. • • •**

1.4 $\pm 0.6 \pm 0.3$	15	<sup>2</sup> BERGFELD	97	CLEO	Repl. by ANASTASSOV 01
< 4.3	95	ARTUSO	92	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
<120	95	ALBRECHT	88M	ARG	$E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$

<sup>1</sup> Weighted average of BERGFELD 97 and ANASTASSOV 01 value of  $(1.5 \pm 0.6 \pm 0.3) \times 10^{-4}$  obtained using  $\eta$ 's reconstructed from  $\eta \rightarrow \pi^+\pi^-\pi^0$  decays.

<sup>2</sup> BERGFELD 97 reconstruct  $\eta$ 's using  $\eta \rightarrow \gamma\gamma$  decays.

### $\Gamma(\eta K^-\nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_{152}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.55 <math>\pm 0.08</math> OUR FIT</b>					

**1.54  $\pm 0.08$  OUR AVERAGE**

1.42 $\pm 0.11 \pm 0.07$	690	DEL-AMO-SA...11E	BABR	$470 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
1.58 $\pm 0.05 \pm 0.09$	1.6k	INAMI	09	BELL	$490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$2.9^{+1.3}_{-1.2} \pm 0.7$  BUSKULIC 97c ALEP 1991–1994 LEP runs

$2.6 \pm 0.5 \pm 0.5$  85 BARTELT 96 CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 4.7$  95 ARTUSO 92 CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

### $\Gamma(\eta K^*(892)^-\nu_\tau)/\Gamma_{\text{total}}$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.38 \pm 0.15</math> OUR AVERAGE</b>				

$1.34 \pm 0.12 \pm 0.09$  245 <sup>1</sup> INAMI 09 BELL  $490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$2.90 \pm 0.80 \pm 0.42$  25 BISHAI 99 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> Not independent of INAMI 09  $B(\tau^- \rightarrow \eta K^- \pi^0 \nu_\tau)$  and  $B(\tau^- \rightarrow \eta \bar{K}^0 \pi^- \nu_\tau)$  values.

### $\Gamma(\eta K^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.48 \pm 0.12</math> OUR FIT</b>				

### **$0.48 \pm 0.12$ OUR AVERAGE**

$0.46 \pm 0.11 \pm 0.04$  270 INAMI 09 BELL  $490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$1.77 \pm 0.56 \pm 0.71$  36 BISHAI 99 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(\eta K^- \pi^0 (\text{non-}K^*(892)) \nu_\tau)/\Gamma_{\text{total}}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 3.5 \times 10^{-5}$				

### $\Gamma(\eta \bar{K}^0 \pi^- \nu_\tau)/\Gamma_{\text{total}}$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.94 \pm 0.15</math> OUR FIT</b>				

### **$0.93 \pm 0.15$ OUR AVERAGE**

$0.88 \pm 0.14 \pm 0.06$  161 <sup>1</sup> INAMI 09 BELL  $490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$2.20 \pm 0.70 \pm 0.22$  15 <sup>2</sup> BISHAI 99 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> We multiply the INAMI 09 measurement  $B(\tau^- \rightarrow \eta K_S^0 \pi^- \nu_\tau) = (0.44 \pm 0.07 \pm 0.03) \times 10^{-4}$  by 2 to obtain the listed value.

<sup>2</sup> We multiply the BISHAI 99 measurement  $B(\tau^- \rightarrow \eta K_S^0 \pi^- \nu_\tau) = (1.10 \pm 0.35 \pm 0.11) \times 10^{-4}$  by 2 to obtain the listed value.

### $\Gamma(\eta \bar{K}^0 \pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 5.0 \times 10^{-5}$				

<sup>1</sup> We multiply the INAMI 09 measurement  $B(\tau^- \rightarrow \eta K_S^0 \pi^- \pi^0 \nu_\tau) < 2.5 \times 10^{-5}$  by 2 to obtain the listed value.

### $\Gamma(\eta K^- K^0 \nu_\tau)/\Gamma_{\text{total}}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 9.0 \times 10^{-6}$				

<sup>1</sup> We multiply the INAMI 09 measurement  $B(\tau^- \rightarrow \eta K^- K_S^0 \nu_\tau) < 4.5 \times 10^{-6}$  by 2 to obtain the listed value.

$\Gamma(\eta\pi^+\pi^-\pi^-\geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{159}/\Gamma$			
VALUE (%)	CL %	DOCUMENT ID	TECN	COMMENT
<0.3	90	ABACHI	87B HRS	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

$\Gamma(\eta\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$	$\Gamma_{160}/\Gamma$			
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>2.20 \pm 0.13</math> OUR FIT</b>				

**2.23 $\pm 0.12$  OUR AVERAGE**

$2.10 \pm 0.09 \pm 0.13$	2.9k	<sup>1</sup> LEES	12X BABR	$\eta \rightarrow \gamma\gamma$
$2.37 \pm 0.12 \pm 0.18$	1.4k	<sup>1</sup> LEES	12X BABR	$\eta \rightarrow \pi^+\pi^-\pi^0$
$2.54 \pm 0.27 \pm 0.25$	315	<sup>1</sup> LEES	12X BABR	$\eta \rightarrow 3\pi^0$
<b>• • • We use the following data for averages but not for fits. • • •</b>				
2.3 ± 0.5	170	<sup>2</sup> ANASTASSOV 01	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
1.60 ± 0.05 ± 0.11	1.8 k	AUBERT	08AE BABR	Repl. by LEES 12X
3.4 $^{+0.6}_{-0.5}$ ± 0.6	89	<sup>3</sup> BERGFELD	97	CLEO Repl. by ANASTASSOV 01

<sup>1</sup> LEES 12X uses  $468 \text{ fb}^{-1}$  of data taken at  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ . It gives the average of the three measurements listed here as  $(2.25 \pm 0.07 \pm 0.12) \times 10^{-4}$ .

<sup>2</sup> Weighted average of BERGFELD 97 and ANASTASSOV 01 measurements using  $\eta$ 's reconstructed from  $\eta \rightarrow \pi^+\pi^-\pi^0$  and  $\eta \rightarrow 3\pi^0$  decays.

<sup>3</sup> BERGFELD 97 reconstruct  $\eta$ 's using  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow 3\pi^0$  decays.

$\Gamma(\eta\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0,f_1(1285)))/\Gamma_{\text{total}}$	$\Gamma_{161}/\Gamma$		
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.99 \pm 0.09 \pm 0.13</math></b>	<sup>1</sup> LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> LEES 12X obtain this result by subtracting their  $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)$  measurement from their  $B(\tau^- \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))$  measurement.

$\Gamma(\eta a_1(1260)^-\nu_\tau \rightarrow \eta\pi^-\rho^0\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{162}/\Gamma$			
VALUE	CL %	DOCUMENT ID	TECN	COMMENT
< $3.9 \times 10^{-4}$	90	BERGFELD	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(\eta\eta\pi^-\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{163}/\Gamma$			
VALUE	CL %	DOCUMENT ID	TECN	COMMENT
< $7.4 \times 10^{-6}$	90	INAMI	09	BELL $490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
< $1.1 \times 10^{-4}$	95	ARTUSO	92	CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
< $8.3 \times 10^{-3}$	95	ALBRECHT	88M ARG	$E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$

$\Gamma(\eta\eta\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{164}/\Gamma$			
VALUE (units $10^{-4}$ )	CL %	DOCUMENT ID	TECN	COMMENT
< 2.0	95	ARTUSO	92	CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
<90	95	ALBRECHT	88M ARG	$E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$

$\Gamma(\eta\eta K^-\nu_\tau)/\Gamma_{\text{total}}$		$\Gamma_{165}/\Gamma$			
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
$<3.0 \times 10^{-6}$	90	INAMI	09	BELL	$490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(\eta'(958)\pi^-\nu_\tau)/\Gamma_{\text{total}}$		$\Gamma_{166}/\Gamma$			
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
$<4.0 \times 10^{-6}$	90	LEES	12X	BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$<7.2 \times 10^{-6}$	90	AUBERT	08AE	BABR	$384 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.4 \times 10^{-5}$	90	BERGFELD	97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(\eta'(958)\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$		$\Gamma_{167}/\Gamma$			
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
$<1.2 \times 10^{-5}$	90	LEES	12X	BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$<8.0 \times 10^{-5}$	90	BERGFELD	97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(\eta'(958)K^-\nu_\tau)/\Gamma_{\text{total}}$		$\Gamma_{168}/\Gamma$			
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
$<2.4 \times 10^{-6}$	90	LEES	12X	BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(\phi\pi^-\nu_\tau)/\Gamma_{\text{total}}$		$\Gamma_{169}/\Gamma$			
<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>3.42 \pm 0.55 \pm 0.25</math></b>	344	AUBERT	08	BABR	$342 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 20$	90	<sup>1</sup> AVERY	97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 35$	90	ALBRECHT	95H	ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$

<sup>1</sup> Avery 97 limit varies from  $(1.2\text{--}2.0) \times 10^{-4}$  depending on decay model assumptions.

$\Gamma(\phi K^-\nu_\tau)/\Gamma_{\text{total}}$		$\Gamma_{170}/\Gamma$			
<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>4.4 \pm 1.6</math> OUR FIT</b>					
<b><math>3.70 \pm 0.33</math> OUR AVERAGE</b>					Error includes scale factor of 1.3.
$\bullet \bullet \bullet$ We use the following data for averages but not for fits. $\bullet \bullet \bullet$					
$3.39 \pm 0.20 \pm 0.28$	274	AUBERT	08	BABR	$342 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$4.05 \pm 0.25 \pm 0.26$	551	INAMI	06	BELL	$401 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 6.7$	90	<sup>1</sup> AVERY	97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> Avery 97 limit varies from  $(5.4\text{--}6.7) \times 10^{-5}$  depending on decay model assumptions.

$\Gamma(f_1(1285)\pi^-\nu_\tau)/\Gamma_{\text{total}}$		$\Gamma_{171}/\Gamma$			
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b><math>3.9 \pm 0.5</math> OUR AVERAGE</b>				Error includes scale factor of 1.9.	
$4.73 \pm 0.28 \pm 0.45$	3.7k	<sup>1</sup> LEES	12X	BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$3.60 \pm 0.18 \pm 0.23$	2.5k	<sup>2</sup> LEES	12X	BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$3.19 \pm 0.18 \pm 1.00$	1.3 k	<sup>3</sup> AUBERT	08AE BABR	Repl. by LEES 12X
$3.9 \pm 0.7 \pm 0.5$	1.4 k	<sup>4</sup> AUBERT,B	05W BABR	Repl. by LEES 12X
$5.8^{+1.4}_{-1.3} \pm 1.8$	54	<sup>5</sup> BERGFELD	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> LEES 12X obtain this value by dividing their  $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow 3\pi^- 2\pi^+\nu_\tau)$  measurement by the PDG 12 value of  $B(f_1(1285) \rightarrow 2\pi^+ 2\pi^-) = 0.111^{+0.007}_{-0.006}$ .

<sup>2</sup> LEES 12X obtain this value by dividing their  $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)$  measurement by 2/3 of the PDG 12 value of  $B(f_1(1285) \rightarrow \eta\pi\pi) = 0.524^{+0.019}_{-0.021}$ .

<sup>3</sup> AUBERT 08AE obtain this value by dividing their  $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)$  measurement by the PDG 06 value of  $B(f_1(1285) \rightarrow \eta\pi^-\pi^+) = 0.35 \pm 0.11$ . The quote  $(3.19 \pm 0.18 \pm 0.16 \pm 0.99) \times 10^{-4}$  where the final error is due to the uncertainty on  $B(f_1(1285) \rightarrow \eta\pi^-\pi^+)$ . We combine the two systematic errors in quadrature.

<sup>4</sup> AUBERT,B 05W use the  $f_1(1285) \rightarrow 2\pi^+ 2\pi^-$  decay mode and the PDG 04 value of  $B(f_1(1285) \rightarrow 2\pi^+ 2\pi^-) = 0.110^{+0.007}_{-0.006}$ .

<sup>5</sup> BERGFELD 97 use the  $f_1(1285) \rightarrow \eta\pi^+\pi^-$  decay mode.

### $\Gamma(f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_{172}/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>1.18 \pm 0.07</math> OUR AVERAGE</b>		Error includes scale factor of 1.3.		
$1.26 \pm 0.06 \pm 0.06$	2.5k	LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$1.11 \pm 0.06 \pm 0.05$	1.3 k	AUBERT	08AE BABR	$384 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)/\Gamma(\eta\pi^-\pi^+\pi^-\nu_\tau (\text{ex. } K^0))$

$\Gamma_{172}/\Gamma_{160}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.69 \pm 0.01 \pm 0.05</math></b>	<sup>1</sup> AUBERT	08AE BABR	$384 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.55 $\pm 0.14$	BERGFELD	97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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<sup>1</sup> Not independent of AUBERT 08AE  $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)$  and  $B(\tau^- \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau (\text{ex. } K^0))$  values.

### $\Gamma(f_1(1285)\pi^-\nu_\tau \rightarrow 3\pi^- 2\pi^+\nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_{173}/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.52 \pm 0.04</math> OUR FIT</b>				
<b><math>0.520 \pm 0.031 \pm 0.037</math></b>	3.7k	LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(\pi(1300)^-\nu_\tau \rightarrow (\rho\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_{174}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.0 \times 10^{-4}</math></b>	90	ASNER	00	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(\pi(1300)^-\nu_\tau \rightarrow ((\pi\pi)_S\text{-wave}\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_{175}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.9 \times 10^{-4}</math></b>	90	ASNER	00	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$$\Gamma(h^- \omega \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{176}/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**2.40±0.08 OUR FIT**

• • • We use the following data for averages but not for fits. • • •

**1.65±0.3 ±0.2**      1513      ALBRECHT      88M ARG       $E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$

$$\Gamma(h^- \omega \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{177}/\Gamma = (\Gamma_{178} + \Gamma_{179})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.99±0.06 OUR FIT****1.92±0.07 OUR AVERAGE**

1.91±0.07±0.06      5803      BUSKULIC      97C ALEP      1991–1994 LEP runs

1.60±0.27±0.41      139      BARINGER      87 CLEO       $E_{\text{cm}}^{\text{ee}} = 10.5 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

1.95±0.07±0.11      2223      <sup>1</sup>BALEST      95C CLEO       $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

<sup>1</sup> Not independent of BALEST 95C  $B(\tau^- \rightarrow h^- \omega \nu_\tau)/B(\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau)$  value.

$$[\Gamma(\pi^- \omega \nu_\tau) + \Gamma(K^- \omega \nu_\tau)]/\Gamma(h^- h^- h^+ \pi^0 \nu_\tau (\text{ex. } K^0)) \quad (\Gamma_{178} + \Gamma_{179})/\Gamma_{74}$$

$$(\Gamma_{178} + \Gamma_{179})/\Gamma_{74} = (\Gamma_{178} + \Gamma_{179})/(\Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{178} + 0.892\Gamma_{179} + 0.0153\Gamma_{180})$$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**43.5±1.4 OUR FIT****45.3±1.9 OUR AVERAGE**

43.1±3.3      2350      <sup>1</sup>BUSKULIC      96 ALEP      LEP 1991–1993 data

46.4±1.6±1.7      2223      <sup>2</sup>BALEST      95C CLEO       $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

37 ±5 ±2      458      <sup>3</sup>ALBRECHT      91D ARG       $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$

<sup>1</sup> BUSKULIC 96 quote the fraction of  $\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau$  (ex.  $K^0$ ) decays which originate in a  $h^- \omega$  final state =  $0.383 \pm 0.029$ . We divide this by the  $\omega(782) \rightarrow \pi^+ \pi^- \pi^0$  branching fraction (0.888).

<sup>2</sup> BALEST 95C quote the fraction of  $\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau$  (ex.  $K^0$ ) decays which originate in a  $h^- \omega$  final state equals  $0.412 \pm 0.014 \pm 0.015$ . We divide this by the  $\omega(782) \rightarrow \pi^+ \pi^- \pi^0$  branching fraction (0.888).

<sup>3</sup> ALBRECHT 91D quote the fraction of  $\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau$  decays which originate in a  $\pi^- \omega$  final state equals  $0.33 \pm 0.04 \pm 0.02$ . We divide this by the  $\omega(782) \rightarrow \pi^+ \pi^- \pi^0$  branching fraction (0.888).

$$\Gamma(\pi^- \omega \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{178}/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID
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**1.95±0.06 OUR FIT**

$$\Gamma(K^- \omega \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{179}/\Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**4.1±0.9 OUR FIT**

**4.1±0.6±0.7**      500      ARMS      05 CLE3       $7.6 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(h^-\omega\pi^0\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{180}/\Gamma$			
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.41±0.04 OUR FIT</b>				
<b>0.43±0.06±0.05</b>	7283	BUSKULIC	97C ALEP	1991–1994 LEP runs

$\Gamma(h^-\omega\pi^0\nu_\tau)/\Gamma(h^-h^-h^+\geq 0 K_L^0\nu_\tau)$	$\Gamma_{180}/\Gamma_{62}$			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>(2.69 ±0.28 ) × 10<sup>-2</sup> OUR FIT</b>				

• • • We use the following data for averages but not for fits. • • •

<b>0.028±0.003±0.003</b>	430	<sup>1</sup> BORTOLETTO 93 CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$	
1 Not independent of BORTOLETTO 93 $\Gamma(\tau^- \rightarrow h^-\omega\pi^0\nu_\tau)/\Gamma(\tau^- \rightarrow h^-h^-h^+2\pi^0\nu_\tau(\text{ex. } K^0))$ value.				

$\Gamma(h^-\omega\pi^0\nu_\tau)/\Gamma(h^-h^-h^+2\pi^0\nu_\tau(\text{ex. } K^0))$	$\Gamma_{180}/\Gamma_{84}$			
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>82±8 OUR FIT</b>				

<b>81±6±6</b>		BORTOLETTO93 CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$	
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$\Gamma(h^-\omega 2\pi^0\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{181}/\Gamma$			
VALUE (units 10 <sup>-4</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.4 ±0.4 ±0.3</b>	53	ANASTASSOV 01 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				

$1.89^{+0.74}_{-0.67} \pm 0.40$       19      ANDERSON      97      CLEO      Repl. by ANASTASSOV 01

$\Gamma(\pi^-\omega 2\pi^0\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{182}/\Gamma$			
VALUE (units 10 <sup>-4</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.72±0.16 OUR FIT</b>				
<b>0.73±0.12±0.12</b>	1.1k	LEES	12x BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(h^-2\omega\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{183}/\Gamma$			
VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;5.4 × 10<sup>-7</sup></b>	90	AUBERT,B 06 BABR	$232 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	

$\Gamma(2h^-\omega\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{184}/\Gamma$			
VALUE (units 10 <sup>-4</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.2±0.2±0.1</b>	110	ANASTASSOV 01 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	

$\Gamma(2\pi^-\pi^+\omega\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$	$\Gamma_{185}/\Gamma$			
VALUE (units 10 <sup>-4</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.84±0.06 OUR FIT</b>				
<b>0.84±0.04±0.06</b>	2.4k	LEES	12x BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(e^- \gamma)/\Gamma_{\text{total}}$ 

Test of lepton family number conservation.

 $\Gamma_{186}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<3.3 \times 10^{-8}$	90	AUBERT	10B	BABR	$516 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$<5.6 \times 10^{-8}$	90	UNO	21	BELL	$988 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.2 \times 10^{-7}$	90	HAYASAKA	08	BELL	$535 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.1 \times 10^{-7}$	90	AUBERT	06C	BABR	$232 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.9 \times 10^{-7}$	90	HAYASAKA	05	BELL	$86.7 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.7 \times 10^{-6}$	90	EDWARDS	97	CLEO	
$<1.1 \times 10^{-4}$	90	ABREU	95U	DLPH	1990–1993 LEP runs
$<1.2 \times 10^{-4}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<2.0 \times 10^{-4}$	90	KEH	88	CBAL	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<6.4 \times 10^{-4}$	90	HAYES	82	MRK2	$E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

 $\Gamma(e^- \gamma\gamma)/\Gamma_{\text{total}}$  $\Gamma_{187}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<2.5 \times 10^{-4}$	90	<sup>1</sup> BRYMAN	21	RVUE	$516 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> BRYMAN 21 reinterprets the upper limit result on  $B(\tau^- \rightarrow e^- \gamma)$  and  $B(\tau^- \rightarrow \mu^- \gamma)$  by AUBERT 10B, estimating with a simulation the efficiency for this decay mode to be detected as the corresponding AUBERT 10B decay mode.

 $\Gamma(\mu^- \gamma)/\Gamma_{\text{total}}$  $\Gamma_{188}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 4.2 \times 10^{-8}$	90	UNO	21	BELL	$988 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 4.4 \times 10^{-8}$	90	AUBERT	10B	BABR	$516 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 4.5 \times 10^{-8}$	90	HAYASAKA	08	BELL	$535 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 6.8 \times 10^{-8}$	90	AUBERT,B	05A	BABR	$232 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 3.1 \times 10^{-7}$	90	ABE	04B	BELL	$86.3 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.1 \times 10^{-6}$	90	AHMED	00	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 3.0 \times 10^{-6}$	90	EDWARDS	97	CLEO	
$< 6.2 \times 10^{-5}$	90	ABREU	95U	DLPH	1990–1993 LEP runs
$< 0.42 \times 10^{-5}$	90	BEAN	93	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 3.4 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 55 \times 10^{-5}$	90	HAYES	82	MRK2	$E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

 $\Gamma(\mu^- \gamma\gamma)/\Gamma_{\text{total}}$  $\Gamma_{189}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<5.8 \times 10^{-4}$	90	<sup>1</sup> BRYMAN	21	RVUE	$516 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> BRYMAN 21 reinterprets the upper limit result on  $B(\tau^- \rightarrow e^- \gamma)$  and  $B(\tau^- \rightarrow \mu^- \gamma)$  by AUBERT 10B, estimating with a simulation the efficiency for this decay mode to be detected as the corresponding AUBERT 10B decay mode.

$\Gamma(e^- \pi^0)/\Gamma_{\text{total}}$  $\Gamma_{190}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
< <b>8.0 × 10<sup>-8</sup></b>	90	MIYAZAKI	07	BELL 401 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
< 1.3 × 10 <sup>-7</sup>	90	AUBERT	07I	BABR 339 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
< 1.9 × 10 <sup>-7</sup>	90	ENARI	05	BELL 154 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
< 3.7 × 10 <sup>-6</sup>	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
< 17 × 10 <sup>-5</sup>	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}}=10$ GeV
< 14 × 10 <sup>-5</sup>	90	KEH	88	CBAL $E_{\text{cm}}^{\text{ee}}=10$ GeV
< 210 × 10 <sup>-5</sup>	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}}=3.8\text{--}6.8$ GeV

 $\Gamma(\mu^- \pi^0)/\Gamma_{\text{total}}$  $\Gamma_{191}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
< <b>1.1 × 10<sup>-7</sup></b>	90	AUBERT	07I	BABR 339 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
< 1.2 × 10 <sup>-7</sup>	90	MIYAZAKI	07	BELL 401 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
< 4.1 × 10 <sup>-7</sup>	90	ENARI	05	BELL 154 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
< 4.0 × 10 <sup>-6</sup>	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
< 4.4 × 10 <sup>-5</sup>	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}}=10$ GeV
< 82 × 10 <sup>-5</sup>	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}}=3.8\text{--}6.8$ GeV

 $\Gamma(e^- K_S^0)/\Gamma_{\text{total}}$  $\Gamma_{192}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
< <b>2.6 × 10<sup>-8</sup></b>	90	MIYAZAKI	10A	BELL 671 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
< 3.3 × 10 <sup>-8</sup>	90	AUBERT	09D	BABR 469 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
< 5.6 × 10 <sup>-8</sup>	90	MIYAZAKI	06A	BELL 281 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
< 9.1 × 10 <sup>-7</sup>	90	CHEN	02C	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
< 1.3 × 10 <sup>-3</sup>	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}}=3.8\text{--}6.8$ GeV

 $\Gamma(\mu^- K_S^0)/\Gamma_{\text{total}}$  $\Gamma_{193}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
< <b>2.3 × 10<sup>-8</sup></b>	90	MIYAZAKI	10A	BELL 671 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
< 4.0 × 10 <sup>-8</sup>	90	AUBERT	09D	BABR 469 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
< 4.9 × 10 <sup>-8</sup>	90	MIYAZAKI	06A	BELL 281 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
< 9.5 × 10 <sup>-7</sup>	90	CHEN	02C	CLEO $E_{\text{cm}}^{\text{ee}}=10.6$ GeV
< 1.0 × 10 <sup>-3</sup>	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}}=3.8\text{--}6.8$ GeV

 $\Gamma(e^- \eta)/\Gamma_{\text{total}}$  $\Gamma_{194}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
< <b>9.2 × 10<sup>-8</sup></b>	90	MIYAZAKI	07	BELL 401 fb <sup>-1</sup> , $E_{\text{cm}}^{\text{ee}}=10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 1.6 \times 10^{-7}$	90	AUBERT	07I	BABR	$339 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.4 \times 10^{-7}$	90	ENARI	05	BELL	$154 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 8.2 \times 10^{-6}$	90	BONVICINI	97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 6.3 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 24 \times 10^{-5}$	90	KEH	88	CBAL	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$

### $\Gamma(\mu^- \eta)/\Gamma_{\text{total}}$

$\Gamma_{195}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 6.5 \times 10^{-8}$	90	MIYAZAKI	07	BELL $401 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 1.5 \times 10^{-7}$	90	AUBERT	07I	BABR	$339 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.5 \times 10^{-7}$	90	ENARI	05	BELL	$154 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 3.4 \times 10^{-7}$	90	ENARI	04	BELL	$84.3 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 9.6 \times 10^{-6}$	90	BONVICINI	97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 7.3 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$

### $\Gamma(e^- \rho^0)/\Gamma_{\text{total}}$

$\Gamma_{196}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.8 \times 10^{-8}$	90	MIYAZAKI	11	BELL $854 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 4.6 \times 10^{-8}$	90	AUBERT	09W	BABR	$451 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 6.3 \times 10^{-8}$	90	NISHIO	08	BELL	$543 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 6.5 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 4.2 \times 10^{-6}$	90	<sup>1</sup> BARTELTT	94	CLEO	Repl. by BLISS 98
$< 1.9 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 37 \times 10^{-5}$	90	HAYES	82	MRK2	$E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> BARTELTT 94 assume phase space decays.

### $\Gamma(\mu^- \rho^0)/\Gamma_{\text{total}}$

$\Gamma_{197}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.2 \times 10^{-8}$	90	MIYAZAKI	11	BELL $854 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 2.6 \times 10^{-8}$	90	AUBERT	09W	BABR	$451 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 6.8 \times 10^{-8}$	90	NISHIO	08	BELL	$543 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 6.3 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 5.7 \times 10^{-6}$	90	<sup>1</sup> BARTELTT	94	CLEO	Repl. by BLISS 98
$< 2.9 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 44 \times 10^{-5}$	90	HAYES	82	MRK2	$E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> BARTELTT 94 assume phase space decays.

$\Gamma(e^- \omega)/\Gamma_{\text{total}}$  $\Gamma_{198}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt;4.8 \times 10^{-8}</math></b>	90	MIYAZAKI	11	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$<1.1 \times 10^{-7}$	90	AUBERT	08K	BABR $384 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.8 \times 10^{-7}$	90	NISHIO	08	BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\mu^- \omega)/\Gamma_{\text{total}}$  $\Gamma_{199}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt;4.7 \times 10^{-8}</math></b>	90	MIYAZAKI	11	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$<1.0 \times 10^{-7}$	90	AUBERT	08K	BABR $384 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<8.9 \times 10^{-8}$	90	NISHIO	08	BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(e^- K^*(892)^0)/\Gamma_{\text{total}}$  $\Gamma_{200}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt;3.2 \times 10^{-8}</math></b>	90	MIYAZAKI	11	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$<5.9 \times 10^{-8}$	90	AUBERT	09W	BABR $451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.8 \times 10^{-8}$	90	NISHIO	08	BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.0 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<5.1 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<6.3 \times 10^{-6}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$<3.8 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$

<sup>1</sup> BARTEL 94 assume phase space decays. $\Gamma(\mu^- K^*(892)^0)/\Gamma_{\text{total}}$  $\Gamma_{201}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt;5.9 \times 10^{-8}</math></b>	90	NISHIO	08	BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$<7.2 \times 10^{-8}$	90	MIYAZAKI	11	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.7 \times 10^{-7}$	90	AUBERT	09W	BABR $451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.9 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.5 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<9.4 \times 10^{-6}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$<4.5 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$

<sup>1</sup> BARTEL 94 assume phase space decays. $\Gamma(e^- \bar{K}^*(892)^0)/\Gamma_{\text{total}}$  $\Gamma_{202}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt;3.4 \times 10^{-8}</math></b>	90	MIYAZAKI	11	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<4.6 \times 10^{-8}$	90	AUBERT	09W	BABR	$451 \text{ fb}^{-1}$	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.7 \times 10^{-8}$	90	NISHIO	08	BELL	$543 \text{ fb}^{-1}$	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<4.0 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1}$	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.4 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$<1.1 \times 10^{-5}$	90	<sup>1</sup> BARTEL	94	CLEO	Repl. by BLISS 98	

<sup>1</sup> BARTEL 94 assume phase space decays.

### $\Gamma(\mu^- \bar{K}^*(892)^0)/\Gamma_{\text{total}}$ $\Gamma_{203}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<7.0 \times 10^{-8}$	90	MIYAZAKI	11	BELL $854 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<7.3 \times 10^{-8}$	90	AUBERT	09W	BABR	$451 \text{ fb}^{-1}$	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.0 \times 10^{-7}$	90	NISHIO	08	BELL	$543 \text{ fb}^{-1}$	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<4.0 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1}$	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.5 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$<8.7 \times 10^{-6}$	90	<sup>1</sup> BARTEL	94	CLEO	Repl. by BLISS 98	

<sup>1</sup> BARTEL 94 assume phase space decays.

### $\Gamma(e^- \eta'(958))/\Gamma_{\text{total}}$ $\Gamma_{204}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.6 \times 10^{-7}$	90	MIYAZAKI	07	BELL $401 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 2.4 \times 10^{-7}$	90	AUBERT	07I	BABR $339 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<10. \times 10^{-7}$	90	ENARI	05	BELL $154 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(\mu^- \eta'(958))/\Gamma_{\text{total}}$ $\Gamma_{205}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.3 \times 10^{-7}$	90	MIYAZAKI	07	BELL $401 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<1.4 \times 10^{-7}$	90	AUBERT	07I	BABR $339 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<4.7 \times 10^{-7}$	90	ENARI	05	BELL $154 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(e^- f_0(980) \rightarrow e^- \pi^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{206}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.2 \times 10^{-8}$	90	MIYAZAKI	09	BELL $671 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(\mu^- f_0(980) \rightarrow \mu^- \pi^+ \pi^-)/\Gamma_{\text{total}}$ $\Gamma_{207}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.4 \times 10^{-8}$	90	MIYAZAKI	09	BELL $671 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(e^- \phi)/\Gamma_{\text{total}}$ $\Gamma_{208}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.1 \times 10^{-8}$	90	MIYAZAKI	11	BELL $854 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.1 \times 10^{-8}$	90	AUBERT	09W	BABR $451 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<7.3 \times 10^{-8}$	90	NISHIO	08	BELL	$543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.3 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<6.9 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(\mu^- \phi)/\Gamma_{\text{total}}$

$\Gamma_{209}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<8.4 \times 10^{-8}$	90	MIYAZAKI	11	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.9 \times 10^{-7}$	90	AUBERT	09W	BABR $451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.3 \times 10^{-7}$	90	NISHIO	08	BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.7 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.0 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(e^- e^+ e^-)/\Gamma_{\text{total}}$

$\Gamma_{210}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.7 \times 10^{-8}$	90	HAYASAKA	10	BELL $782 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 2.9 \times 10^{-8}$	90	LEES	10A	BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 3.6 \times 10^{-8}$	90	MIYAZAKI	08	BELL $535 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 4.3 \times 10^{-8}$	90	AUBERT	07BK	BABR $376 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	AUBERT	04J	BABR $91.5 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 3.5 \times 10^{-7}$	90	YUSA	04	BELL $87.1 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.9 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 0.33 \times 10^{-5}$	90	<sup>1</sup> BARTELTT	94	CLEO Repl. by BLISS 98
$< 1.3 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 2.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$
$< 40 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> BARTELTT 94 assume phase space decays.

### $\Gamma(e^- \mu^+ \mu^-)/\Gamma_{\text{total}}$

$\Gamma_{211}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.7 \times 10^{-8}$	90	HAYASAKA	10	BELL $782 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 3.2 \times 10^{-8}$	90	LEES	10A	BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 4.1 \times 10^{-8}$	90	MIYAZAKI	08	BELL $535 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 3.7 \times 10^{-8}$	90	AUBERT	07BK	BABR $376 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 3.3 \times 10^{-7}$	90	AUBERT	04J	BABR $91.5 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	YUSA	04	BELL $87.1 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.8 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 0.36 \times 10^{-5}$	90	<sup>1</sup> BARTELTT	94	CLEO Repl. by BLISS 98
$< 1.9 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 2.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$
$< 33 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(e^+ \mu^- \mu^-)/\Gamma_{\text{total}}$

Test of lepton family number conservation.

### $\Gamma_{212}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.7 \times 10^{-8}$	90	HAYASAKA	10	BELL $782 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$<2.6 \times 10^{-8}$	90	LEES	10A	BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.3 \times 10^{-8}$	90	MIYAZAKI	08	BELL $535 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<5.6 \times 10^{-8}$	90	AUBERT	07BK	BABR $376 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.3 \times 10^{-7}$	90	AUBERT	04J	BABR $91.5 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.0 \times 10^{-7}$	90	YUSA	04	BELL $87.1 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.5 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<0.35 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98
$<1.8 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<1.6 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(\mu^- e^+ e^-)/\Gamma_{\text{total}}$

Test of lepton family number conservation.

### $\Gamma_{213}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.8 \times 10^{-8}$	90	HAYASAKA	10	BELL $782 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$< 2.2 \times 10^{-8}$	90	LEES	10A	BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.7 \times 10^{-8}$	90	MIYAZAKI	08	BELL $535 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 8.0 \times 10^{-8}$	90	AUBERT	07BK	BABR $376 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.7 \times 10^{-7}$	90	AUBERT	04J	BABR $91.5 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.9 \times 10^{-7}$	90	YUSA	04	BELL $87.1 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.7 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 0.34 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98
$< 1.4 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 2.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$
$< 44 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(\mu^+ e^- e^-)/\Gamma_{\text{total}}$

Test of lepton family number conservation.

### $\Gamma_{214}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.5 \times 10^{-8}$	90	HAYASAKA	10	BELL $782 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$< 1.8 \times 10^{-8}$	90	LEES	10A	BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-8}$	90	MIYAZAKI	08	BELL $535 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 5.8 \times 10^{-8}$	90	AUBERT	07BK	BABR $376 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.1 \times 10^{-7}$	90	AUBERT	04J	BABR $91.5 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	YUSA	04	BELL $87.1 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.5 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 0.34 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98

$<1.4 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<1.6 \times 10^{-5}$	90	BOWCOCK	90	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(\mu^-\mu^+\mu^-)/\Gamma_{\text{total}}$

### $\Gamma_{215}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 2.1 \times 10^{-8}$	90	HAYASAKA	10	BELL	$782 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 8.0 \times 10^{-8}$	90	SIRUNYAN	21D	CMS	$33.2 \text{ fb}^{-1}, pp \text{ at } 13 \text{ TeV}$
$< 3.8 \times 10^{-7}$	90	AAD	16BA	ATLS	$20.3 \text{ fb}^{-1}, pp \text{ at } 8 \text{ TeV}$
$< 4.6 \times 10^{-8}$	90	AAIJ	15AI	LHCb	$3.0 \text{ fb}^{-1}, pp \text{ at } 7, 8 \text{ TeV}$
$< 8.0 \times 10^{-8}$	90	<sup>1</sup> AAIJ	13AH	LHCb	$1.0 \text{ fb}^{-1}, pp \text{ at } 7 \text{ TeV}$
$< 3.3 \times 10^{-8}$	90	LEES	10A	BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 3.2 \times 10^{-8}$	90	MIYAZAKI	08	BELL	$535 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 5.3 \times 10^{-8}$	90	AUBERT	07BK	BABR	$376 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.9 \times 10^{-7}$	90	AUBERT	04J	BABR	$91.5 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	YUSA	04	BELL	$87.1 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.9 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 0.43 \times 10^{-5}$	90	<sup>2</sup> BARTELT	94	CLEO	Repl. by BLISS 98
$< 1.9 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 1.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$
$< 49 \times 10^{-5}$	90	HAYES	82	MRK2	$E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> Repl. by AAIJ 15AI.

<sup>2</sup> BARTELT 94 assume phase space decays.

### $\Gamma(e^-\pi^+\pi^-)/\Gamma_{\text{total}}$

### $\Gamma_{216}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 2.3 \times 10^{-8}$	90	MIYAZAKI	13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 4.4 \times 10^{-8}$	90	MIYAZAKI	10	BELL	Repl. by MIYAZAKI 13
$< 7.3 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.2 \times 10^{-7}$	90	AUBERT,BE	05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.2 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 4.4 \times 10^{-6}$	90	<sup>1</sup> BARTELT	94	CLEO	Repl. by BLISS 98
$< 2.7 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 6.0 \times 10^{-5}$	90	BOWCOCK	90	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

### $\Gamma(e^+\pi^-\pi^-)/\Gamma_{\text{total}}$

### $\Gamma_{217}/\Gamma$

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 2.0 \times 10^{-8}$	90	MIYAZAKI	13	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 8.8 \times 10^{-8}$	90	MIYAZAKI	10	BELL	Repl. by MIYAZAKI 13
$< 2.0 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$<2.7 \times 10^{-7}$	90	AUBERT,BE	05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.9 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<4.4 \times 10^{-6}$	90	<sup>1</sup> BARTELTT	94	CLEO	Repl. by BLISS 98
$<1.8 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<1.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTELTT 94 assume phase space decays.

### $\Gamma(\mu^- \pi^+ \pi^-)/\Gamma_{\text{total}}$

$\Gamma_{218}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;2.1 \times 10^{-8}</math></b>	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<3.3 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<4.8 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.9 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<8.2 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.4 \times 10^{-6}$	90	<sup>1</sup> BARTELTT	94	CLEO Repl. by BLISS 98
$<3.6 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<3.9 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTELTT 94 assume phase space decays.

### $\Gamma(\mu^+ \pi^- \pi^-)/\Gamma_{\text{total}}$

$\Gamma_{219}/\Gamma$

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.9 \times 10^{-8}</math></b>	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<3.7 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<3.4 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7 \times 10^{-8}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.4 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<6.9 \times 10^{-6}$	90	<sup>1</sup> BARTELTT	94	CLEO Repl. by BLISS 98
$<6.3 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<3.9 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTELTT 94 assume phase space decays.

### $\Gamma(e^- \pi^+ K^-)/\Gamma_{\text{total}}$

$\Gamma_{220}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.7 \times 10^{-8}</math></b>	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<5.8 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<7.2 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.2 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<6.4 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.7 \times 10^{-6}$	90	<sup>1</sup> BARTELTT	94	CLEO Repl. by BLISS 98
$<2.9 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<5.8 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTELTT 94 assume phase space decays.

$\Gamma(e^- \pi^- K^+)/\Gamma_{\text{total}}$  $\Gamma_{221}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.1 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$<5.2 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<1.6 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.7 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.8 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<4.6 \times 10^{-6}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$<5.8 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTEL 94 assume phase space decays. $\Gamma(e^+ \pi^- K^-)/\Gamma_{\text{total}}$  $\Gamma_{222}/\Gamma$ 

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.2 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$<6.7 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<1.9 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.8 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.1 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<4.5 \times 10^{-6}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$<2.0 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<4.9 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTEL 94 assume phase space decays. $\Gamma(e^- K_S^0 K_S^0)/\Gamma_{\text{total}}$  $\Gamma_{223}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<7.1 \times 10^{-8}$	90	MIYAZAKI	10A	BELL $671 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$<2.2 \times 10^{-6}$	90	CHEN	02C	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(e^- K^+ K^-)/\Gamma_{\text{total}}$  $\Gamma_{224}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.4 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$<5.4 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<3.0 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.4 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<6.0 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(e^+ K^- K^-)/\Gamma_{\text{total}}$  $\Gamma_{225}/\Gamma$ 

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.3 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<6.0 \times 10^{-8}$	90	MIYAZAKI	10	BELL	Repl. by MIYAZAKI 13
$<3.1 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.5 \times 10^{-7}$	90	AUBERT,BE	05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.8 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(\mu^- \pi^+ K^-)/\Gamma_{\text{total}}$

### $\Gamma_{226}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 8.6 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.6 \times 10^{-7}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$< 2.7 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.6 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 7.5 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 8.7 \times 10^{-6}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$< 11 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 7.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTEL 94 assume phase space decays.

### $\Gamma(\mu^- \pi^- K^+)/\Gamma_{\text{total}}$

### $\Gamma_{227}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 4.5 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.0 \times 10^{-7}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$< 7.3 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 3.2 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 7.4 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.5 \times 10^{-5}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$< 7.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTEL 94 assume phase space decays.

### $\Gamma(\mu^+ \pi^- K^-)/\Gamma_{\text{total}}$

### $\Gamma_{228}/\Gamma$

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 4.8 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 9.4 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$< 2.9 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.2 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 7.0 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-5}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$< 5.8 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 4.0 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTEL 94 assume phase space decays.

$\Gamma(\mu^- K_S^0 K_S^0)/\Gamma_{\text{total}}$  $\Gamma_{229}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;8.0 × 10<sup>-8</sup></b>	90	MIYAZAKI	10A BELL	$671 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<3.4 \times 10^{-6}$	90	CHEN	02C CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\mu^- K^+ K^-)/\Gamma_{\text{total}}$  $\Gamma_{230}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt; 4.4 × 10<sup>-8</sup></b>	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$< 6.8 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$< 8.0 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.5 \times 10^{-7}$	90	AUBERT,BE	05D BABR	$221 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 15 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\mu^+ K^- K^-)/\Gamma_{\text{total}}$  $\Gamma_{231}/\Gamma$ 

Test of lepton number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;4.7 × 10<sup>-8</sup></b>	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$< 9.6 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$< 4.4 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 4.8 \times 10^{-7}$	90	AUBERT,BE	05D BABR	$221 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 6.0 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(e^- \pi^0 \pi^0)/\Gamma_{\text{total}}$  $\Gamma_{232}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;6.5 × 10<sup>-6</sup></b>	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\mu^- \pi^0 \pi^0)/\Gamma_{\text{total}}$  $\Gamma_{233}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;14 × 10<sup>-6</sup></b>	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(e^- \eta\eta)/\Gamma_{\text{total}}$  $\Gamma_{234}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;35 × 10<sup>-6</sup></b>	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\mu^- \eta\eta)/\Gamma_{\text{total}}$  $\Gamma_{235}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;60 × 10<sup>-6</sup></b>	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(e^- \pi^0 \eta)/\Gamma_{\text{total}}$ 

Test of lepton family number conservation.

 $\Gamma_{236}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<24 \times 10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\mu^- \pi^0 \eta)/\Gamma_{\text{total}}$ 

Test of lepton family number conservation.

 $\Gamma_{237}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<22 \times 10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(p e^- e^-)/\Gamma_{\text{total}}$  $\Gamma_{238}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<3.0 \times 10^{-8}$	90	SAHOO	20	BELL    921 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\bar{p} e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{239}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<3.0 \times 10^{-8}$	90	SAHOO	20	BELL    921 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\bar{p} e^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{240}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<2.0 \times 10^{-8}$	90	SAHOO	20	BELL    921 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\bar{p} e^- \mu^+)/\Gamma_{\text{total}}$  $\Gamma_{241}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<1.8 \times 10^{-8}$	90	SAHOO	20	BELL    921 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(p \mu^- \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{242}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<4.0 \times 10^{-8}$	90	SAHOO	20	BELL    921 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $<4.4 \times 10^{-7}$     90    AAIJ    13AH LHCb    1.0 fb<sup>-1</sup>, pp at 7 TeV $\Gamma(\bar{p} \mu^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{243}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<1.8 \times 10^{-8}$	90	SAHOO	20	BELL    921 fb <sup>-1</sup> $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $<3.3 \times 10^{-7}$     90    AAIJ    13AH LHCb    1.0 fb<sup>-1</sup>, pp at 7 TeV $\Gamma(\bar{p} \gamma)/\Gamma_{\text{total}}$  $\Gamma_{244}/\Gamma$ 

Test of lepton number and baryon number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$< 3.5 \times 10^{-6}$	90	GODANG	99	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

 $<29 \times 10^{-5}$     90    ALBRECHT    92K ARG     $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$  $\Gamma(\bar{p} \pi^0)/\Gamma_{\text{total}}$  $\Gamma_{245}/\Gamma$ 

Test of lepton number and baryon number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<15 \times 10^{-6}$	90	GODANG	99	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<66 \times 10^{-5}$  90 ALBRECHT 92K ARG  $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$

### $\Gamma(\bar{p}2\pi^0)/\Gamma_{\text{total}}$

$\Gamma_{246}/\Gamma$

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<33 \times 10^{-6}$	90	GODANG	99	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(\bar{p}\eta)/\Gamma_{\text{total}}$

$\Gamma_{247}/\Gamma$

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 8.9 \times 10^{-6}$	90	GODANG	99	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<130 \times 10^{-5}$  90 ALBRECHT 92K ARG  $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$

### $\Gamma(\bar{p}\pi^0\eta)/\Gamma_{\text{total}}$

$\Gamma_{248}/\Gamma$

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<27 \times 10^{-6}$	90	GODANG	99	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(\Lambda\pi^-)/\Gamma_{\text{total}}$

$\Gamma_{249}/\Gamma$

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<0.72 \times 10^{-7}$	90	MIYAZAKI	06	BELL $154 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(\bar{\Lambda}\pi^-)/\Gamma_{\text{total}}$

$\Gamma_{250}/\Gamma$

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.4 \times 10^{-7}$	90	MIYAZAKI	06	BELL $154 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(e^- \text{ light boson})/\Gamma(e^- \bar{\nu}_e \nu_\tau)$

$\Gamma_{251}/\Gamma_5$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<0.015$	95	1 ALBRECHT	95G	ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<0.008$  95 2 BRYMAN 21 RVUE

$<0.018$  95 3 ALBRECHT 90E ARG  $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$

$<0.040$  95 4 BALTRUSAITIS..85 MRK3  $E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$

<sup>1</sup> ALBRECHT 95G limit holds for bosons with mass  $< 0.4 \text{ GeV}$ . The limit rises to 0.036 for a mass of 1.0 GeV, then falls to 0.006 at the upper mass limit of 1.6 GeV.

<sup>2</sup> BRYMAN 21 reports indirect limits obtained from the consistency of the world averages of  $B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  and  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$  with their Standard Model predictions, without a simulation of the efficiency as a function of the  $X$  mass for the searched decay modes to be detected as the corresponding Standard Model decay modes.

<sup>3</sup> ALBRECHT 90E limit applies for spinless boson with mass  $< 100 \text{ MeV}$ , and rises to 0.050 for mass = 500 MeV.

<sup>4</sup> BALTRUSAITIS 85 limit applies for spinless boson with mass  $< 100 \text{ MeV}$ .

$\Gamma(\mu^- \text{ light boson})/\Gamma(e^- \bar{\nu}_e \nu_\tau)$  $\Gamma_{252}/\Gamma_5$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.026</b>	95	<sup>1</sup> ALBRECHT	95G ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
<0.011	95	<sup>2</sup> BRYMAN	21 RVUE	
<0.033	95	<sup>3</sup> ALBRECHT	90E ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
<0.125	95	<sup>4</sup> BALTRUSAITIS	..85 MRK3	$E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$

<sup>1</sup> ALBRECHT 95G limit holds for bosons with mass < 1.3 GeV. The limit rises to 0.034 for a mass of 1.4 GeV, then falls to 0.003 at the upper mass limit of 1.6 GeV.

<sup>2</sup> BRYMAN 21 reports indirect limits obtained from the consistency of the world averages of  $B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  and  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$  with their Standard Model predictions, without a simulation of the efficiency as a function of the  $X$  mass for the searched decay modes to be detected as the corresponding Standard Model decay modes.

<sup>3</sup> ALBRECHT 90E limit applies for spinless boson with mass < 100 MeV, and rises to 0.071 for mass = 500 MeV.

<sup>4</sup> BALTRUSAITIS 85 limit applies for spinless boson with mass < 100 MeV.

 **$\tau$ -DECAY PARAMETERS**

See the related review(s):

[τ-Lepton Decay Parameters](#) **$\rho(e \text{ or } \mu)$  PARAMETER**(V-A) theory predicts  $\rho = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.745±0.008 OUR FIT</b>				
<b>0.749±0.008 OUR AVERAGE</b>				
0.742±0.014±0.006	81k	HEISTER	01E ALEP	1991–1995 LEP runs
0.775±0.023±0.020	36k	ABREU	00L DLPH	1992–1995 runs
0.781±0.028±0.018	46k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.762±0.035	54k	ACCIARRI	98R L3	1991–1995 LEP runs
0.731±0.031		<sup>1</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$
0.72 ± 0.09 ± 0.03		<sup>2</sup> ABE	970 SLD	1993–1995 SLC runs
0.747±0.010±0.006	55k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
0.79 ± 0.10 ± 0.10	3732	FORD	87B MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
0.71 ± 0.09 ± 0.03	1426	BEHRENDS	85 CLEO	$e^+ e^-$ near $\gamma(4S)$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
0.735±0.013±0.008	31k	AMMAR	97B CLEO	Repl. by ALEXANDER 97F
0.794±0.039±0.031	18k	ACCIARRI	96H L3	Repl. by ACCIARRI 98R
0.732±0.034±0.020	8.2k	<sup>3</sup> ALBRECHT	95 ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$
0.738±0.038		<sup>4</sup> ALBRECHT	95C ARG	Repl. by ALBRECHT 98
0.751±0.039±0.022		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
0.742±0.035±0.020	8000	ALBRECHT	90E ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$

<sup>1</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

<sup>2</sup> ABE 970 assume  $\eta = 0$  in their fit. Letting  $\eta$  vary in the fit gives a  $\rho$  value of  $0.69 \pm 0.13 \pm 0.05$ .

<sup>3</sup> Value is from a simultaneous fit for the  $\rho$  and  $\eta$  decay parameters to the lepton energy spectrum. Not independent of ALBRECHT 90E  $\rho(e \text{ or } \mu)$  value which assumes  $\eta = 0$ . Result is strongly correlated with ALBRECHT 95C.

<sup>4</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E.

## $\rho(e)$ PARAMETER

( $V-A$ ) theory predicts  $\rho = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.747±0.010 OUR FIT</b>				
<b>0.744±0.010 OUR AVERAGE</b>				
0.747±0.019±0.014	44k	HEISTER 01E	ALEP	1991–1995 LEP runs
0.744±0.036±0.037	17k	ABREU 00L	DLPH	1992–1995 runs
0.779±0.047±0.029	25k	ACKERSTAFF 99D	OPAL	1990–1995 LEP runs
0.68 ± 0.04 ± 0.07		<sup>1</sup> ALBRECHT 98	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
0.71 ± 0.14 ± 0.05		ABE 970	SLD	1993–1995 SLC runs
0.747±0.012±0.004	34k	ALEXANDER 97F	CLEO	$E_{cm}^{ee} = 10.6 \text{ GeV}$
0.735±0.036±0.020	4.7k	<sup>2</sup> ALBRECHT 95	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
0.79 ± 0.08 ± 0.06	3230	<sup>3</sup> ALBRECHT 93G	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6 \text{ GeV}$
0.64 ± 0.06 ± 0.07	2753	JANSSEN 89	CBAL	$E_{cm}^{ee} = 9.4\text{--}10.6 \text{ GeV}$
0.62 ± 0.17 ± 0.14	1823	FORD 87B	MAC	$E_{cm}^{ee} = 29 \text{ GeV}$
0.60 ± 0.13	699	BEHREND 85	CLEO	$e^+ e^- \text{ near } \gamma(4S)$
0.72 ± 0.10 ± 0.11	594	BACINO 79B	DLCO	$E_{cm}^{ee} = 3.5\text{--}7.4 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.732±0.014±0.009	19k	AMMAR 97B	CLEO	Repl. by ALEXANDER 97F
0.793±0.050±0.025		BUSKULIC 95D	ALEP	Repl. by HEISTER 01E
0.747±0.045±0.028	5106	ALBRECHT 90E	ARG	Repl. by ALBRECHT 95

<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

<sup>2</sup> ALBRECHT 95 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(h^+ h^- (\pi^0) \bar{\nu}_\tau)$  and their charged conjugates.

<sup>3</sup> ALBRECHT 93G use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\mu^- \bar{\nu}_\mu \nu_\tau)(e^+ \nu_e \bar{\nu}_\tau)$  and their charged conjugates.

## $\rho(\mu)$ PARAMETER

( $V-A$ ) theory predicts  $\rho = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.763±0.020 OUR FIT</b>				
<b>0.770±0.022 OUR AVERAGE</b>				
0.776±0.045±0.019	46k	HEISTER 01E	ALEP	1991–1995 LEP runs
0.999±0.098±0.045	22k	ABREU 00L	DLPH	1992–1995 runs
0.777±0.044±0.016	27k	ACKERSTAFF 99D	OPAL	1990–1995 LEP runs
0.69 ± 0.06 ± 0.06		<sup>1</sup> ALBRECHT 98	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
0.54 ± 0.28 ± 0.14		ABE 970	SLD	1993–1995 SLC runs
0.750±0.017±0.045	22k	ALEXANDER 97F	CLEO	$E_{cm}^{ee} = 10.6 \text{ GeV}$
0.76 ± 0.07 ± 0.08	3230	ALBRECHT 93G	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6 \text{ GeV}$
0.734±0.055±0.027	3041	ALBRECHT 90E	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6 \text{ GeV}$
0.89 ± 0.14 ± 0.08	1909	FORD 87B	MAC	$E_{cm}^{ee} = 29 \text{ GeV}$
0.81 ± 0.13	727	BEHREND 85	CLEO	$e^+ e^- \text{ near } \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.747 \pm 0.048 \pm 0.044$	13k	AMMAR	97B	CLEO	Repl. by ALEXANDER 97F
$0.693 \pm 0.057 \pm 0.028$		BUSKULIC	95D	ALEP	Repl. by HEISTER 01E

<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

## $\xi(e \text{ or } \mu)$ PARAMETER

(V-A) theory predicts  $\xi = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.985 ± 0.030 OUR FIT</b>				
<b>0.981 ± 0.031 OUR AVERAGE</b>				
$0.986 \pm 0.068 \pm 0.031$	81k	HEISTER	01E	ALEP 1991–1995 LEP runs
$0.929 \pm 0.070 \pm 0.030$	36k	ABREU	00L	DLPH 1992–1995 runs
$0.98 \pm 0.22 \pm 0.10$	46k	ACKERSTAFF	99D	OPAL 1990–1995 LEP runs
$0.70 \pm 0.16$	54k	ACCIARRI	98R	L3 1991–1995 LEP runs
$1.03 \pm 0.11$		<sup>1</sup> ALBRECHT	98	ARG $E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
$1.05 \pm 0.35 \pm 0.04$		<sup>2</sup> ABE	970	SLD 1993–1995 SLC runs
$1.007 \pm 0.040 \pm 0.015$	55k	ALEXANDER	97F	CLEO $E_{cm}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$0.94 \pm 0.21 \pm 0.07$				
$0.97 \pm 0.14$				
$1.18 \pm 0.15 \pm 0.16$				
$0.90 \pm 0.15 \pm 0.10$	3230	<sup>3</sup> ALBRECHT	96H	L3 Repl. by ACCIARRI 98R
		<sup>3</sup> ALBRECHT	95C	ARG Repl. by ALBRECHT 98
		BUSKULIC	95D	ALEP Repl. by HEISTER 01E
		<sup>4</sup> ALBRECHT	93G	ARG $E_{cm}^{ee} = 9.4\text{--}10.6 \text{ GeV}$

<sup>1</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

<sup>2</sup> ABE 970 assume  $\eta = 0$  in their fit. Letting  $\eta$  vary in the fit gives a  $\xi$  value of  $1.02 \pm 0.36 \pm 0.05$ .

<sup>3</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 95C uses events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(h^+ h^- h^+ \bar{\nu}_\tau)$  and their charged conjugates.

<sup>4</sup> ALBRECHT 93G measurement determines  $|\xi|$  for the case  $\xi(e) = \xi(\mu)$ , but the authors point out that other LEP experiments determine the sign to be positive.

## $\xi(e)$ PARAMETER

(V-A) theory predicts  $\xi = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.994 ± 0.040 OUR FIT</b>				
<b>1.00 ± 0.04 OUR AVERAGE</b>				
$1.011 \pm 0.094 \pm 0.038$	44k	HEISTER	01E	ALEP 1991–1995 LEP runs
$1.01 \pm 0.12 \pm 0.05$	17k	ABREU	00L	DLPH 1992–1995 runs
$1.13 \pm 0.39 \pm 0.14$	25k	ACKERSTAFF	99D	OPAL 1990–1995 LEP runs
$1.11 \pm 0.20 \pm 0.08$		<sup>1</sup> ALBRECHT	98	ARG $E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
$1.16 \pm 0.52 \pm 0.06$		ABE	970	SLD 1993–1995 SLC runs
$0.979 \pm 0.048 \pm 0.016$	34k	ALEXANDER	97F	CLEO $E_{cm}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.03 \pm 0.23 \pm 0.09$	BUSKULIC	95D	ALEP	Repl. by HEISTER 01E
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<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

**$\xi(\mu)$  PARAMETER**(V-A) theory predicts  $\xi = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.030±0.059 OUR FIT</b>				
<b>1.06 ±0.06 OUR AVERAGE</b>				
1.030±0.120±0.050	46k	HEISTER 01E	ALEP	1991–1995 LEP runs
1.16 ±0.19 ±0.06	22k	ABREU 00L	DLPH	1992–1995 runs
0.79 ±0.41 ±0.09	27k	ACKERSTAFF 99D	OPAL	1990–1995 LEP runs
1.26 ±0.27 ±0.14		<sup>1</sup> ALBRECHT 98	ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$
0.75 ±0.50 ±0.14		ABE 970	SLD	1993–1995 SLC runs
1.054±0.069±0.047	22k	ALEXANDER 97F	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.23 ±0.22 ±0.10		BUSKULIC 95D	ALEP	Repl. by HEISTER 01E

<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

 **$\eta(e \text{ or } \mu)$  PARAMETER**(V-A) theory predicts  $\eta = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.013±0.020 OUR FIT</b>				
<b>0.015±0.021 OUR AVERAGE</b>				
0.012±0.026±0.004	81k	HEISTER 01E	ALEP	1991–1995 LEP runs
-0.005±0.036±0.037		ABREU 00L	DLPH	1992–1995 runs
0.027±0.055±0.005	46k	ACKERSTAFF 99D	OPAL	1990–1995 LEP runs
0.27 ±0.14	54k	ACCIARRI 98R	L3	1991–1995 LEP runs
-0.13 ±0.47 ±0.15		ABE 970	SLD	1993–1995 SLC runs
-0.015±0.061±0.062	31k	AMMAR 97B	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
0.03 ±0.18 ±0.12	8.2k	ALBRECHT 95	ARG	$E_{\text{cm}}^{\text{ee}} = 9.5\text{--}10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.25 ±0.17 ±0.11	18k	ACCIARRI 96H	L3	Repl. by ACCIARRI 98R
-0.04 ±0.15 ±0.11		BUSKULIC 95D	ALEP	Repl. by HEISTER 01E

 **$\eta(\mu)$  PARAMETER**(V-A) theory predicts  $\eta = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.094±0.073 OUR FIT</b>				
<b>0.17 ±0.15 OUR AVERAGE</b>				
Error includes scale factor of 1.2.				
0.160±0.150±0.060	46k	HEISTER 01E	ALEP	1991–1995 LEP runs
0.72 ±0.32 ±0.15		ABREU 00L	DLPH	1992–1995 runs
-0.59 ±0.82 ±0.45		<sup>1</sup> ABE 970	SLD	1993–1995 SLC runs
0.010±0.149±0.171	13k	<sup>2</sup> AMMAR 97B	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.010±0.065±0.001	27k	<sup>3</sup> ACKERSTAFF 99D	OPAL	1990–1995 LEP runs
-0.24 ±0.23 ±0.18		BUSKULIC 95D	ALEP	Repl. by HEISTER 01E

<sup>1</sup> Highly correlated (corr. = 0.92) with ABE 970  $\rho(\mu)$  measurement.

<sup>2</sup> Highly correlated (corr. = 0.949) with AMMAR 97B  $\rho(\mu)$  value.

<sup>3</sup> ACKERSTAFF 99D result is dominated by a constraint on  $\eta$  from the OPAL measurements of the  $\tau$  lifetime and  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$  assuming lepton universality for the total coupling strength.

**( $\delta\xi$ )(e or  $\mu$ ) PARAMETER** $(V-A)$  theory predicts  $(\delta\xi) = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.746±0.021 OUR FIT</b>				
<b>0.744±0.022 OUR AVERAGE</b>				
0.776±0.045±0.024	81k	HEISTER 01E	ALEP	1991–1995 LEP runs
0.779±0.070±0.028	36k	ABREU 00L	DLPH	1992–1995 runs
0.65 ± 0.14 ± 0.07	46k	ACKERSTAFF 99D	OPAL	1990–1995 LEP runs
0.70 ± 0.11	54k	ACCIARRI 98R	L3	1991–1995 LEP runs
0.63 ± 0.09		<sup>1</sup> ALBRECHT 98	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
0.88 ± 0.27 ± 0.04		<sup>2</sup> ABE 970	SLD	1993–1995 SLC runs
0.745±0.026±0.009	55k	ALEXANDER 97F	CLEO	$E_{cm}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.81 ± 0.14 ± 0.06	18k	ACCIARRI 96H	L3	Repl. by ACCIARRI 98R
0.65 ± 0.12		<sup>3</sup> ALBRECHT 95C	ARG	Repl. by ALBRECHT 98
0.88 ± 0.11 ± 0.07		BUSKULIC 95D	ALEP	Repl. by HEISTER 01E

<sup>1</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

<sup>2</sup> ABE 970 assume  $\eta = 0$  in their fit. Letting  $\eta$  vary in the fit gives a  $(\delta\xi)$  value of  $0.87 \pm 0.27 \pm 0.04$ .

<sup>3</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 95C uses events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(h^+ h^- h^+ \bar{\nu}_\tau)$  and their charged conjugates.

**( $\delta\xi$ )(e) PARAMETER** $(V-A)$  theory predicts  $(\delta\xi) = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.734±0.028 OUR FIT</b>				
<b>0.731±0.029 OUR AVERAGE</b>				
0.778±0.066±0.024	44k	HEISTER 01E	ALEP	1991–1995 LEP runs
0.85 ± 0.12 ± 0.04	17k	ABREU 00L	DLPH	1992–1995 runs
0.72 ± 0.31 ± 0.14	25k	ACKERSTAFF 99D	OPAL	1990–1995 LEP runs
0.56 ± 0.14 ± 0.06		<sup>1</sup> ALBRECHT 98	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
0.85 ± 0.43 ± 0.08		ABE 970	SLD	1993–1995 SLC runs
0.720±0.032±0.010	34k	ALEXANDER 97F	CLEO	$E_{cm}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.11 ± 0.17 ± 0.07		BUSKULIC 95D	ALEP	Repl. by HEISTER 01E

<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

**( $\delta\xi$ )( $\mu$ ) PARAMETER** $(V-A)$  theory predicts  $(\delta\xi) = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.778±0.037 OUR FIT</b>				
<b>0.79 ± 0.04 OUR AVERAGE</b>				
0.786±0.066±0.028	46k	HEISTER 01E	ALEP	1991–1995 LEP runs
0.86 ± 0.13 ± 0.04	22k	ABREU 00L	DLPH	1992–1995 runs
0.63 ± 0.23 ± 0.05	27k	ACKERSTAFF 99D	OPAL	1990–1995 LEP runs
0.73 ± 0.18 ± 0.10		<sup>1</sup> ALBRECHT 98	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
0.82 ± 0.32 ± 0.07		ABE 970	SLD	1993–1995 SLC runs

$0.786 \pm 0.041 \pm 0.032$	22k	ALEXANDER 97F	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$0.71 \pm 0.14 \pm 0.06$		BUSKULIC 95D	ALEP	Repl. by HEISTER 01E
<sup>1</sup> ALBRECHT 98 use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.				

### $\xi(\pi)$ PARAMETER

(V-A) theory predicts  $\xi(\pi) = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.993 ± 0.022 OUR FIT</b>				
<b>0.994 ± 0.023 OUR AVERAGE</b>				
$0.994 \pm 0.020 \pm 0.014$	27k	HEISTER 01E	ALEP	1991–1995 LEP runs
$0.81 \pm 0.17 \pm 0.02$		ABE 970	SLD	1993–1995 SLC runs
$1.03 \pm 0.06 \pm 0.04$	2.0k	COAN 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$0.987 \pm 0.057 \pm 0.027$		BUSKULIC 95D	ALEP	Repl. by HEISTER 01E
$0.95 \pm 0.11 \pm 0.05$		<sup>1</sup> BUSKULIC 94D	ALEP	1990+1991 LEP run

<sup>1</sup> Superseded by BUSKULIC 95D.

### $\xi(\rho)$ PARAMETER

(V-A) theory predicts  $\xi(\rho) = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.994 ± 0.008 OUR FIT</b>				
<b>0.994 ± 0.009 OUR AVERAGE</b>				
$0.987 \pm 0.012 \pm 0.011$	59k	HEISTER 01E	ALEP	1991–1995 LEP runs
$0.99 \pm 0.12 \pm 0.04$		ABE 970	SLD	1993–1995 SLC runs
$0.995 \pm 0.010 \pm 0.003$	66k	ALEXANDER 97F	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$1.022 \pm 0.028 \pm 0.030$	1.7k	<sup>1</sup> ALBRECHT 94E	ARG	$E_{\text{cm}}^{\text{ee}} = 9.4–10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$1.045 \pm 0.058 \pm 0.032$		BUSKULIC 95D	ALEP	Repl. by HEISTER 01E
$1.03 \pm 0.11 \pm 0.05$		<sup>2</sup> BUSKULIC 94D	ALEP	1990+1991 LEP run

<sup>1</sup> ALBRECHT 94E measure the square of this quantity and use the sign determined by ALBRECHT 90I to obtain the quoted result.

<sup>2</sup> Superseded by BUSKULIC 95D.

### $\xi(a_1)$ PARAMETER

(V-A) theory predicts  $\xi(a_1) = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.001 ± 0.027 OUR FIT</b>				
<b>1.002 ± 0.028 OUR AVERAGE</b>				
$1.000 \pm 0.016 \pm 0.024$	35k	<sup>1</sup> HEISTER 01E	ALEP	1991–1995 LEP runs
$1.02 \pm 0.13 \pm 0.03$	17.2k	ASNER 00	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$1.29 \pm 0.26 \pm 0.11$	7.4k	<sup>2</sup> ACKERSTAFF 97R	OPAL	1992–1994 LEP runs
$0.85 \pm 0.15 \pm 0.05$		ALBRECHT 95C	ARG	$E_{\text{cm}}^{\text{ee}} = 9.5–10.6 \text{ GeV}$
$1.25 \pm 0.23 \pm 0.15$	7.5k	ALBRECHT 93C	ARG	$E_{\text{cm}}^{\text{ee}} = 9.4–10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$1.08 \pm 0.46 \pm 0.14$	2.6k	<sup>3</sup> AKERS 95P	OPAL	Repl. by ACKER-STAFF 97R
$0.937 \pm 0.116 \pm 0.064$		BUSKULIC 95D	ALEP	Repl. by HEISTER 01E

- <sup>1</sup> HEISTER 01E quote  $1.000 \pm 0.016 \pm 0.013 \pm 0.020$  where the errors are statistical, systematic, and an uncertainty due to the final state model. We combine the systematic error and model uncertainty.
- <sup>2</sup> ACKERSTAFF 97R obtain this result with a model independent fit to the hadronic structure functions. Fitting with the model of Kuhn and Santamaria (ZPHY **C48**, 445 (1990)) gives  $0.87 \pm 0.16 \pm 0.04$ , and with the model of Isgur *et al.* (PR **D39**, 1357 (1989)) they obtain  $1.20 \pm 0.21 \pm 0.14$ .
- <sup>3</sup> AKERS 95P obtain this result with a model independent fit to the hadronic structure functions. Fitting with the model of Kuhn and Santamaria (ZPHY **C48**, 445 (1990)) gives  $0.87 \pm 0.27^{+0.05}_{-0.06}$ , and with the model of Isgur *et al.* (PR **D39**, 1357 (1989)) they obtain  $1.10 \pm 0.31^{+0.13}_{-0.14}$ .

## $\xi$ (all hadronic modes) PARAMETER

( $V-A$ ) theory predicts  $\xi = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.995±0.007 OUR FIT</b>				
<b>0.997±0.007 OUR AVERAGE</b>				
$0.992 \pm 0.007 \pm 0.008$	102k	<sup>1</sup> HEISTER 01E	ALEP	1991–1995 LEP runs
$0.997 \pm 0.027 \pm 0.011$	39k	<sup>2</sup> ABREU 00L	DLPH	1992–1995 runs
$1.02 \pm 0.13 \pm 0.03$	17.2k	<sup>3</sup> ASNER 00	CLEO	$E_{cm}^{ee} = 10.6$ GeV
$1.032 \pm 0.031$	37k	<sup>4</sup> ACCIARRI 98R	L3	1991–1995 LEP runs
$0.93 \pm 0.10 \pm 0.04$		ABE 970	SLD	1993–1995 SLC runs
$1.29 \pm 0.26 \pm 0.11$	7.4k	<sup>5</sup> ACKERSTAFF 97R	OPAL	1992–1994 LEP runs
$0.995 \pm 0.010 \pm 0.003$	66k	<sup>6</sup> ALEXANDER 97F	CLEO	$E_{cm}^{ee} = 10.6$ GeV
$1.03 \pm 0.06 \pm 0.04$	2.0k	<sup>7</sup> COAN 97	CLEO	$E_{cm}^{ee} = 10.6$ GeV
$1.017 \pm 0.039$		<sup>8</sup> ALBRECHT 95C	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6$ GeV
$1.25 \pm 0.23^{+0.15}_{-0.08}$	7.5k	<sup>9</sup> ALBRECHT 93C	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.970 \pm 0.053 \pm 0.011$	14k	<sup>10</sup> ACCIARRI 96H	L3	Repl. by ACCIARRI 98R
$1.08^{+0.46}_{-0.41}{}^{+0.14}_{-0.25}$	2.6k	<sup>11</sup> AKERS 95P	OPAL	Repl. by ACKERSTAFF 97R
$1.006 \pm 0.032 \pm 0.019$		<sup>12</sup> BUSKULIC 95D	ALEP	Repl. by HEISTER 01E
$1.022 \pm 0.028 \pm 0.030$	1.7k	<sup>13</sup> ALBRECHT 94E	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV
$0.99 \pm 0.07 \pm 0.04$		<sup>14</sup> BUSKULIC 94D	ALEP	1990+1991 LEP run
$1.14 \pm 0.34^{+0.34}_{-0.17}$	3.9k	<sup>9</sup> ALBRECHT 90I	ARG	Repl. by ALBRECHT 93C

<sup>1</sup> HEISTER 01E quote  $0.992 \pm 0.007 \pm 0.006 \pm 0.005$  where the errors are statistical, systematic, and an uncertainty due to the final state model. We combine the systematic error and model uncertainty. They use  $\tau \rightarrow \pi\nu_\tau$ ,  $\tau \rightarrow K\nu_\tau$ ,  $\tau \rightarrow \rho\nu_\tau$ , and  $\tau \rightarrow a_1\nu_\tau$  decays.

<sup>2</sup> ABREU 00L use  $\tau^- \rightarrow h^- \geq 0\pi^0\nu_\tau$  decays.

<sup>3</sup> ASNER 00 use  $\tau^- \rightarrow \pi^- 2\pi^0\nu_\tau$  decays.

<sup>4</sup> ACCIARRI 98R use  $\tau \rightarrow \pi\nu_\tau$ ,  $\tau \rightarrow K\nu_\tau$ , and  $\tau \rightarrow \rho\nu_\tau$  decays.

<sup>5</sup> ACKERSTAFF 97R use  $\tau \rightarrow a_1\nu_\tau$  decays.

<sup>6</sup> ALEXANDER 97F use  $\tau \rightarrow \rho\nu_\tau$  decays.

<sup>7</sup> COAN 97 use  $h^+ h^-$  energy correlations.

<sup>8</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E.

<sup>9</sup> Uses  $\tau \rightarrow a_1\nu_\tau$  decays. Replaced by ALBRECHT 95C.

<sup>10</sup> ACCIARRI 96H use  $\tau \rightarrow \pi\nu_\tau$ ,  $\tau \rightarrow K\nu_\tau$ , and  $\tau \rightarrow \rho\nu_\tau$  decays.

<sup>11</sup> AKERS 95P use  $\tau \rightarrow a_1 \nu_\tau$  decays.

<sup>12</sup> BUSKULIC 95D use  $\tau \rightarrow \pi \nu_\tau$ ,  $\tau \rightarrow \rho \nu_\tau$ , and  $\tau \rightarrow a_1 \nu_\tau$  decays.

<sup>13</sup> ALBRECHT 94E measure the square of this quantity and use the sign determined by ALBRECHT 90I to obtain the quoted result. Uses  $\tau \rightarrow a_1 \nu_\tau$  decays. Replaced by ALBRECHT 95C.

<sup>14</sup> BUSKULIC 94D use  $\tau \rightarrow \pi \nu_\tau$  and  $\tau \rightarrow \rho \nu_\tau$  decays. Superseded by BUSKULIC 95D.

## $\bar{\eta}(\mu)$ PARAMETER

( $V-A$ ) theory predicts  $\bar{\eta}(\mu) = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-1.3±1.5±0.8</b>	71k	<sup>1</sup> SHIMIZU	18A BELL	$\tau^- \rightarrow \nu_\tau \mu^- \bar{\nu}_\mu \gamma$

<sup>1</sup> SHIMIZU 18A measurement procedure fits a distribution affected by  $\bar{\eta}(\mu)$ ,  $\xi\kappa(\mu)$  and  $\eta''(\mu)$ , floating  $\bar{\eta}(\mu)$  and  $\xi\kappa(\mu)$  and fixing  $\eta''(\mu) = 0$ . The contribution of  $\eta''(\mu)$  is suppressed by  $m_\mu/m_\tau$ .

## $(\xi\kappa)(e \text{ or } \mu)$ PARAMETER

( $V-A$ ) theory predicts  $(\xi\kappa)(e \text{ or } \mu) = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.5±0.4±0.2</b>	149k	<sup>1,2</sup> SHIMIZU	18A BELL	$\tau^- \rightarrow \nu_\tau e^- \bar{\nu}_e \gamma$ and $\tau^- \rightarrow \nu_\tau \mu^- \bar{\nu}_\mu \gamma$

<sup>1</sup> SHIMIZU 18A measurement procedure fits a distribution of radiative tau decays into both electrons and muons affected by  $\bar{\eta}(e \text{ or } \mu)$ ,  $\xi\kappa(e \text{ or } \mu)$  and  $\eta''(e \text{ or } \mu)$ , floating  $\bar{\eta}(e \text{ or } \mu)$  and  $\xi\kappa(e \text{ or } \mu)$  and fixing  $\eta''(e \text{ or } \mu) = 0$ . The contribution of  $\eta''(e \text{ or } \mu)$  is suppressed by  $m_e/m_\tau$  for tau decaying to electrons and by  $m_\mu/m_\tau$  for tau decaying to muons.

<sup>2</sup> Error correlated with SHIMIZU 18A  $(\xi\kappa)(e)$  and  $(\xi\kappa)(\mu)$  values.

## $(\xi\kappa)(e)$ PARAMETER

( $V-A$ ) theory predicts  $(\xi\kappa)(e) = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.4±0.8±0.9</b>	78k	<sup>1,2</sup> SHIMIZU	18A BELL	$\tau^- \rightarrow \nu_\tau e^- \bar{\nu}_e \gamma$

<sup>1</sup> SHIMIZU 18A measurement procedure fits a distribution affected by  $\bar{\eta}(e)$ ,  $(\xi\kappa)(e)$  and  $\eta''(e)$ , floating  $(\xi\kappa)(e)$  and fixing  $\bar{\eta}(e) = 0$  and  $\eta''(e) = 0$ . The contribution of  $\eta''(e)$  is suppressed by  $m_e/m_\tau$ .

<sup>2</sup> Error correlated with SHIMIZU 18A  $(\xi\kappa)(e)$  value.

## $(\xi\kappa)(\mu)$ PARAMETER

( $V-A$ ) theory predicts  $(\xi\kappa)(\mu) = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.8±0.5±0.3</b>	71k	<sup>1,2</sup> SHIMIZU	18A BELL	$\tau^- \rightarrow \nu_\tau \mu^- \bar{\nu}_\mu \gamma$

<sup>1</sup> SHIMIZU 18A measurement procedure fits a distribution affected by  $\bar{\eta}(\mu)$ ,  $\xi\kappa(\mu)$  and  $\eta''(\mu)$ , floating  $\bar{\eta}(\mu)$  and  $\xi\kappa(\mu)$  and fixing  $\eta''(\mu) = 0$ . The contribution of  $\eta''(\mu)$  is suppressed by  $m_\mu/m_\tau$ .

<sup>2</sup> Error correlated with SHIMIZU 18A  $(\xi\kappa)(e \text{ or } \mu)$  value.

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EDWARDS	97	PR D55 3919	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
EDWARDS	97B	PR D56 5297	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
ESCRIBANO	97	PL B395 369	R. Escribano, E. Masso	(BARC, PARIT)
ABREU	96B	PL B365 448	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	96H	PL B377 313	M. Acciarri <i>et al.</i>	(L3 Collab.)

ACCIARRI	96K	PL B389 187	M. Acciari <i>et al.</i>	(L3 Collab.)
ALAM	96	PRL 76 2637	M.S. Alam <i>et al.</i>	(CLEO Collab.)
ALBRECHT	96E	PRPL 276 223	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALEXANDER	96D	PL B369 163	G. Alexander <i>et al.</i>	(OPAL Collab.)
ALEXANDER	96E	PL B374 341	G. Alexander <i>et al.</i>	(OPAL Collab.)
ALEXANDER	96S	PL B388 437	G. Alexander <i>et al.</i>	(OPAL Collab.)
BAI	96	PR D53 20	J.Z. Bai <i>et al.</i>	(BES Collab.)
BALEST	96	PL B388 402	R. Balest <i>et al.</i>	(CLEO Collab.)
BARTELT	96	PRL 76 4119	J.E. Bartelt <i>et al.</i>	(CLEO Collab.)
BUSKULIC	96	ZPHY C70 579	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	96C	ZPHY C70 561	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
COAN	96	PR D53 6037	T.E. Coan <i>et al.</i>	(CLEO Collab.)
ABE	95Y	PR D52 4828	K. Abe <i>et al.</i>	(SLD Collab.)
ABREU	95T	PL B357 715	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	95U	PL B359 411	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	95	PL B345 93	M. Acciari <i>et al.</i>	(L3 Collab.)
ACCIARRI	95F	PL B352 487	M. Acciari <i>et al.</i>	(L3 Collab.)
AKERS	95F	ZPHY C66 31	R. Akers <i>et al.</i>	(OPAL Collab.)
AKERS	95I	ZPHY C66 543	R. Akers <i>et al.</i>	(OPAL Collab.)
AKERS	95P	ZPHY C67 45	R. Akers <i>et al.</i>	(OPAL Collab.)
AKERS	95Y	ZPHY C68 555	R. Akers <i>et al.</i>	(OPAL Collab.)
ALBRECHT	95	PL B341 441	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	95C	PL B349 576	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	95G	ZPHY C68 25	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	95H	ZPHY C68 215	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BALEST	95C	PRL 75 3809	R. Balest <i>et al.</i>	(CLEO Collab.)
BERNABEU	95	NP B436 474	J. Bernabeu <i>et al.</i>	
BUSKULIC	95C	PL B346 371	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	95D	PL B346 379	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
Also		PL B363 265 (erratum)	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
ABREU	94K	PL B334 435	P. Abreu <i>et al.</i>	(DELPHI Collab.)
AKERS	94E	PL B328 207	R. Akers <i>et al.</i>	(OPAL Collab.)
AKERS	94G	PL B339 278	R. Akers <i>et al.</i>	(OPAL Collab.)
ALBRECHT	94E	PL B337 383	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ARTUSO	94	PRL 72 3762	M. Artuso <i>et al.</i>	(CLEO Collab.)
BARTELT	94	PRL 73 1890	J.E. Bartelt <i>et al.</i>	(CLEO Collab.)
BATTLE	94	PRL 73 1079	M. Battle <i>et al.</i>	(CLEO Collab.)
BAUER	94	PR D50 13	D.A. Bauer <i>et al.</i>	(TPC/2gamma Collab.)
BUSKULIC	94D	PL B321 168	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	94E	PL B332 209	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	94F	PL B332 219	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
GIBAUT	94B	PRL 73 934	D. Gibaut <i>et al.</i>	(CLEO Collab.)
ADRIANI	93M	PRPL 236 1	O. Adriani <i>et al.</i>	(L3 Collab.)
ALBRECHT	93C	ZPHY C58 61	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	93G	PL B316 608	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BALEST	93	PR D47 3671	R. Balest <i>et al.</i>	(CLEO Collab.)
BEAN	93	PRL 70 138	A. Bean <i>et al.</i>	(CLEO Collab.)
BORTOLETTO	93	PRL 71 1791	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
ESCRIBANO	93	PL B301 419	R. Escribano, E. Masso	(BARC)
PROCARIO	93	PRL 70 1207	M. Procario <i>et al.</i>	
ABREU	92N	ZPHY C55 555	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACTON	92F	PL B281 405	D.P. Acton <i>et al.</i>	(OPAL Collab.)
ACTON	92H	PL B288 373	P.D. Acton <i>et al.</i>	(OPAL Collab.)
AKERIB	92	PRL 69 3610	D.S. Akerib <i>et al.</i>	(CLEO Collab.)
Also		PRL 71 3395 (erratum)	D.S. Akerib <i>et al.</i>	(CLEO Collab.)
ALBRECHT	92D	ZPHY C53 367	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	92K	ZPHY C55 179	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	92M	PL B292 221	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	92Q	ZPHY C56 339	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AMMAR	92	PR D45 3976	R. Ammar <i>et al.</i>	(CLEO Collab.)
ARTUSO	92	PRL 69 3278	M. Artuso <i>et al.</i>	(CLEO Collab.)
BAI	92	PRL 69 3021	J.Z. Bai <i>et al.</i>	(BES Collab.)
BATTLE	92	PL B291 488	M. Battle <i>et al.</i>	(CLEO Collab.)
BUSKULIC	92J	PL B297 459	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
DECAMP	92C	ZPHY C54 211	D. Decamp <i>et al.</i>	(ALEPH Collab.)
ADEVA	91F	PL B265 451	B. Adeva <i>et al.</i>	(L3 Collab.)
ALBRECHT	91D	PL B260 259	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALEXANDER	91D	PL B266 201	G. Alexander <i>et al.</i>	(OPAL Collab.)
ANTREASYAN	91	PL B259 216	D. Antreasyan <i>et al.</i>	(Crystal Ball Collab.)
GRIFOLS	91	PL B255 611	J.A. Grifols, A. Mendez	(BARC)
ABACHI	90	PR D41 1414	S. Abachi <i>et al.</i>	(HRS Collab.)

ALBRECHT	90E	PL B246 278	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	90I	PL B250 164	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BEHREND	90	ZPHY C46 537	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
BOWCOCK	90	PR D41 805	T.J.V. Bowcock <i>et al.</i>	(CLEO Collab.)
DELAGUILA	90	PL B252 116	F. del Aguila, M. Sher	(BARC, WILL)
GOLDBERG	90	PL B251 223	M. Goldberg <i>et al.</i>	(CLEO Collab.)
WU	90	PR D41 2339	D.Y. Wu <i>et al.</i>	(Mark II Collab.)
ABACHI	89B	PR D40 902	S. Abachi <i>et al.</i>	(HRS Collab.)
BEHREND	89B	PL B222 163	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
JANSSEN	89	PL B228 273	H. Janssen <i>et al.</i>	(Crystal Ball Collab.)
KLEINWORT	89	ZPHY C42 7	C. Kleinwort <i>et al.</i>	(JADE Collab.)
ADEVA	88	PR D38 2665	B. Adeva <i>et al.</i>	(Mark-J Collab.)
ALBRECHT	88B	PL B202 149	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	88L	ZPHY C41 1	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	88M	ZPHY C41 405	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AMIDEI	88	PR D37 1750	D. Amidei <i>et al.</i>	(Mark II Collab.)
BEHREND	88	PL B200 226	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
BRAUNSCH...	88C	ZPHY C39 331	W. Braunschweig <i>et al.</i>	(TASSO Collab.)
KEH	88	PL B212 123	S. Keh <i>et al.</i>	(Crystal Ball Collab.)
TSCHIRHART	88	PL B205 407	R. Tschirhart <i>et al.</i>	(HRS Collab.)
ABACHI	87B	PL B197 291	S. Abachi <i>et al.</i>	(HRS Collab.)
ABACHI	87C	PRL 59 2519	S. Abachi <i>et al.</i>	(HRS Collab.)
ADLER	87B	PRL 59 1527	J. Adler <i>et al.</i>	(Mark III Collab.)
AIHARA	87B	PR D35 1553	H. Aihara <i>et al.</i>	(TPC Collab.)
AIHARA	87C	PRL 59 751	H. Aihara <i>et al.</i>	(TPC Collab.)
ALBRECHT	87L	PL B185 223	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	87P	PL B199 580	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BAND	87	PL B198 297	H.R. Band <i>et al.</i>	(MAC Collab.)
BAND	87B	PRL 59 415	H.R. Band <i>et al.</i>	(MAC Collab.)
BARINGER	87	PRL 59 1993	P. Baringer <i>et al.</i>	(CLEO Collab.)
BEBEK	87C	PR D36 690	C. Bebek <i>et al.</i>	(CLEO Collab.)
BURCHAT	87	PR D35 27	P.R. Burchat <i>et al.</i>	(Mark II Collab.)
BYLSMA	87	PR D35 2269	B.G. Bylsma <i>et al.</i>	(HRS Collab.)
COFFMAN	87	PR D36 2185	D.M. Coffman <i>et al.</i>	(Mark III Collab.)
DERRICK	87	PL B189 260	M. Derrick <i>et al.</i>	(HRS Collab.)
FORD	87	PR D35 408	W.T. Ford <i>et al.</i>	(MAC Collab.)
FORD	87B	PR D36 1971	W.T. Ford <i>et al.</i>	(MAC Collab.)
GAN	87	PRL 59 411	K.K. Gan <i>et al.</i>	(Mark II Collab.)
GAN	87B	PL B197 561	K.K. Gan <i>et al.</i>	(Mark II Collab.)
AIHARA	86E	PRL 57 1836	H. Aihara <i>et al.</i>	(TPC Collab.)
BARTEL	86D	PL B182 216	W. Bartel <i>et al.</i>	(JADE Collab.)
PDG	86	PL 170B 1	M. Aguilar-Benitez <i>et al.</i>	(CERN, CIT+)
RUCKSTUHL	86	PRL 56 2132	W. Ruckstuhl <i>et al.</i>	(DELCO Collab.)
SCHMIDKE	86	PRL 57 527	W.B. Schmidke <i>et al.</i>	(Mark II Collab.)
YELTON	86	PRL 56 812	J.M. Yelton <i>et al.</i>	(Mark II Collab.)
ALTHOFF	85	ZPHY C26 521	M. Althoff <i>et al.</i>	(TASSO Collab.)
ASH	85B	PRL 55 2118	W.W. Ash <i>et al.</i>	(MAC Collab.)
BALTRUSAIT...	85	PRL 55 1842	R.M. Baltrusaitis <i>et al.</i>	(Mark III Collab.)
BARTEL	85F	PL 161B 188	W. Bartel <i>et al.</i>	(JADE Collab.)
BEHRENDS	85	PR D32 2468	S. Behrends <i>et al.</i>	(CLEO Collab.)
BELTRAMI	85	PRL 54 1775	I. Beltrami <i>et al.</i>	(HRS Collab.)
BERGER	85	ZPHY C28 1	C. Berger <i>et al.</i>	(PLUTO Collab.)
BURCHAT	85	PRL 54 2489	P.R. Burchat <i>et al.</i>	(Mark II Collab.)
FERNANDEZ	85	PRL 54 1624	E. Fernandez <i>et al.</i>	(MAC Collab.)
MILLS	85	PRL 54 624	G.B. Mills <i>et al.</i>	(DELCO Collab.)
AIHARA	84C	PR D30 2436	H. Aihara <i>et al.</i>	(TPC Collab.)
BEHREND	84	ZPHY C23 103	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
MILLS	84	PRL 52 1944	G.B. Mills <i>et al.</i>	(DELCO Collab.)
BEHREND	83C	PL 127B 270	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
SILVERMAN	83	PR D27 1196	D.J. Silverman, G.L. Shaw	(UCI)
BEHREND	82	PL 114B 282	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
BLOCKER	82B	PRL 48 1586	C.A. Blocker <i>et al.</i>	(Mark II Collab.)
BLOCKER	82D	PL 109B 119	C.A. Blocker <i>et al.</i>	(Mark II Collab.) J
FELDMAN	82	PRL 48 66	G.J. Feldman <i>et al.</i>	(Mark II Collab.)
HAYES	82	PR D25 2869	K.G. Hayes <i>et al.</i>	(Mark II Collab.)
BERGER	81B	PL 99B 489	C. Berger <i>et al.</i>	(PLUTO Collab.)
DORFAN	81	PRL 46 215	J.M. Dorfan <i>et al.</i>	(Mark II Collab.)
BRANDELIK	80	PL 92B 199	R. Brandelik <i>et al.</i>	(TASSO Collab.)
ZHOLENTZ	80	PL 96B 214	A.A. Zholents <i>et al.</i>	(NOVO)
Also		SJNP 34 814	A.A. Zholents <i>et al.</i>	(NOVO)

Translated from YAF 34 1471.

BACINO	79B	PRL 42 749	W.J. Bacino <i>et al.</i>	(DELCO Collab.)
KIRKBY	79	SLAC-PUB-2419	J. Kirkby	(SLAC) J
Batavia Lepton Photon Conference.				
BACINO	78B	PRL 41 13	W.J. Bacino <i>et al.</i>	(DELCO Collab.) J
Also		Tokyo Conf. 249	J. Kirz	(STON)
Also		PL 96B 214	A.A. Zholents <i>et al.</i>	(NOVO)
BRANDELIK	78	PL 73B 109	R. Brandelik <i>et al.</i>	(DASP Collab.) J
FELDMAN	78	Tokyo Conf. 777	G.J. Feldman	(SLAC) J
JAROS	78	PRL 40 1120	J. Jaros <i>et al.</i>	(LGW Collab.)
PERL	75	PRL 35 1489	M.L. Perl <i>et al.</i>	(LBL, SLAC)

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## OTHER RELATED PAPERS

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GENTILE	96	PRPL 274 287	S. Gentile, M. Pohl	(ROMAI, ETH)
WEINSTEIN	93	ARNPS 43 457	A.J. Weinstein, R. Stroynowski	(CIT, SMU)
PERL	92	RPP 55 653	M.L. Perl	(SLAC)
PICH	90	MPL A5 1995	A. Pich	(VALE)
BARISH	88	PRPL 157 1	B.C. Barish, R. Stroynowski	(CIT)
GAN	88	IJMP A3 531	K.K. Gan, M.L. Perl	(SLAC)
HAYES	88	PR D38 3351	K.G. Hayes, M.L. Perl	(SLAC)
PERL	80	ARNPS 30 299	M.L. Perl	(SLAC)

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